

# Advanced Bistatic and Multistatic SAR Concepts and Applications

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# Future SAR Systems: Motivation

## Application Areas for SAR Data Products

### **Agriculture**

Precision Farming Suite  
Crop Ripeness  
Crop Inventory  
Yield Prediction Cereals



### **Forestry**

Strategic Forestry Inventory  
Reconnaissance Inventory  
Inventory Update



### **Cartography**

Topo Map  
Regional Planing Map  
Environmental Planing Map  
Infrastructure Planing Map



### **Marine**

Ship Detection Service  
Oil Spill Monitoring  
Sea Ice Monitoring



### **Risk / Disaster**

Flood Damage Assessment  
Fire Damage Assessment  
Storm Damage Assessment



### **Security**

Reconnaissance Imagery  
(VHR-SAR)



### **Geology**

Geology Structure Map  
Geology Image Map  
Geology Elevation Map  
Oil Seep Detection



### **Transportation**

Dynamic Traffic Monitoring  
Maps of Roads, Channels, ...

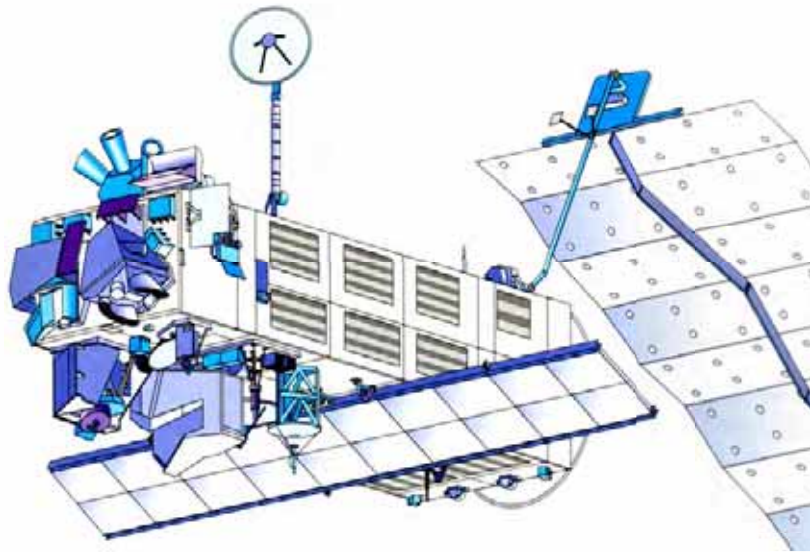


## Requirements

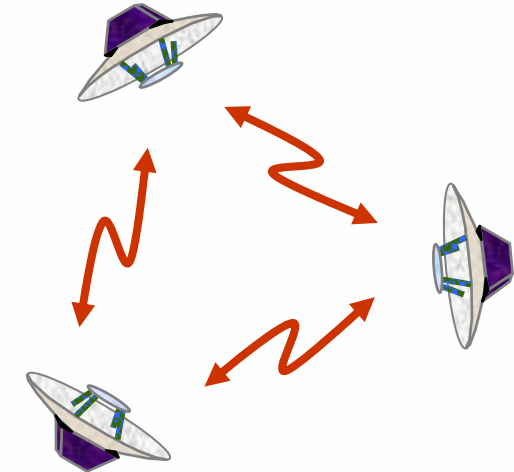
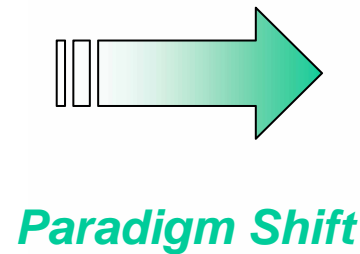
- **wide area coverage and short revisit times**
- **high radiometric and geometric resolution**
- **new data products:**
  - **high precision DEMs**
  - **3-D volume images (biomass, soil, ice, ...)**
  - **dynamic maps (ocean currents, traffic, ...)**
  - ...
- **high reliability and cost efficiency**



# Future SAR Systems: Paradigm Shift to Satellite Clusters



Large, multi-functional satellites



Virtual satellite -  
web of cooperating satellites

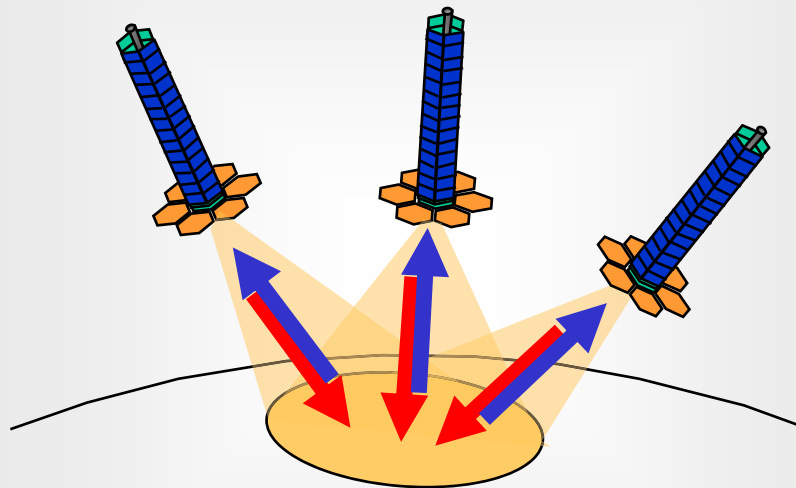
- Smaller, simpler satellites → reduced cost & time
- Modular design → upgradable, improved reliability
- Spatially distributed → improved revisit time / coverage / adaptability
- Separated, sparse apertures → improved performance and resolution

# Bistatic and Multistatic SAR Systems

- Definitions:**
- Radar systems with a spatial separation between transmitter and receiver are called **bistatic**.
  - Systems with multiple receivers are called **multistatic**.

## Fully active system

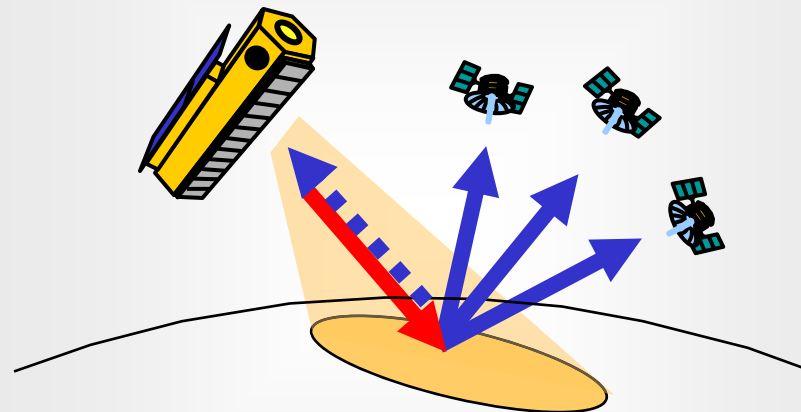
(TechSAT21, Radarsat 2/3, TanDEM-X)



- + more observables & baselines
- + phase synchronisation in ping/pong mode
- + high redundancy & great flexibility

## Semi active system

(BISSAT, Cartwheel, Pendulum)

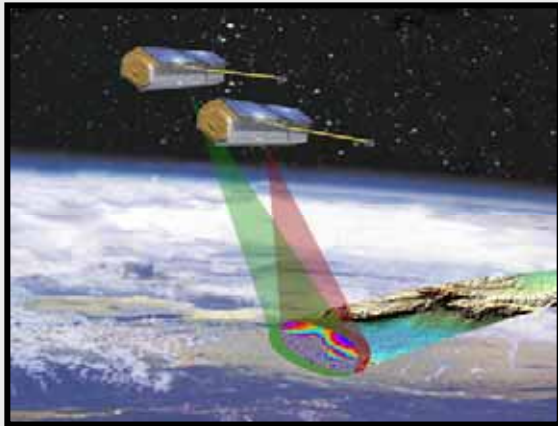


- + reduced size, weight and costs
- + increased sensitivity (no Tx/Rx switches)
- + receiver camouflage, robustness to jamming



# Examples for Suggested Missions

**TanDEM-X (Germany)**



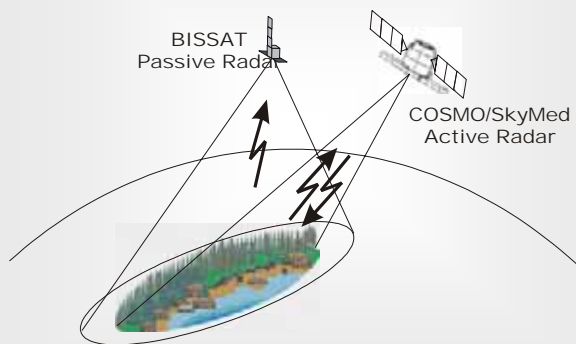
**Radarsat 2/3 (Canada)**



**TechSat 21 (USA)**



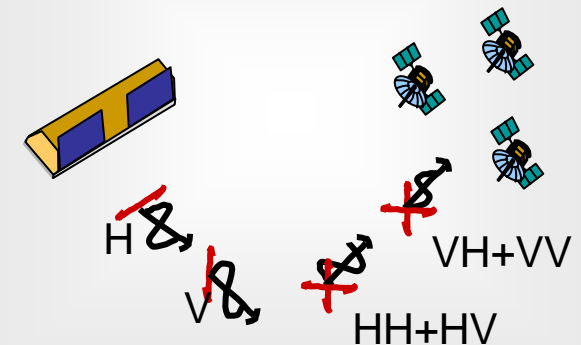
**BISSAT (Italy)**



**TerraSAR-L Cartwheel (ESA)**



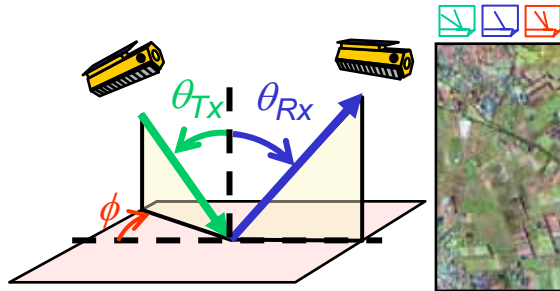
**Voice, Habitat (Earth Explorer Proposals)**



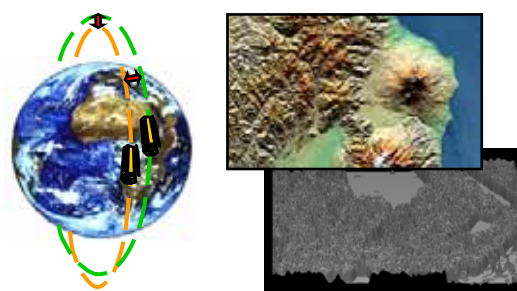
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# Potentials of Bistatic and Multistatic SAR Systems

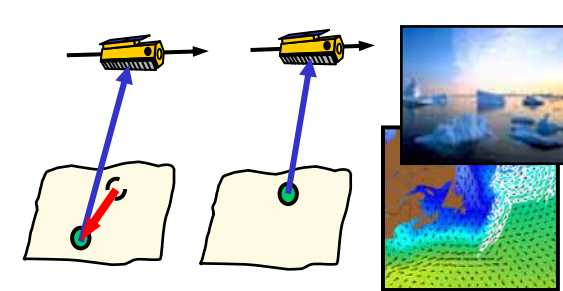
## Bistatic Imaging



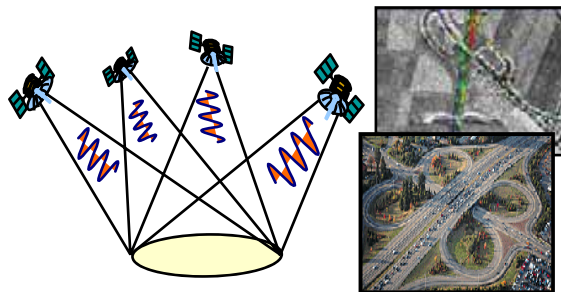
## Cross-Track Interferometry



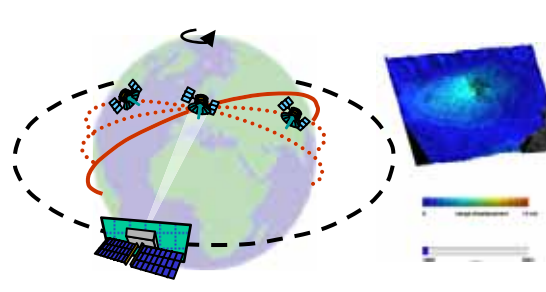
## Along-Track Interferometry



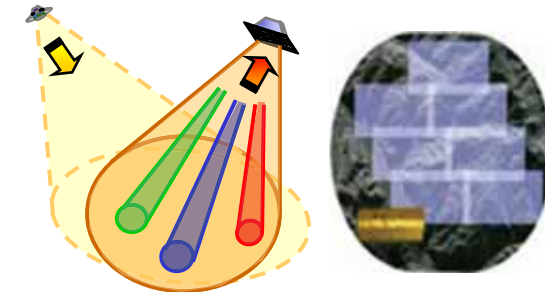
## Moving Target Indication



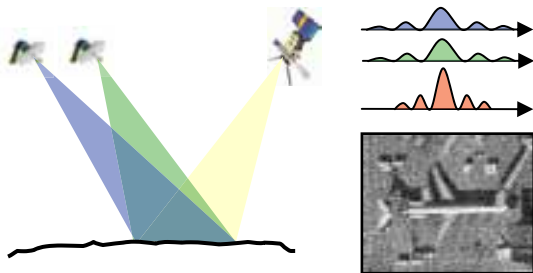
## Frequent Monitoring



## Wide Swath Imaging



## Resolution Enhancement



## SAR Tomography



...  
Interference Suppression  
Silent Operation  
Increased Radiom. Sensitivity  
Double Differential InSAR  
...



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# Geostationary Illuminator / LEO Receivers

## Basic Idea:

- constant illumination by geostationary transmitter
- signal reception by multiple low-cost receivers

## Receivers:

- passive (receive only)
- low power, small antennas
- low-cost micro-satellites
- low earth orbit

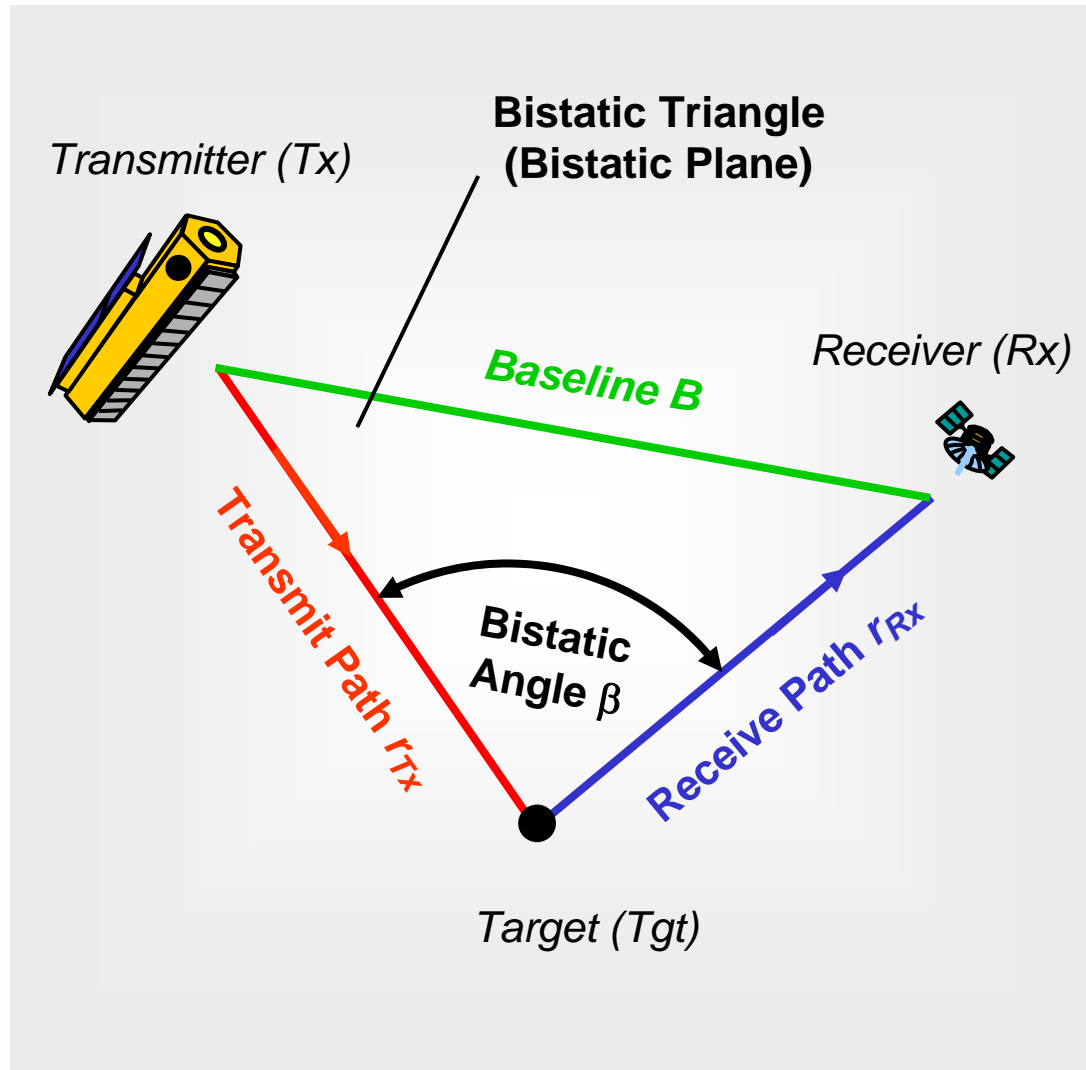
## Illuminator:

- geostationary orbit
- high Tx power (CW)
- large antenna area
- optional: steerable antenna

## Advantages:

- substantially improved revisit times without cost explosion
- multiple missions may share one illuminator

# Bistatic SAR Systems: Basic Definitions

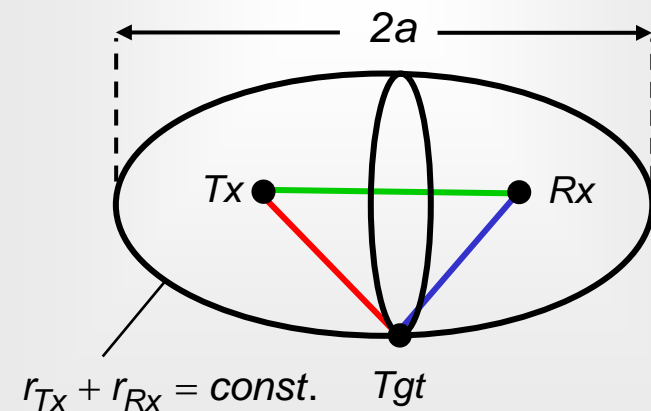


## Bistatic Range:

$$r = r_{Tx} + r_{Rx}$$

→ for a given range  $r$  the target is located on an **ellipsoid** with semi major axis

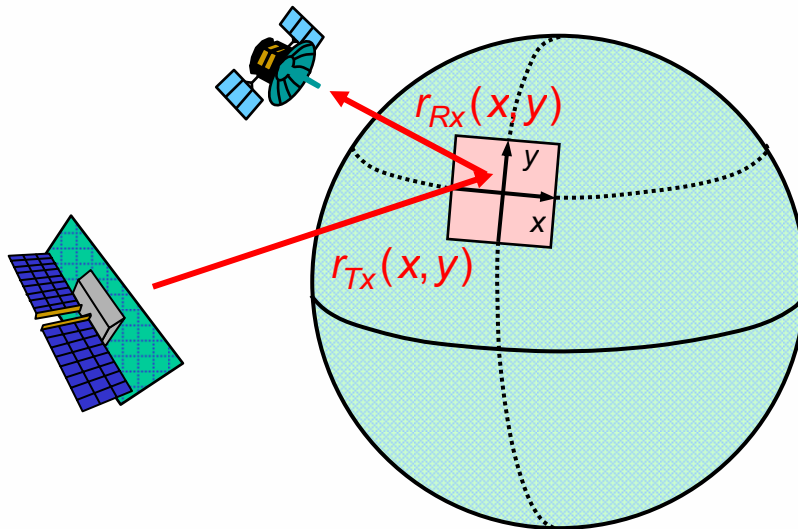
$$a = \frac{r_{Tx} + r_{Rx}}{2}$$





# Iso-Range Contours

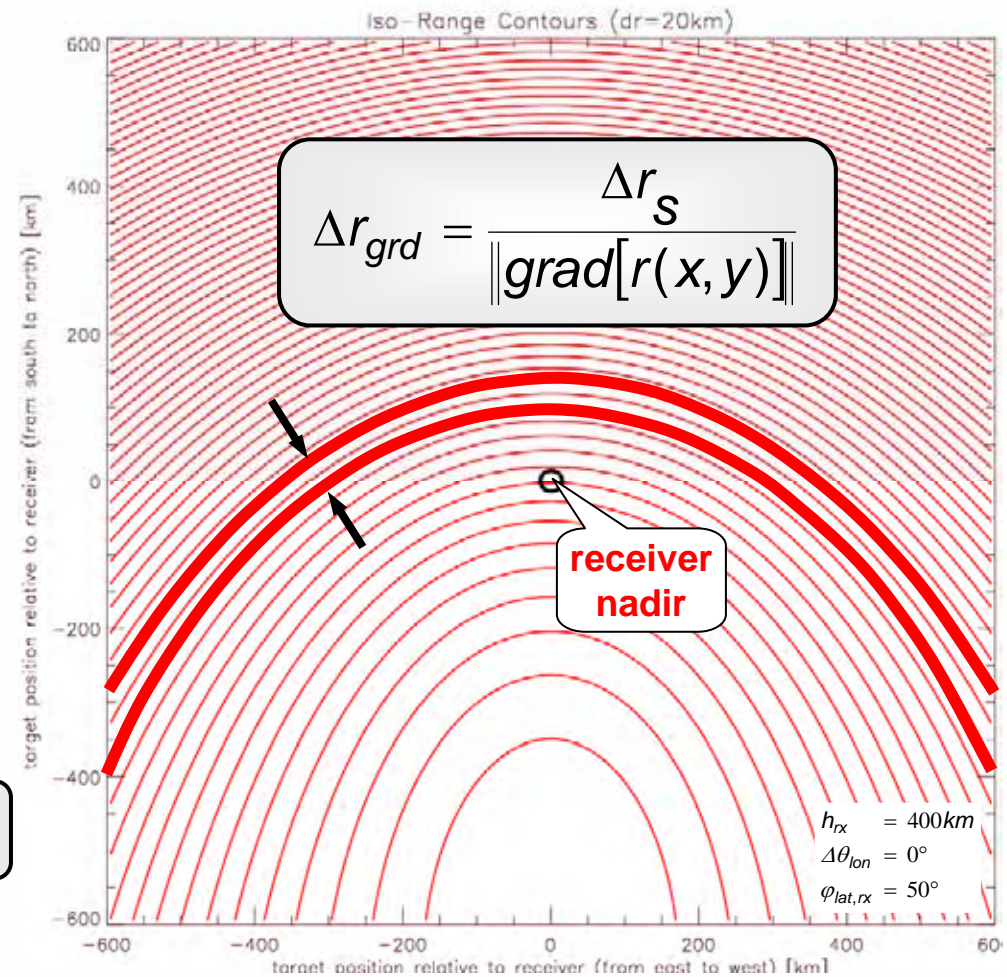
Tangent Plane:



Iso-Range Contours:

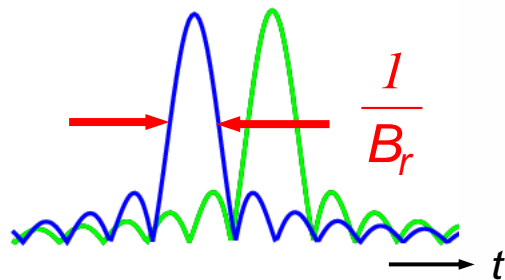
$$r(x, y) = r_{Tx}(x, y) + r_{Rx}(x, y) = \text{const.}$$

*(Intersection of Ellipsoid with Tangent Plane)*



# Ground Range Resolution

Slant Range Resolution:  
(after Pulse Compression)

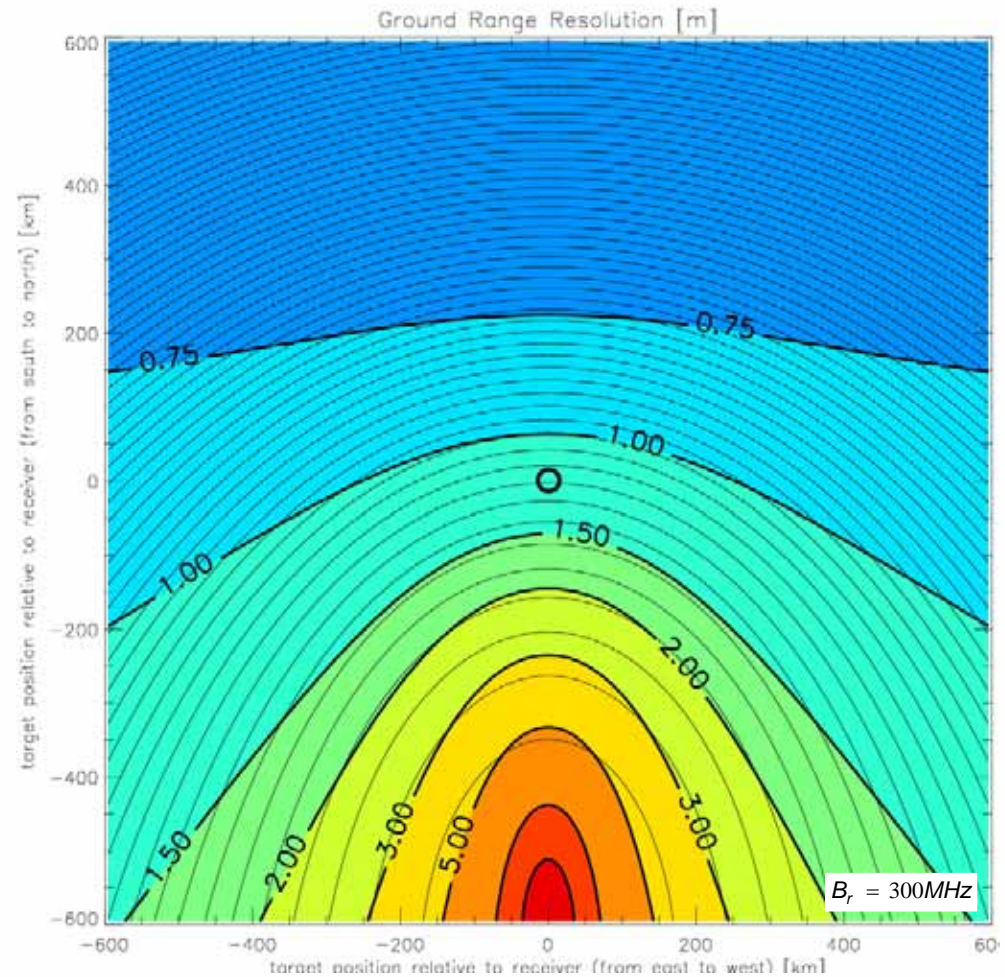


$$\Delta r_s = \frac{c_0}{B_r}$$

$B_r$  : bandwidth of range chirp  
 $c_0$  : velocity of light

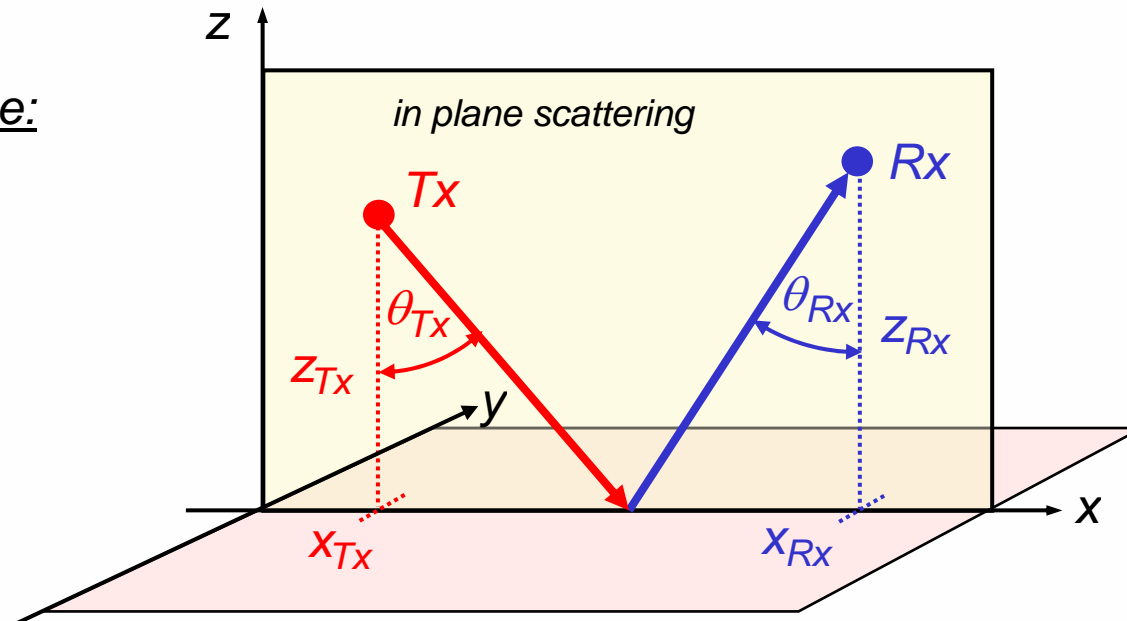
Ground Range Resolution:

$$\Delta r_{grd} = \frac{1}{\|\text{grad}[r(x,y)]\|} \cdot \frac{c_0}{B_r}$$



# Ground Range Resolution

Special Case:



$$r(x, y) = \underbrace{\sqrt{(x - x_{Tx})^2 + z_{Tx}^2}}_{r_{Tx}(x, y)} + \underbrace{\sqrt{(x - x_{Rx})^2 + z_{Rx}^2}}_{r_{Rx}(x, y)} \longrightarrow \frac{\partial}{\partial x} [r(x, y)] = \underbrace{\frac{x - x_{Tx}}{\sqrt{(x - x_{Tx})^2 + z_{Tx}^2}}}_{\sin \theta_{Tx}} + \underbrace{\frac{x - x_{Rx}}{\sqrt{(x - x_{Rx})^2 + z_{Rx}^2}}}_{-\sin \theta_{Rx}}$$

$$\Delta r_{\text{grd}} = \frac{c_0}{B_r |\sin \theta_{Tx} - \sin \theta_{Rx}|}$$



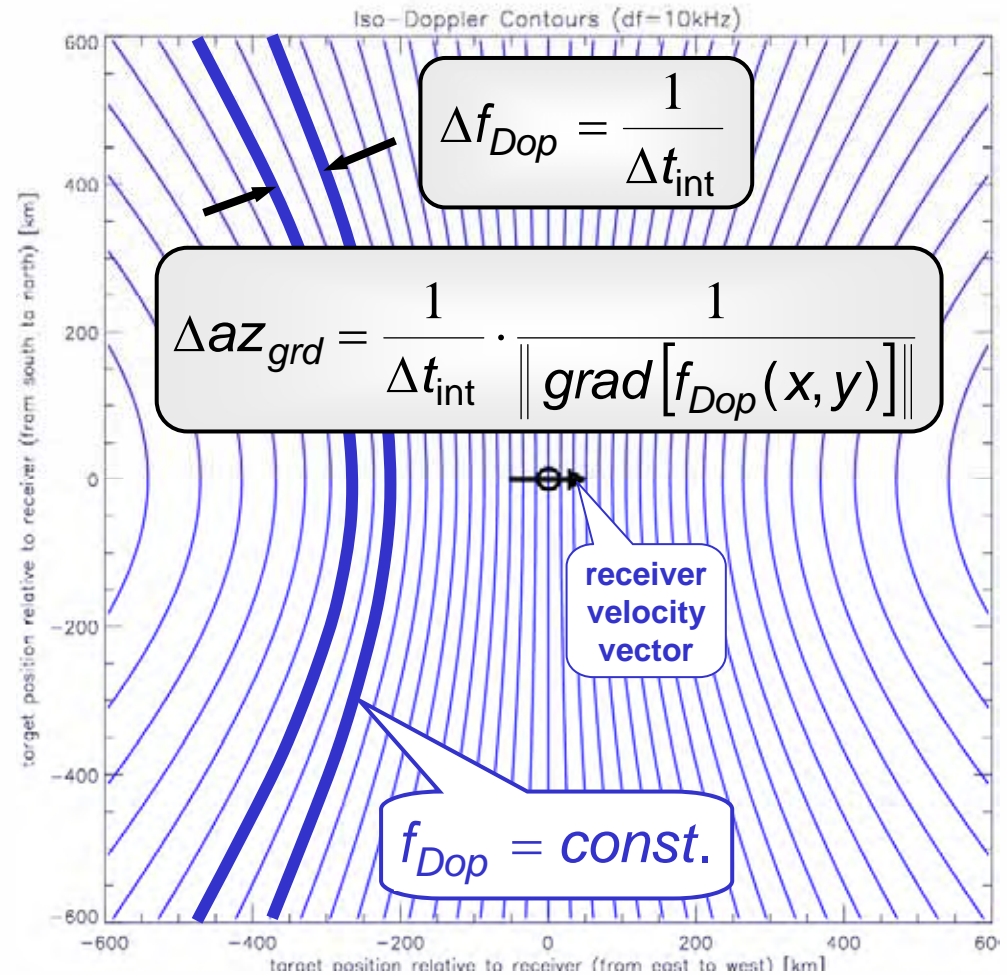
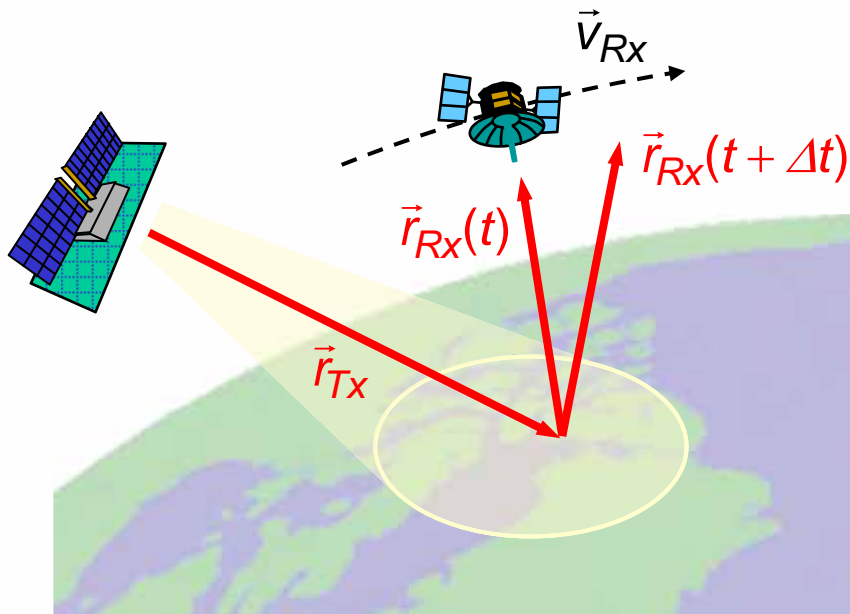


# Bistatic Iso-Doppler Contours

## Doppler Frequencies:

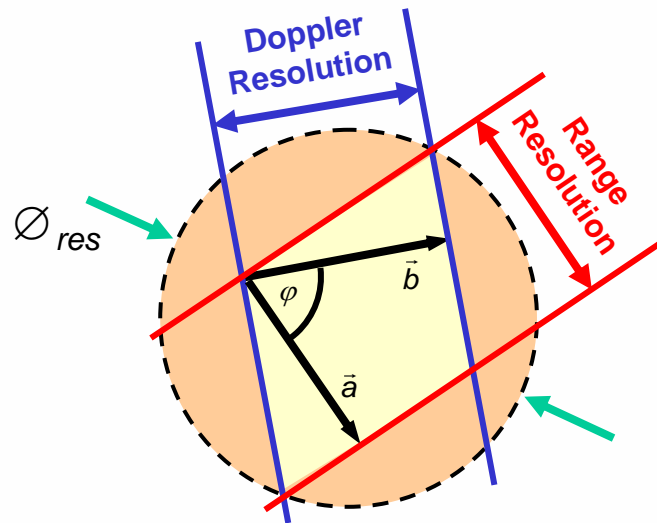
$$f_{Dop} = \frac{-1}{\lambda} \left[ \frac{\partial}{\partial t} (r_{Tx} + r_{Rx}) \right] \approx$$

$$\approx \frac{-1}{\lambda} \left[ \frac{\vec{V}_{Tx} \cdot \vec{r}_{Tx}}{\|\vec{r}_{Tx}\|} + \frac{\vec{V}_{Rx} \cdot \vec{r}_{Rx}}{\|\vec{r}_{Rx}\|} \right]$$





# Resolution Cell: Area and Diameter

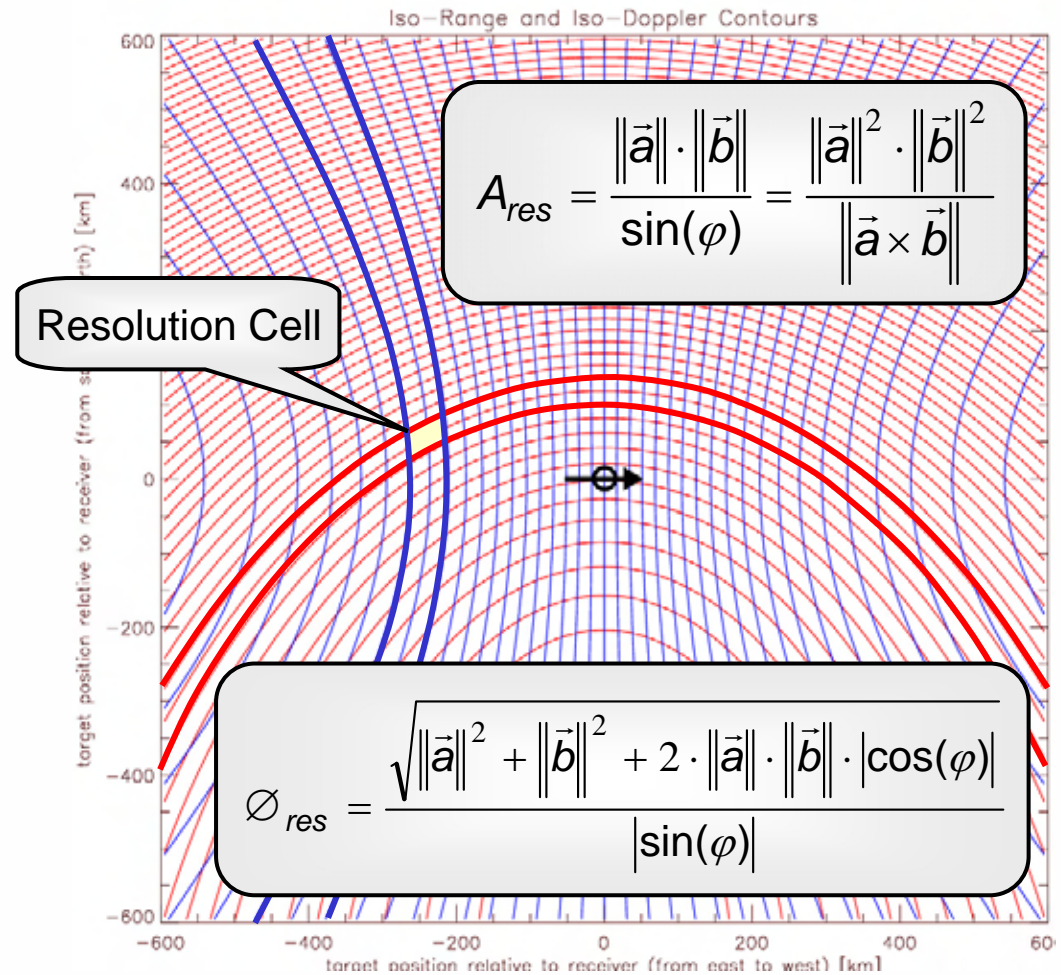


Range Resolution Vector:

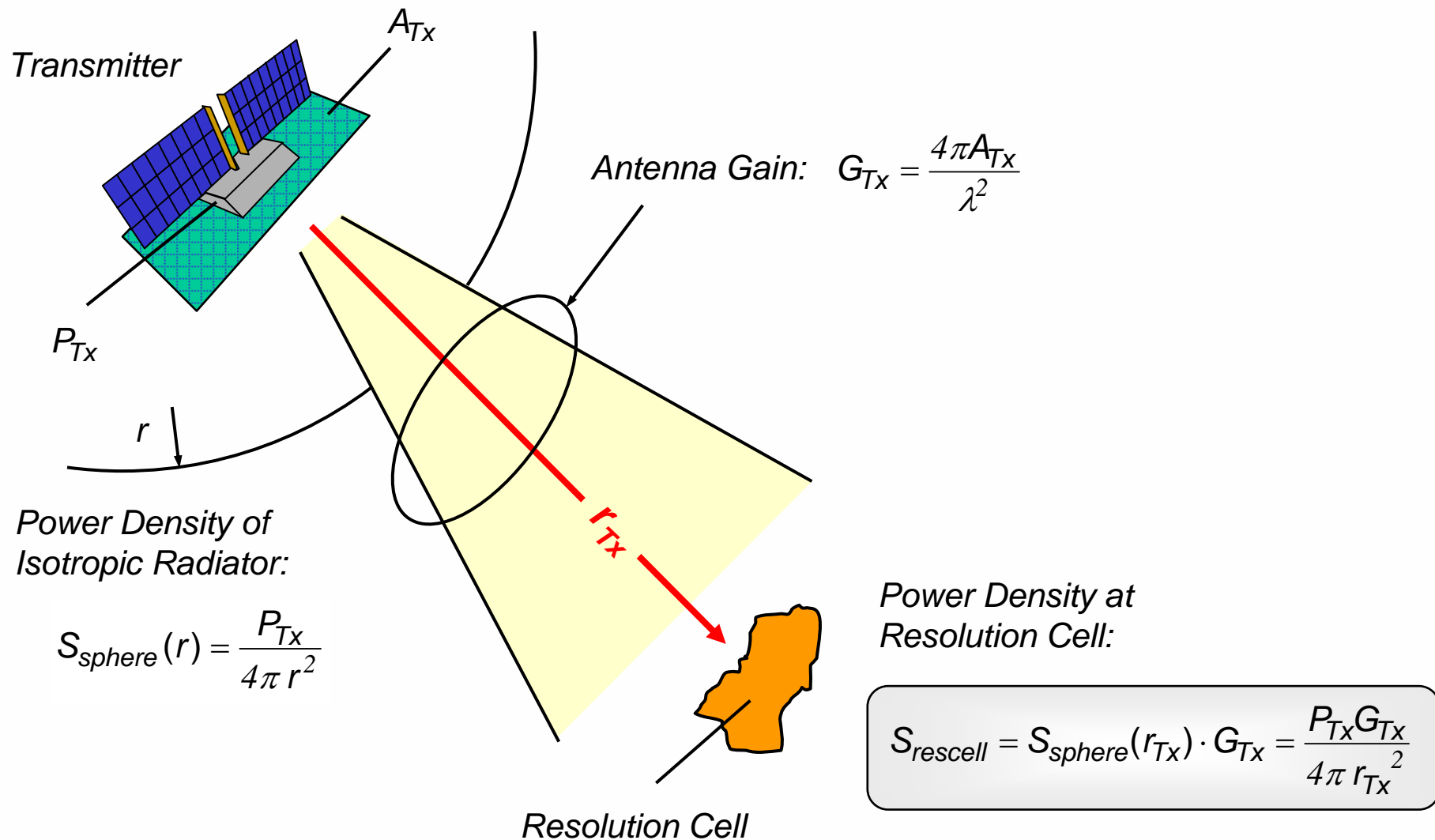
$$\vec{a} = \frac{\text{grad}(r(x,y))}{\|\text{grad}(r(x,y))\|^2} \cdot \frac{c_0}{B_r}$$

Doppler Resolution Vector:

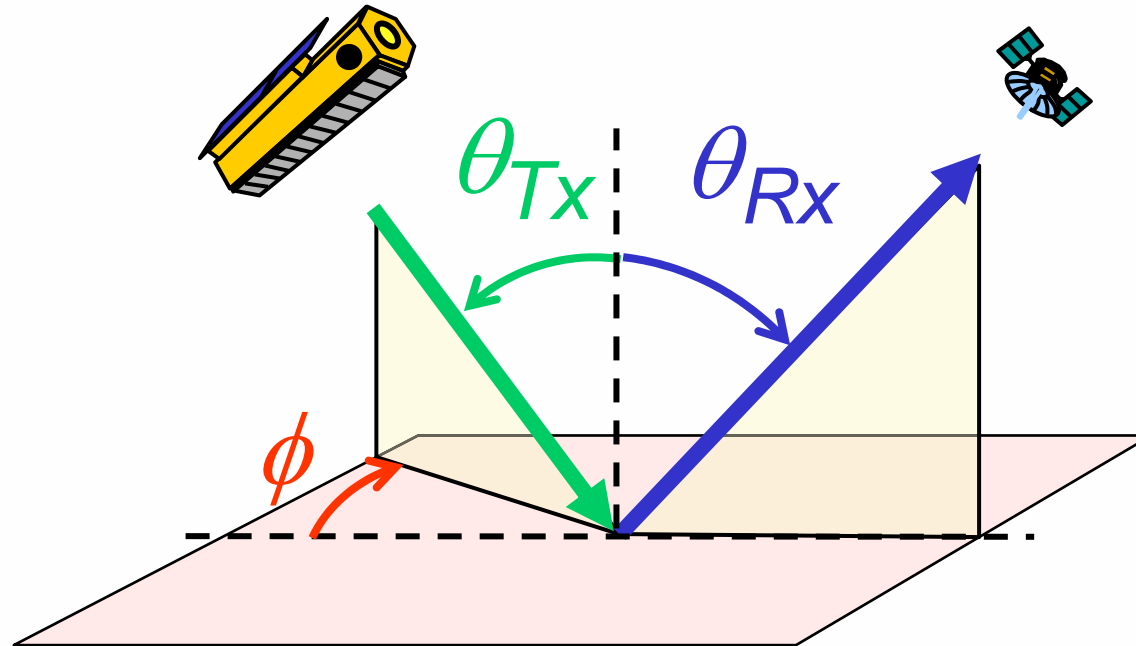
$$\vec{b} = \frac{\text{grad}(f_{Dop}(x,y))}{\|\text{grad}(f_{Dop}(x,y))\|^2} \cdot \frac{1}{t_{\text{int}}(x,y)}$$



# Bistatic Radar Equation (1)



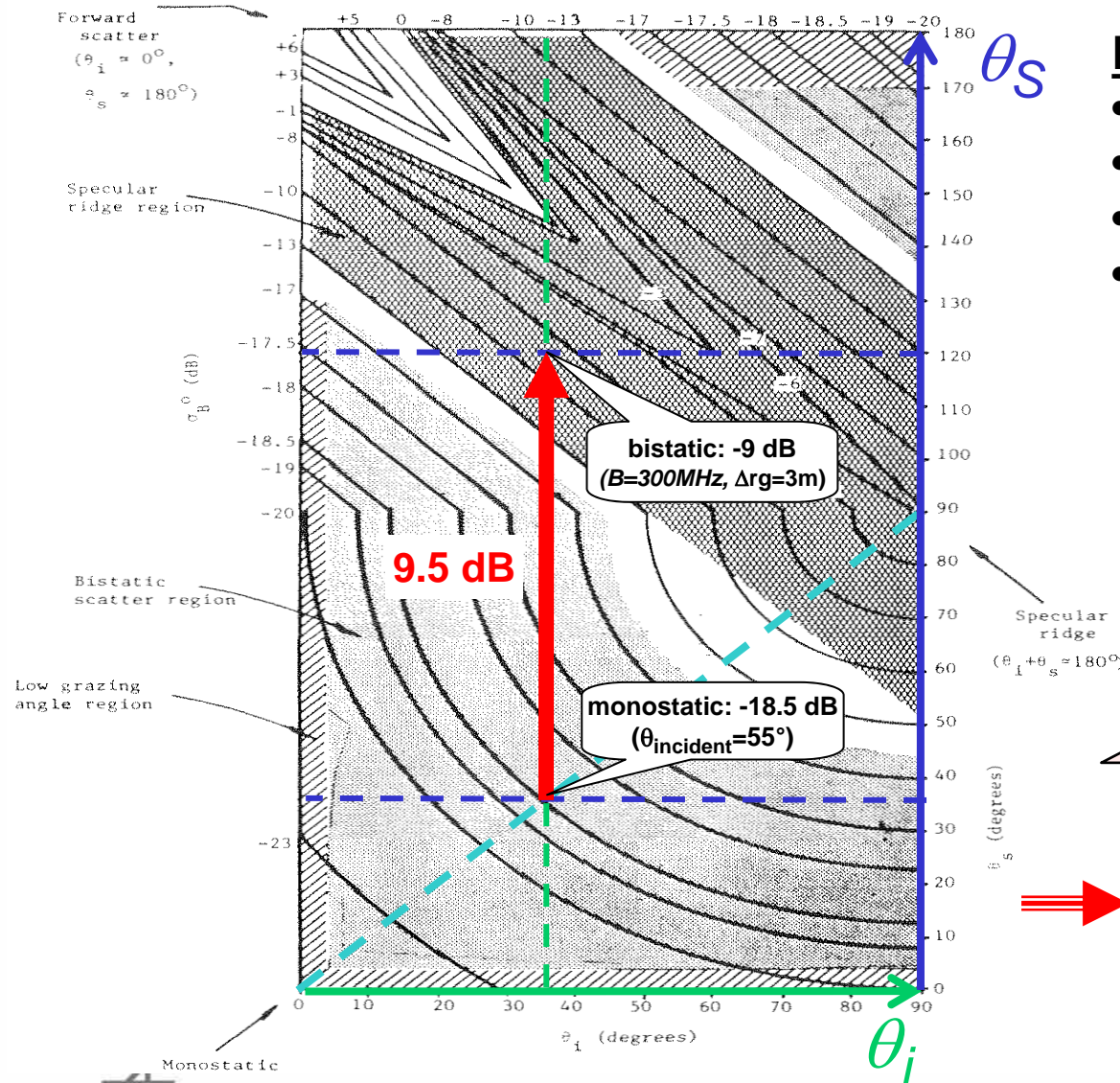
# Bistatic Scattering Coefficient



Bistatic scattering coefficient depends on:

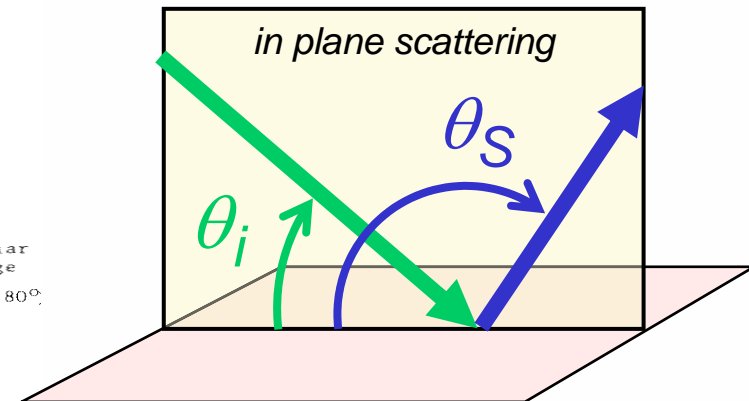
- Incident angle  $\theta_{Tx}$
- Scattering angle  $\theta_{Rx}$
- Out-of-plane angle  $\phi$
- Frequency
- Polarization
- Surface / Object

# Bistatic Scattering Coefficient: Example



## Data from Domville (1967)

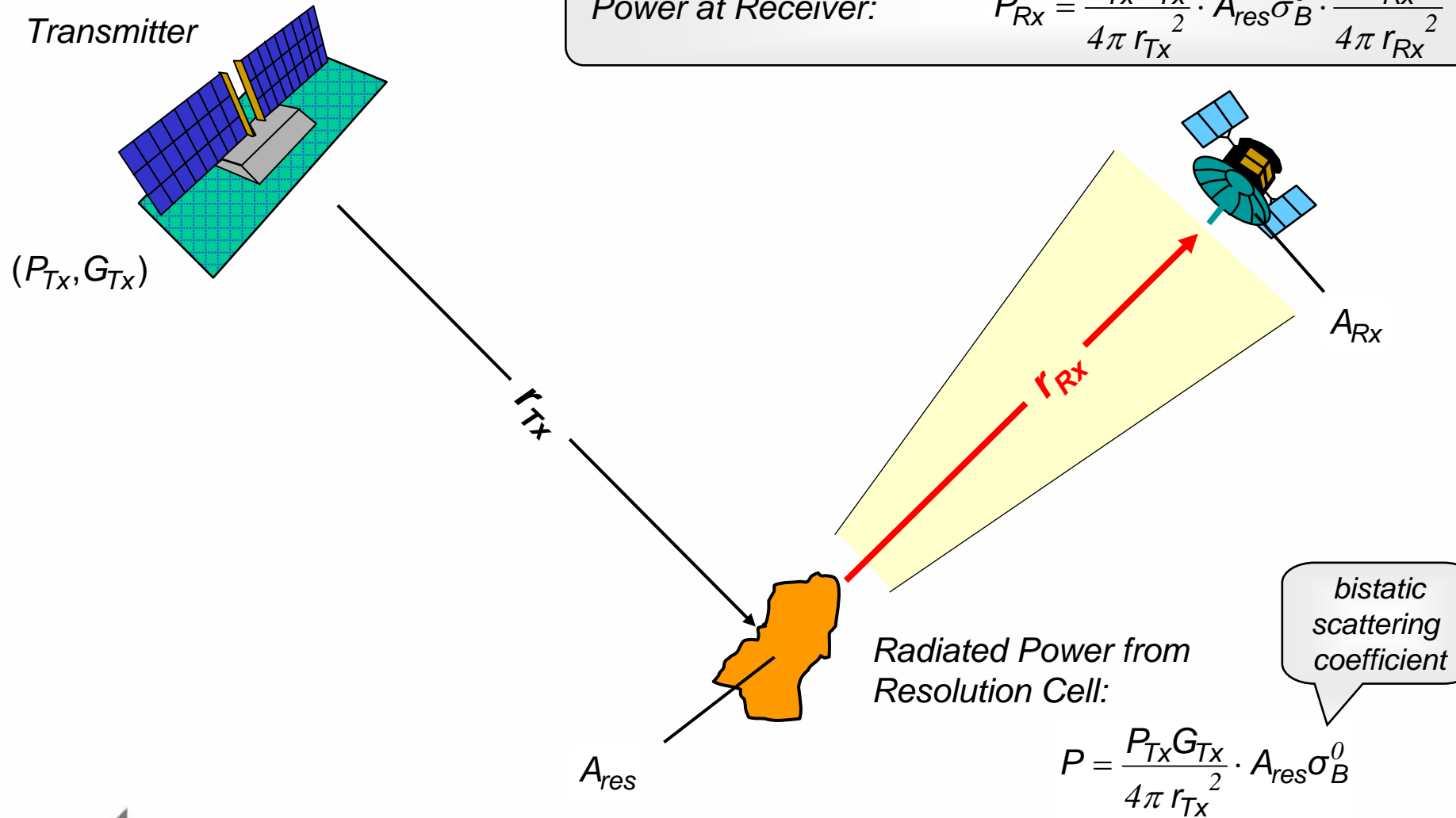
- X-Band scattering for rural land
- vertical polarization (VV)
- only in-plane scattering
- see also Willis, 1991



⇒ **Forward scattering may increase the radar cross section by 10 dB and more**



## Bistatic Radar Equation (2)



# Signal to Noise Ratio

## Single Pulse SNR:

$$SNR_I = \frac{\frac{P_{Tx} G_{Tx}}{4\pi r_{Tx}^2} \cdot A_{res} \sigma_B^0 \cdot \frac{A_{Rx}}{4\pi r_{Rx}^2}}{k T_s B_n FL}$$

power at receiver

thermal noise  
& losses

## Independent Samples:

$$n_{rg} = B_r \cdot \tau_P \approx B_n \cdot \tau_P = B_n \cdot \frac{P_{avg}}{P_{Tx} \cdot PRF}$$

$$n_{az} = t_{int} \cdot PRF$$

## SNR after Coherent Integration:

$$SNR_n = n_{rg} \cdot n_{az} \cdot SNR_I = \frac{P_{avg} G_{Tx}}{4\pi r_{Tx}^2} \cdot A_{res} \sigma_B^0 \cdot \frac{A_{Rx}}{4\pi r_{Rx}^2} \cdot \frac{t_{int}}{k T_s FL}$$

usually  $A_{res} \sim \frac{1}{t_{int}}$   
→ SNR independent  
of azimuth resolution

$r_{Tx}$	Distance from transmitter to scene
$r_{Rx}$	Distance from receiver to scene
$k$	Boltzmann's constant ( $1.3805 \cdot 10^{-23}$ Ws/K)
$T_s$	Receiving system noise temperature
$B_r$ ( $B_n$ )	Range Chirp (Receiver) Bandwidth
$F$	System Noise Figure
$L$	Losses
$\tau_P$	Pulse length
$PRF$	Pulse repetition frequency
$P_{avg}$ ( $P_{Tx}$ )	Average (Peak) transmit power
$G_{Tx}$	Gain of transmit antenna
$A_{Rx}$	Effective aperture of receive antenna
$A_{res}$	Size of resolution cell (1 look)
$t_{int}$	Coherent integration time





## Noise Equivalent Sigma Zero (NESZ)

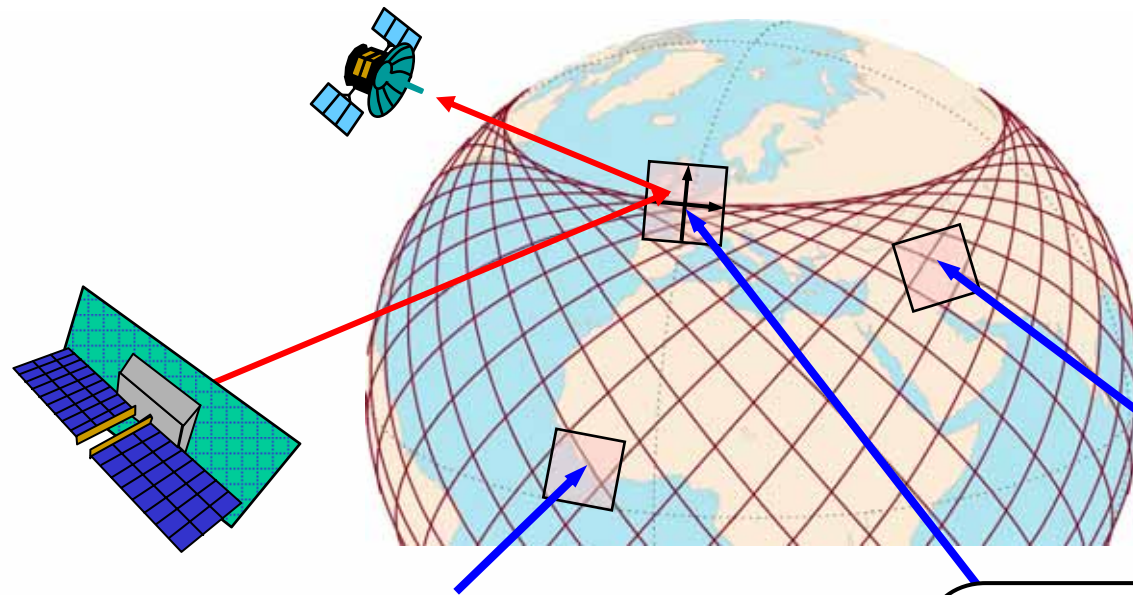
- The Noise Equivalent Sigma Zero (NESZ) measures the **sensitivity** of a given SAR
- The NESZ corresponds to the bistatic scattering coefficient for which the SNR is equal to one

$$SNR = \frac{P_{avg} G_{Tx} A_{Rx} \cdot A_{res} \cdot t_{int}}{(4\pi)^2 r_{Tx}^2 r_{Rx}^2 k T_s FL} \cdot \sigma_B^0 \stackrel{!}{=} 1 \quad (0 \text{ dB})$$

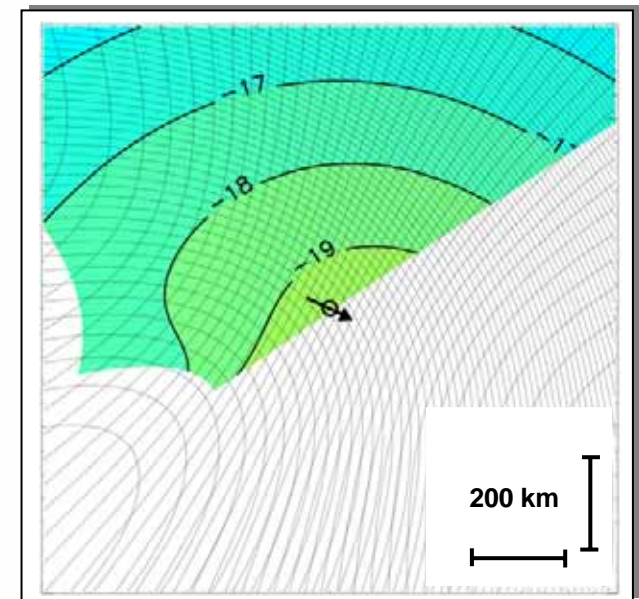
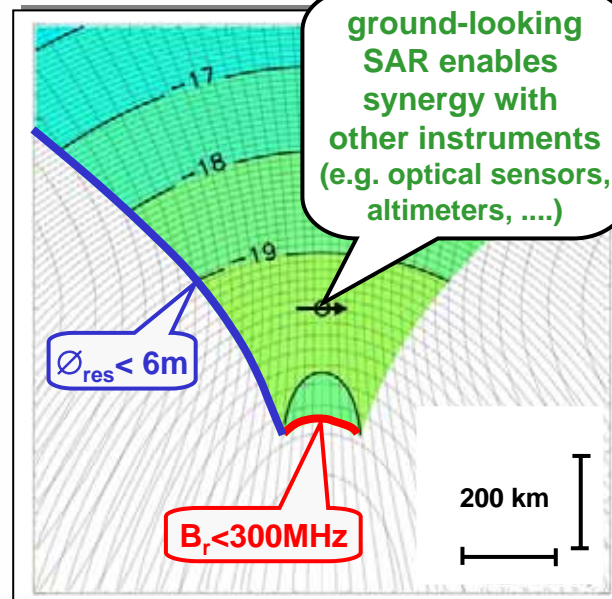
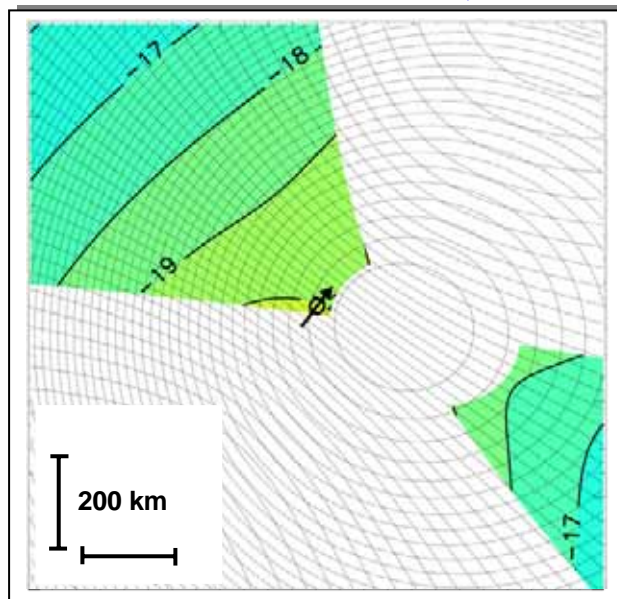
$$\rightarrow \boxed{NESZ = \sigma_B^0(SNR = 1) = \frac{(4\pi)^2 r_{Tx}^2 r_{Rx}^2 k T_s FL}{P_{avg} G_{Tx} A_{Rx} \cdot A_{res} \cdot t_{int}}}$$

- Lower NESZ values are better
- For spaceborne SAR, typical values of the NESZ are in the order of -20 dB
- The SNR is given by:  $SNR = \sigma_B^0 - NESZ \quad [dB]$

# NESZ Example: Geostationary Illuminator



Wavelength	3.1 cm
<b>Max. Bandwidth</b>	<b>300 MHz</b>
Average Transmit Power	1000 W
Antenna Size Tx	100 m <sup>2</sup>
Antenna Size Rx	6 m <sup>2</sup>
Noise Figure + Losses	5 dB
Receiver Altitude	400 km
<b>Ground Resolution</b>	<b>3 m</b>
<b>Max. Res. Cell Diameter</b>	<b>6 m</b>

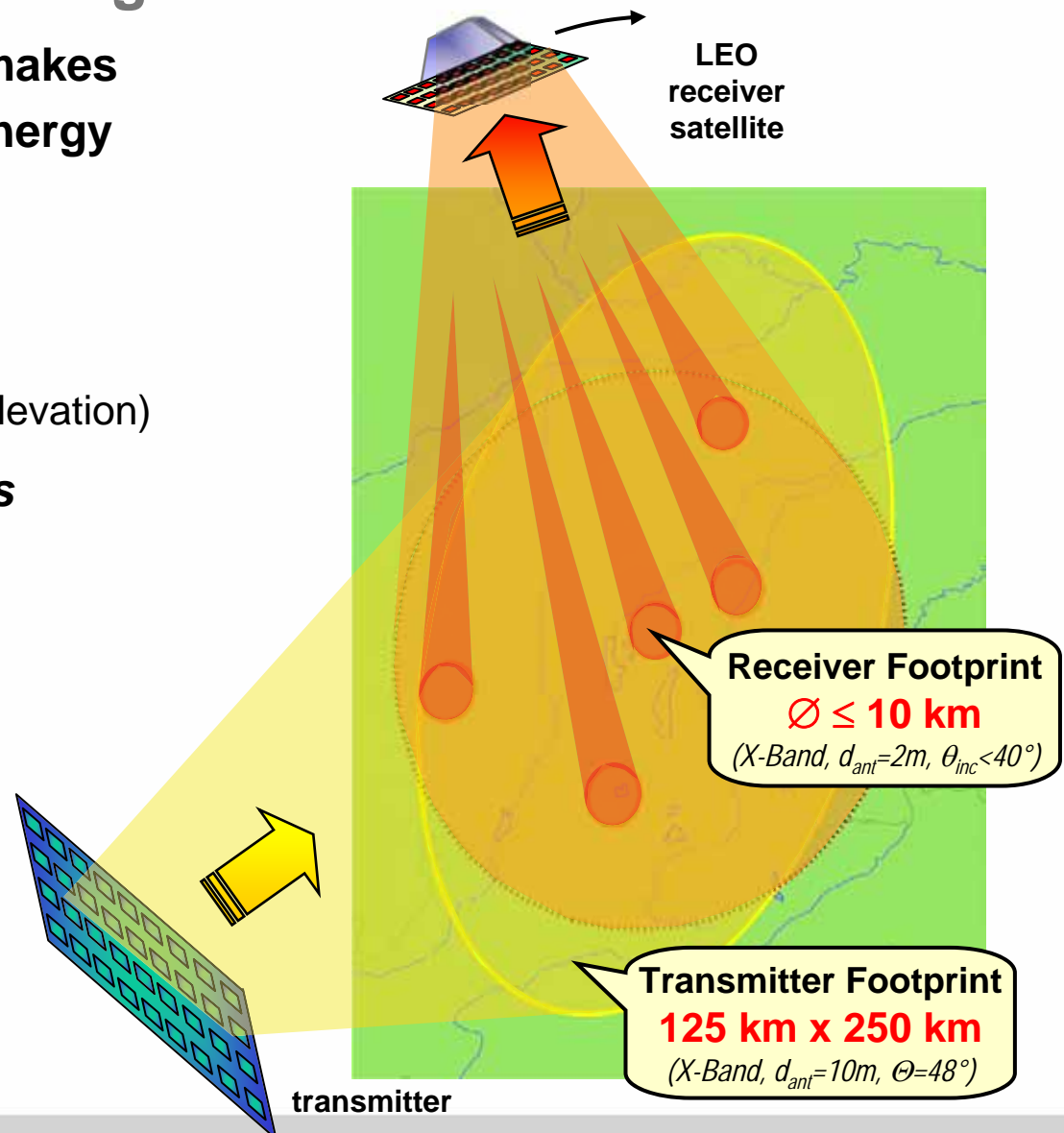




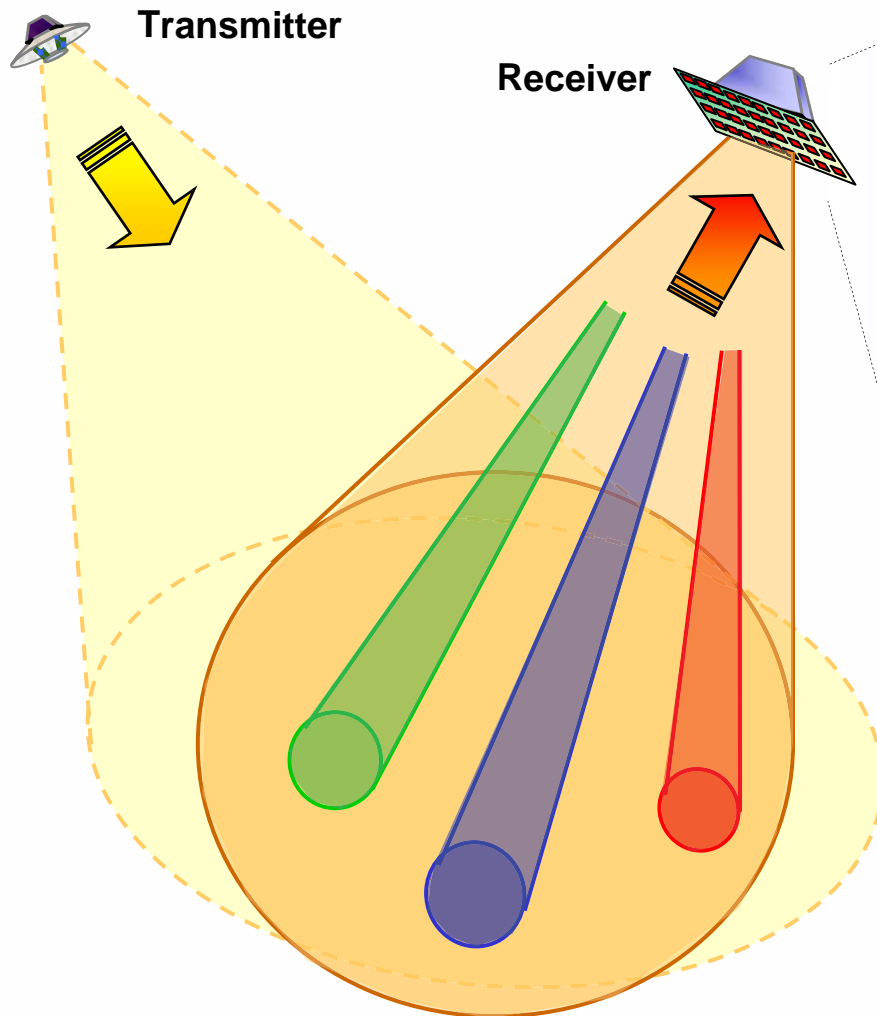
# Digital Beamforming in Passive Receivers

Digital beamforming on receive makes effective use of the total signal energy in the large illuminated footprint:

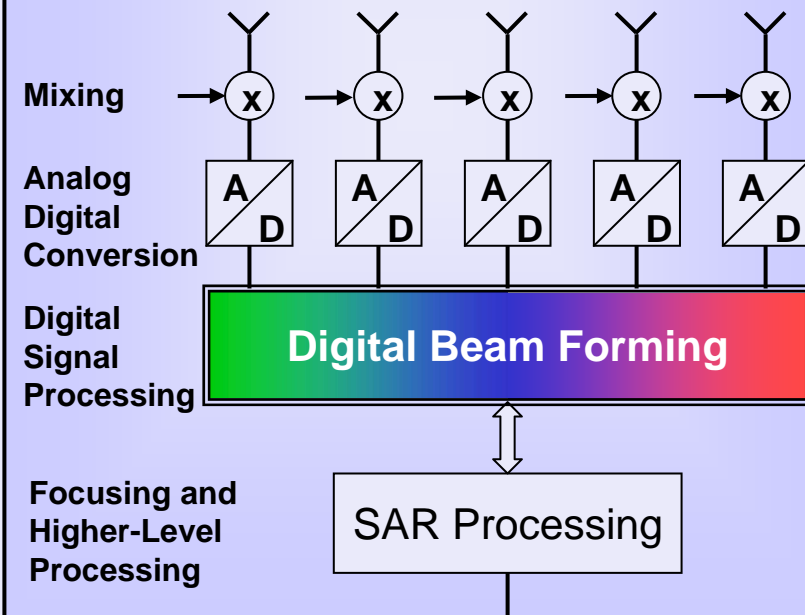
- ⇒ mapping of a *wide swath* or *multiple spots*  
(in spite of extended antennas in elevation)
- ⇒ *very long synthetic apertures*  
(also with long receiver apertures)
  - *high azimuth resolution*
  - *more independent looks*
  - *improved sensitivity*
- ⇒ interference suppression
- ⇒ ambiguity reduction
- ⇒ multiple phase center MTI



# Digital Beamforming on Receive



Multiple beams with adaptable antenna patterns





# Parasitic SAR with Communication Satellite

## Basic Idea:

- illumination by a transmitter of opportunity
- sufficient SNR is provided by very long coherent integration time and moderate resolution

## Receivers:

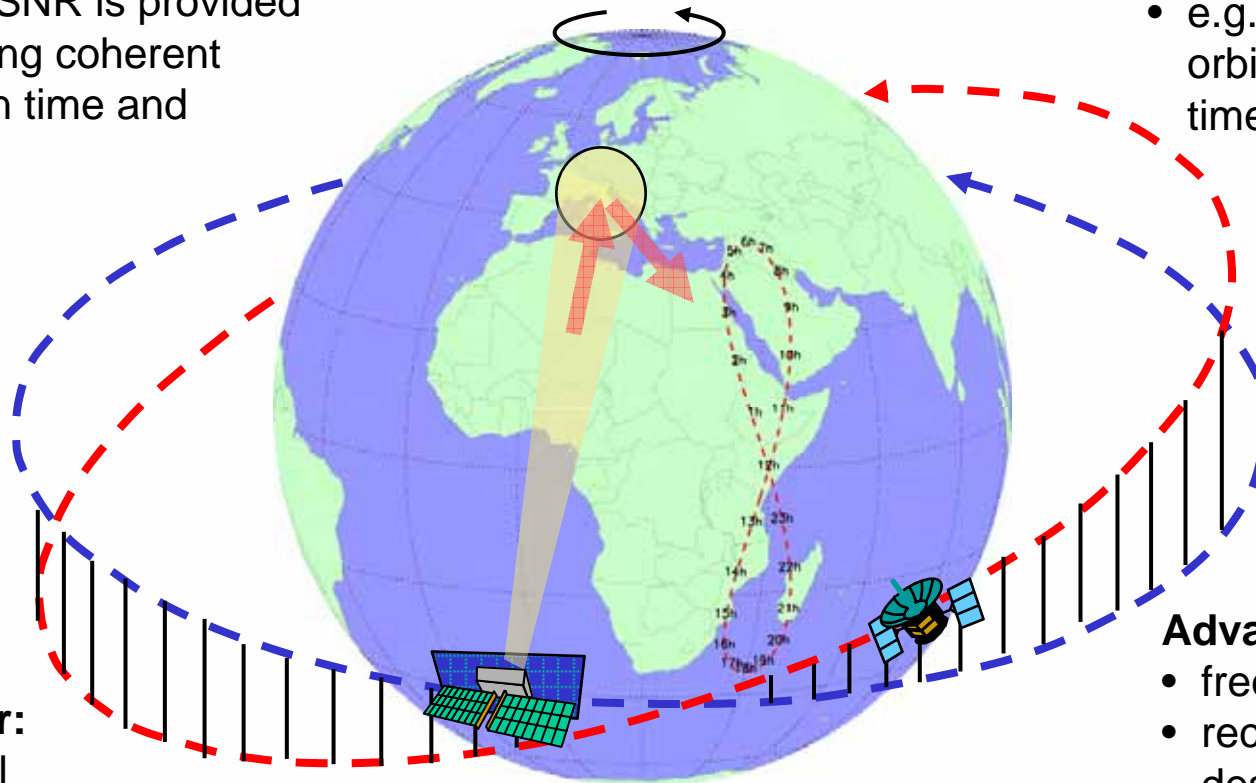
- passive, low-cost mini- or micro-satellites
- e.g. geosynchronous orbit for long integration time (Prati et al., 1998)

## Illuminator:

- e.g. digital communication satellite
- geostationary orbit

## Advantages:

- free transmitter
- receiving part can be designed using commercial DAB- or TV-SAT components



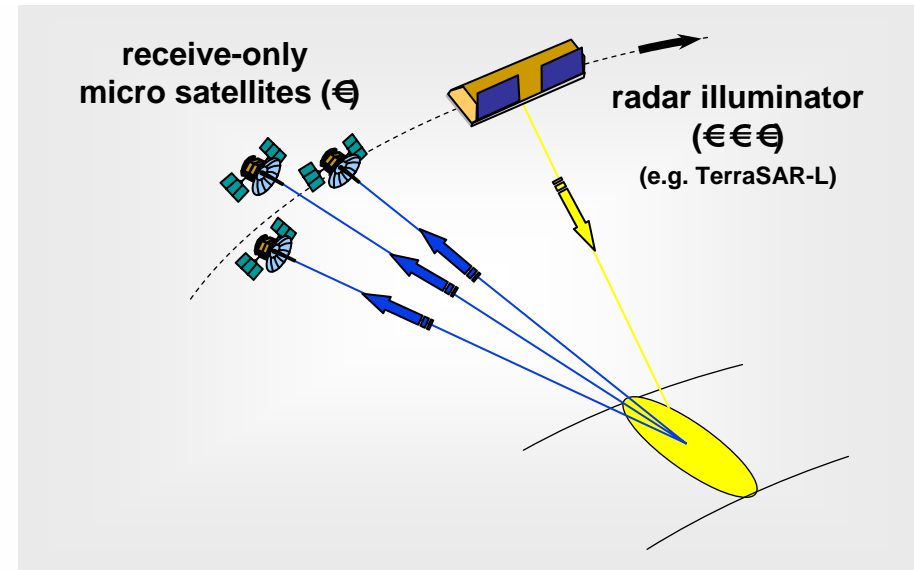
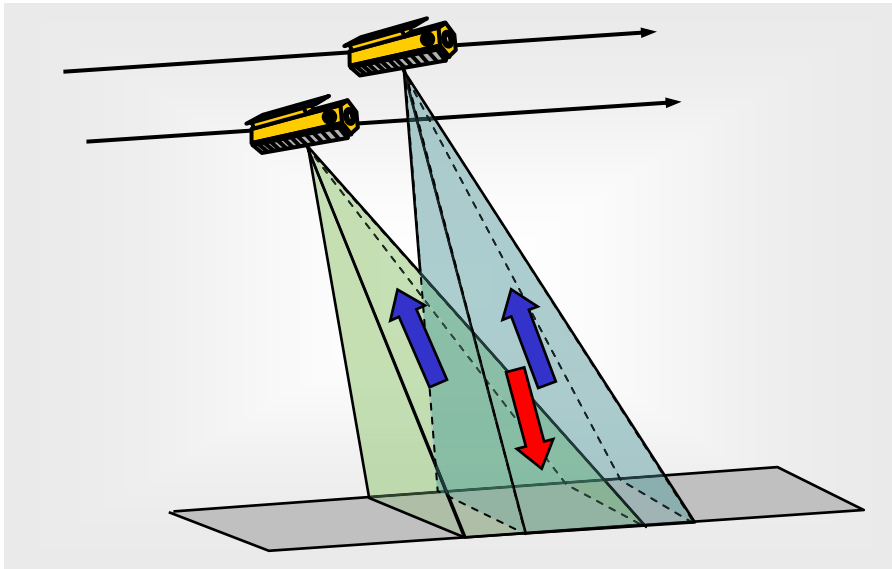


# Performance of a Parasitic SAR with Communication Satellite and Geosynchronous Receiver

Power Budget Example (cf. Prati, Rocca, Giancola, Monti Guarnieri, 1998):

Geostationary Illuminator & Geosynchronous Receiver			
<b>Power Density on Ground</b> $P_{ground} = \frac{P_{Tx} G_{Tx}}{4\pi r_{Tx}^2 B_{Tx}}$	Effective Irradiated Power ( $P_{Tx} G_{Tx}$ )	57 dBW	-171.1 dB [W/m <sup>2</sup> /Hz]
	Transmit Range	36000 km	
	Transmit Bandwidth	4 MHz	
<b>Power at Receiver Satellite</b> $P_{rec} = \frac{P_{ground} B_{Rx}}{4\pi r_{Rx}^2} \cdot A_{Rx} \cdot A_{res} \sigma_0$	Receiver Bandwidth	4 MHz	-232.5 dB [W]
	Receive Range	37000 km	
	Receiver Antenna Area	20 m <sup>2</sup>	
	Ground Resolution	100 m x 100 m	
	Sigma	-18 dB m <sup>2</sup> /m <sup>2</sup>	
<b>SNR</b> $SNR = \frac{P_{rec}}{kTFB} = \frac{P_{rec}}{kTF} \cdot t_{int}$	Receiver Noise Figure + Losses	7 dB	<b>7.1 dB</b>  (NESZ: -25.1 dB)
	Receiver Temperature	290 K	
	Integration Time	5 h	

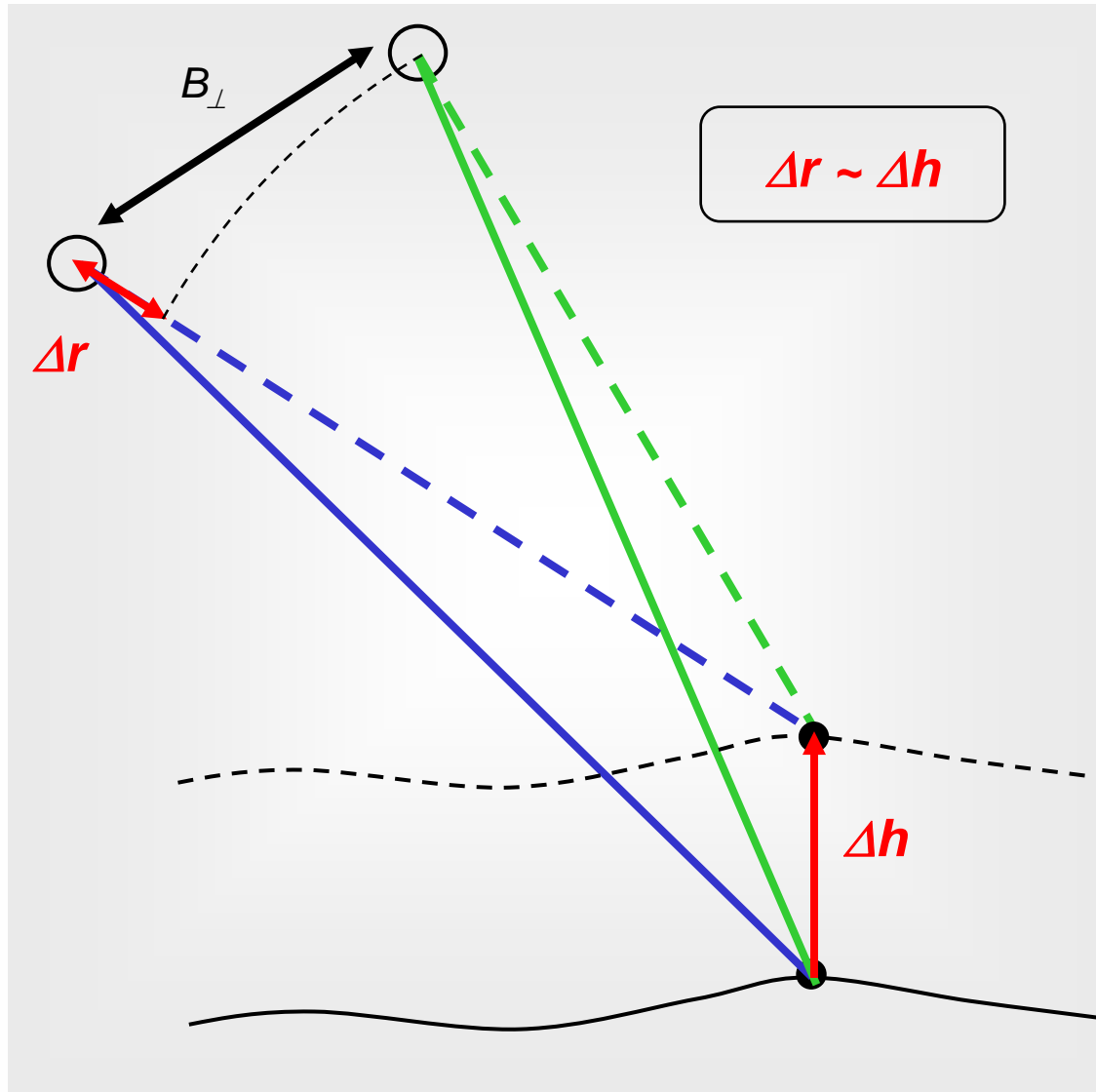
# Multistatic SAR System Concepts



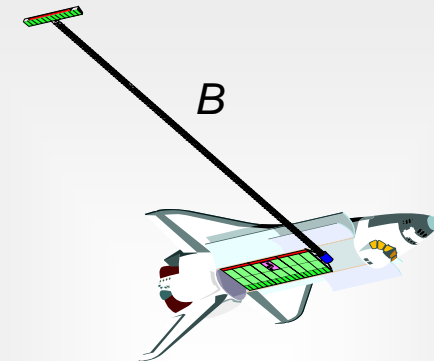
## Combination of multiple receiver signals enables:

- *cross-track interferometry for cost efficient acquisition of high quality global DEMs*
- *along-track interferometry (e.g. oceanography) & moving target indication*
- *increased geometric resolution by super-resolution techniques in azimuth and range*
- *retrieval of vegetation and volume parameters by polarimetric interferometry*
- *real 3-D imaging of semitransparent volume scatterers by SAR tomography*
- *ambiguity reduction and high resolution wide swath SAR imaging*
- *improved classification (e.g. joint evaluation of multiple mono- and bistatic RCS)*
- ...

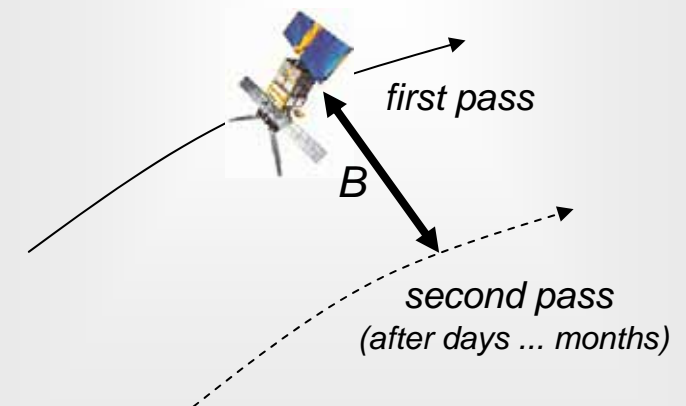
# Basic Principle of Cross-Track Interferometry



single-pass:

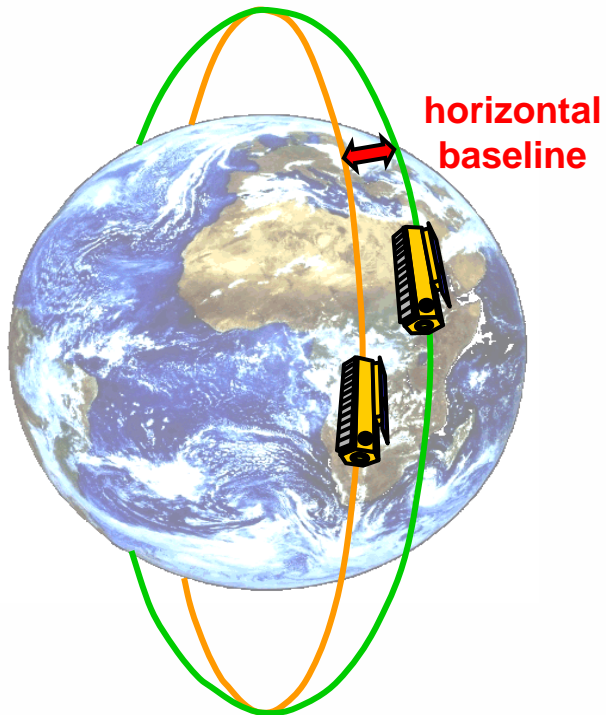


repeat-pass:

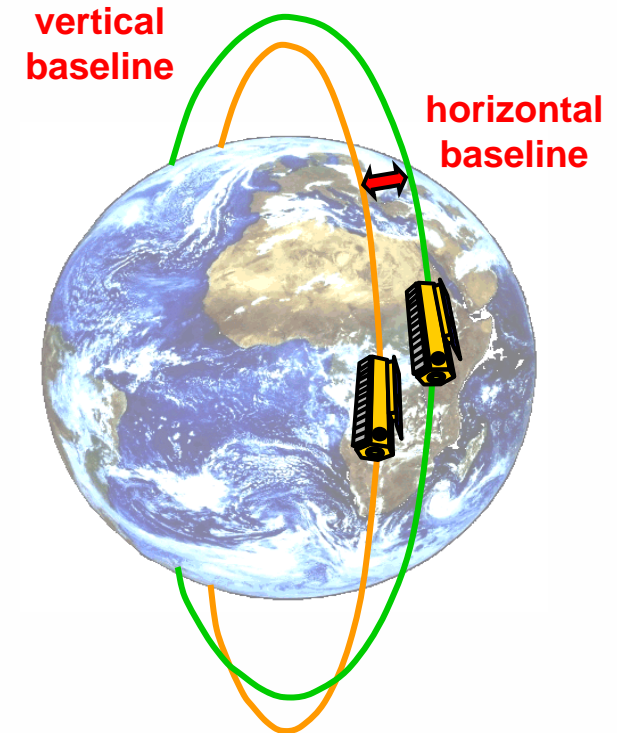




# Single-Pass Cross-Track Interferometry



Digital Elevation Model (DEM)  
from SRTM Mission



## **Single-Pass Cross-Track Interferometry with Multiple Satellites:**

- ***no temporal decorrelation*** (as opposed to repeat-pass interferometry)
- ***no atmospheric distortions*** (as opposed to repeat-pass interferometry)
- ***large interferometric baselines*** (as opposed to e.g. SRTM)

# Relative Movement in Satellite Clusters

- Relative satellite movement is described in a **rotating reference frame**
- **Linearization** of the equations of motions in a circular Kepler orbit leads to Clohessy-Wiltshire (or Hill's) Equations:

$$\ddot{x} - 2n\dot{y} - 3n^2 x = 0$$

$$\ddot{y} + 2n\dot{x} = 0$$

$$\ddot{z} + n^2 z = 0$$

$$\text{with } n = \sqrt{\frac{GM_{\oplus}}{a_{sat}^3}}$$

- Solution to Clohessy-Wiltshire Equations :

$$\text{with } T_0 = 2\pi \sqrt{\frac{a_{sat}^3}{GM_{\oplus}}}$$



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in der Helmholtz-Gemeinschaft

$$x_i(t) = A_i \sin\left(\frac{2\pi}{T_0} t + \alpha_i\right)$$

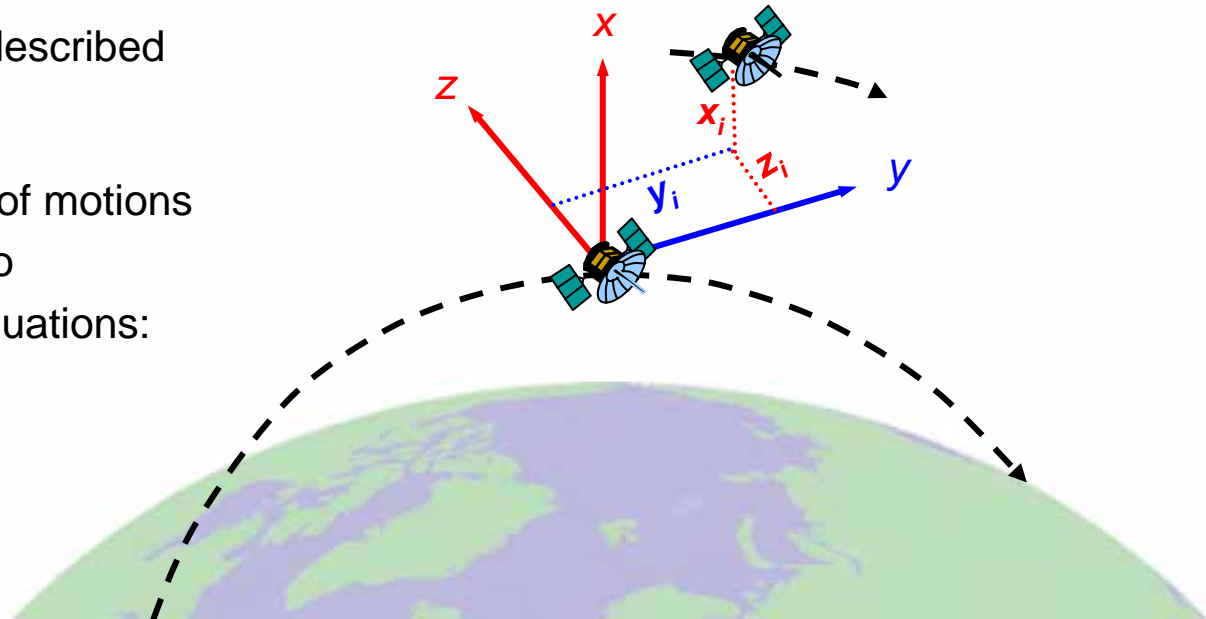
$$y_i(t) = 2A_i \cos\left(\frac{2\pi}{T_0} t + \alpha_i\right) + \Delta y_i$$

$$z_i(t) = B_i \sin\left(\frac{2\pi}{T_0} t + \beta_i\right)$$

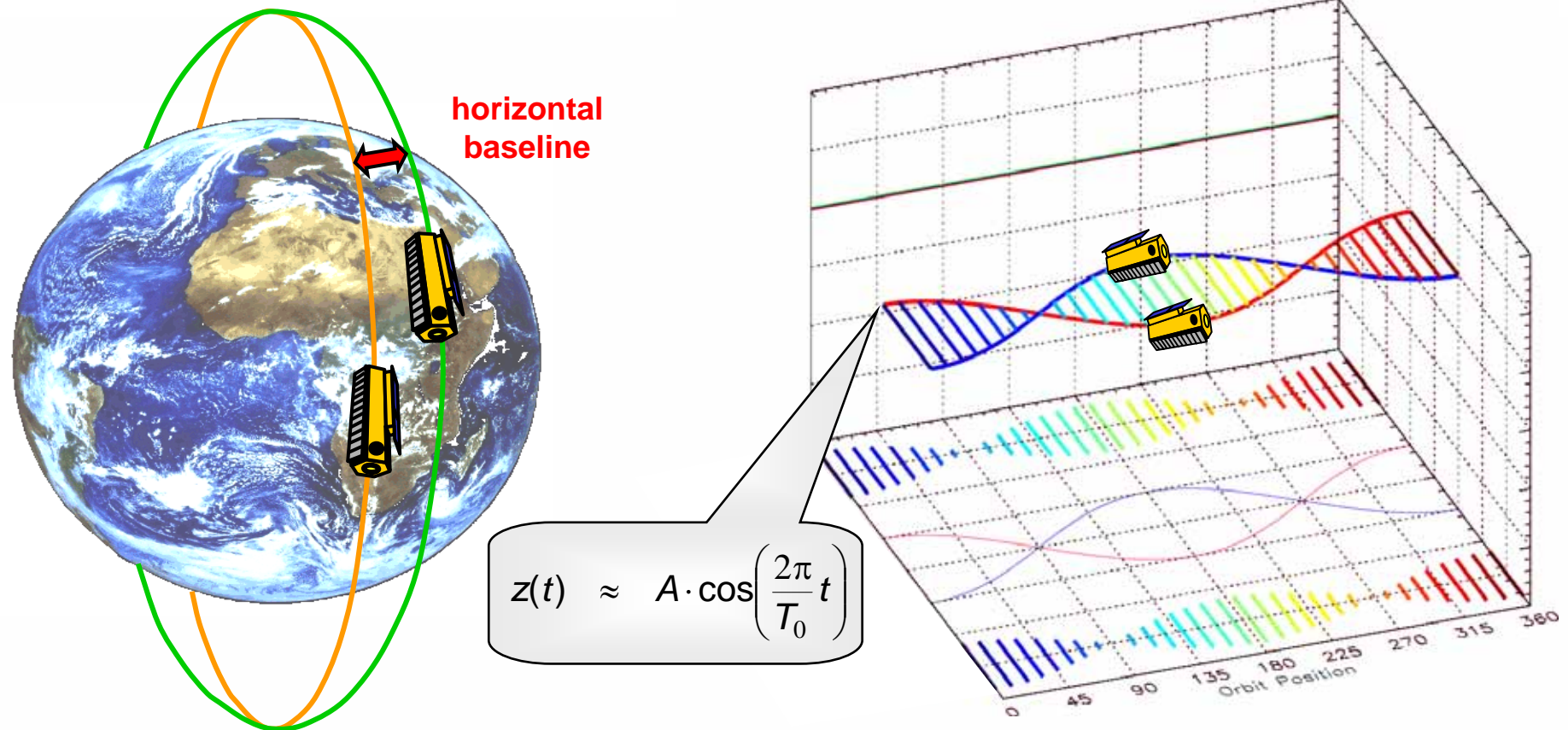
vertical cross-track baseline

along-track displacement

horizontal cross-track baseline



# 2-SAT Pendulum



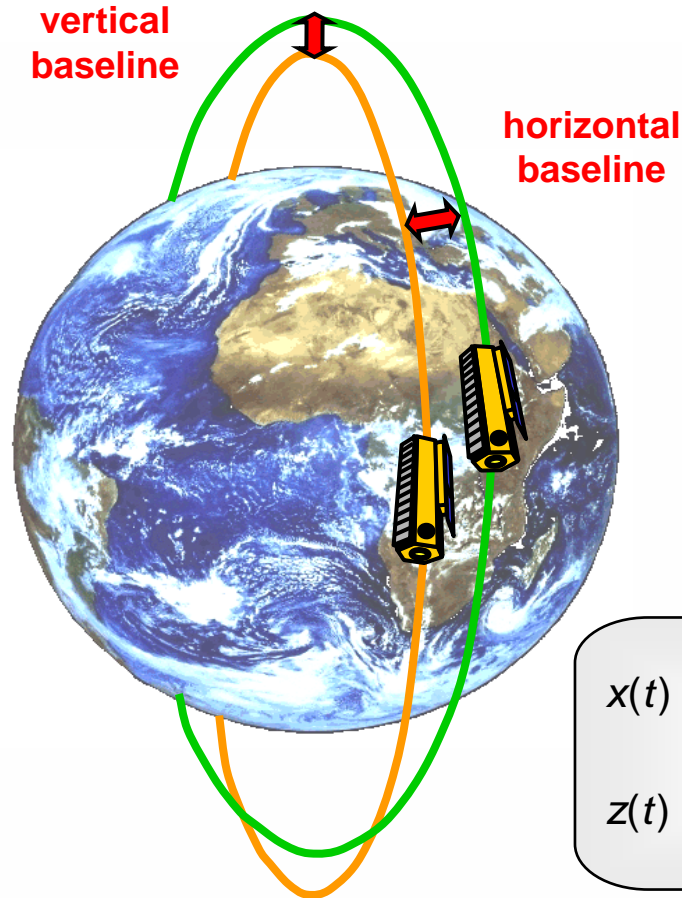
## 2-SAT Pendulum (Hartl 1989, Zebker, 1992)

- *horizontal cross-track separation at equator by different ascending nodes*
- *requires along-track displacement to avoid satellite collision at orbit crossing*
- *insufficient baselines for polar regions*



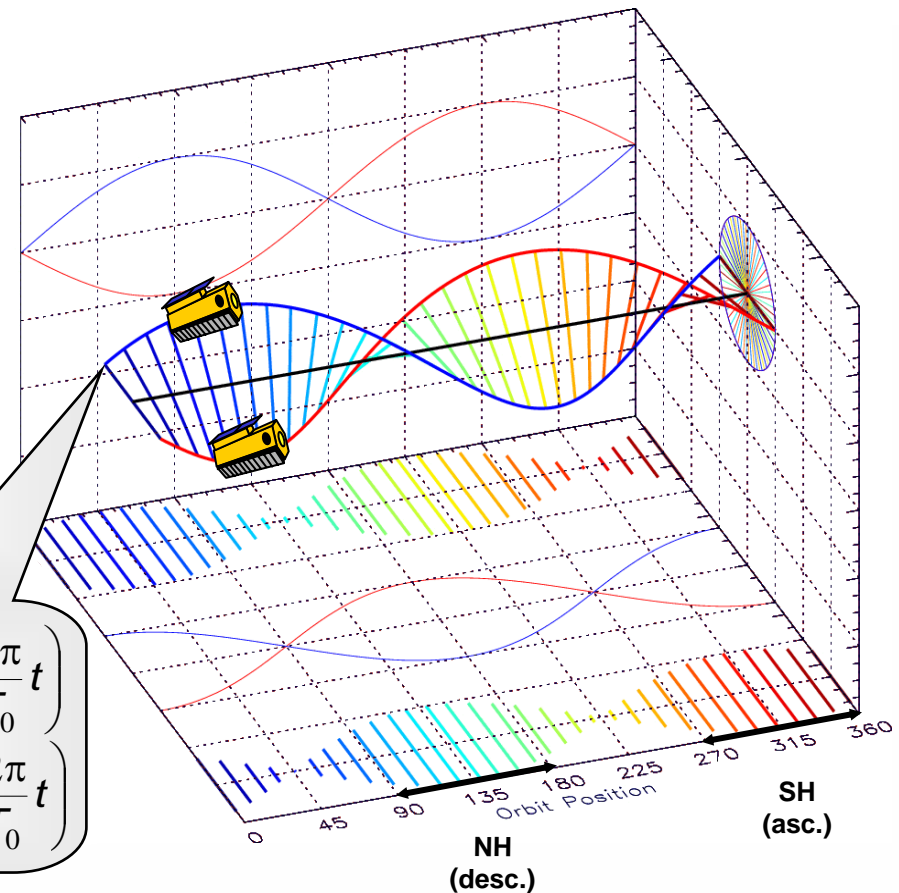


# HELIX Formation



$$x(t) \approx A \cdot \sin\left(\frac{2\pi}{T_0} t\right)$$

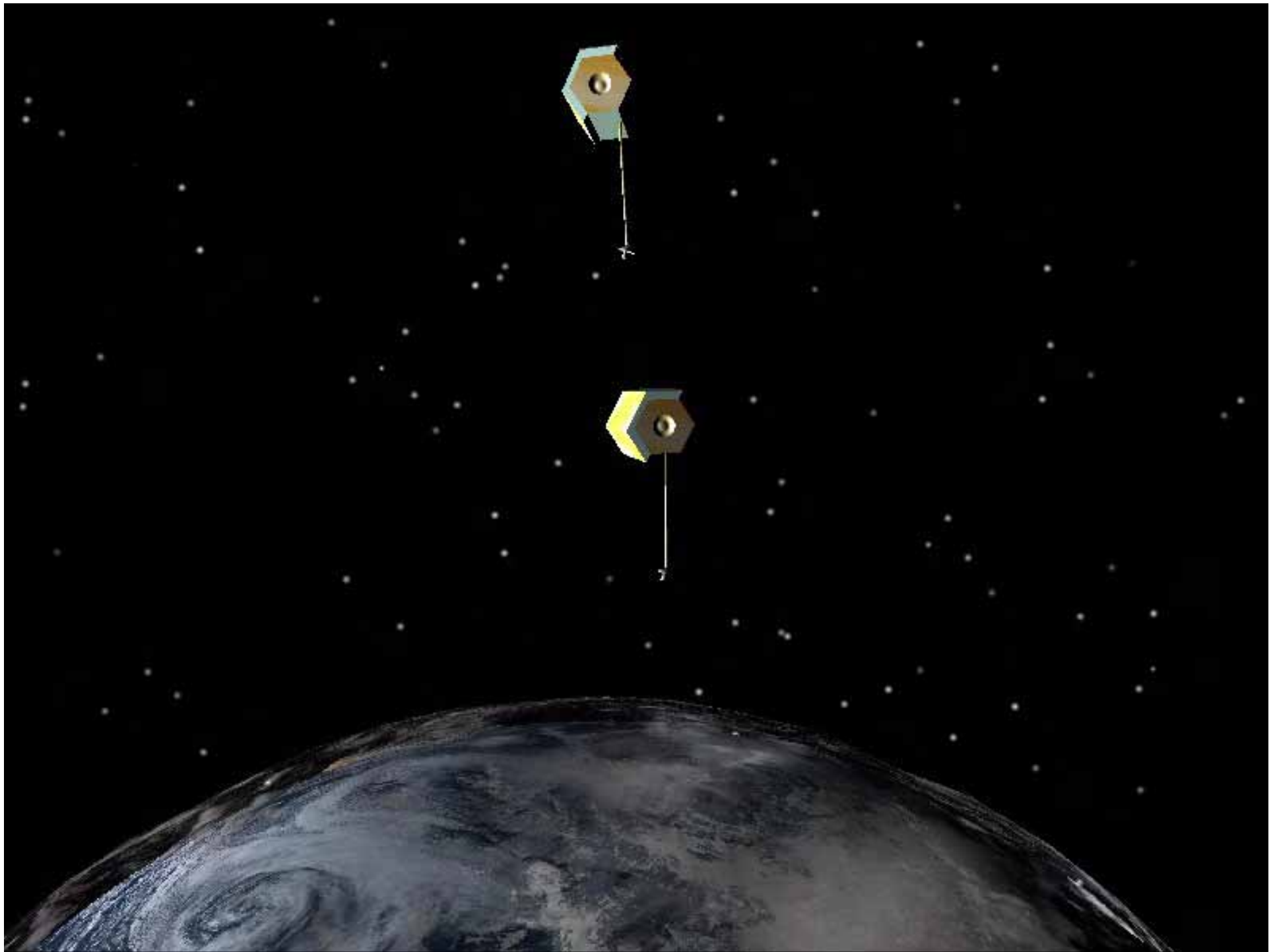
$$z(t) \approx B \cdot \cos\left(\frac{2\pi}{T_0} t\right)$$



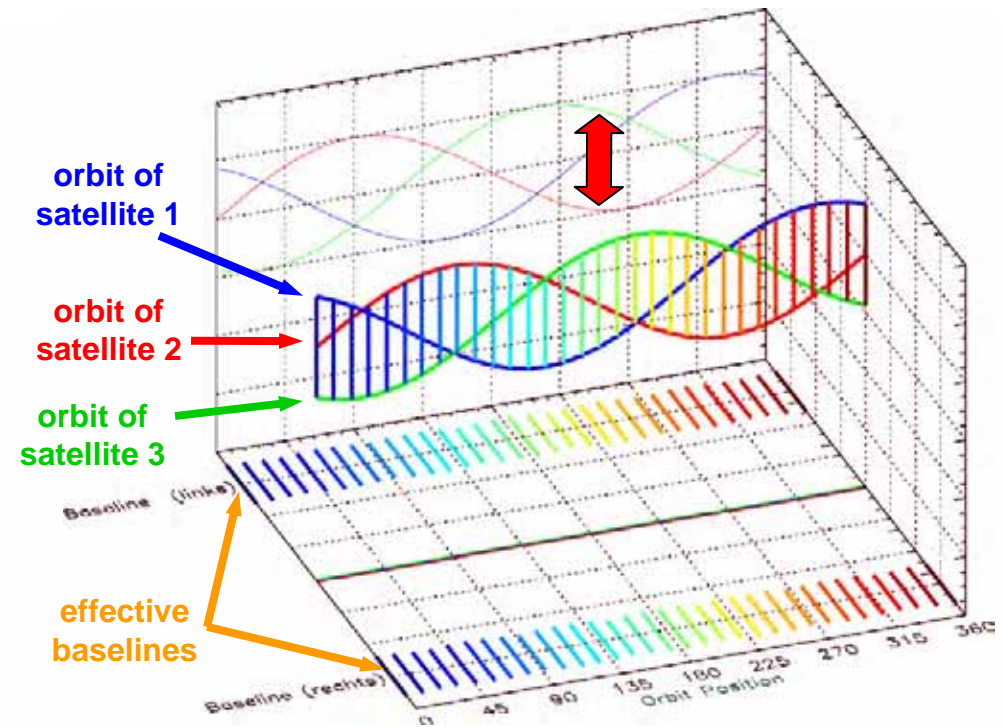
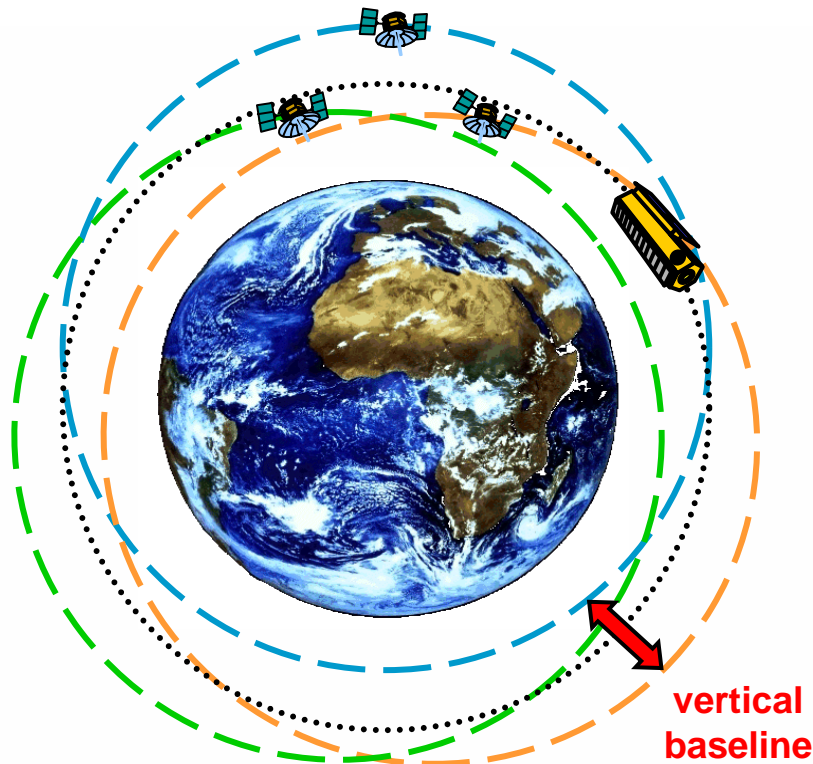
## **HELIX satellite formation enables safe operation**

- *horizontal cross-track separation at equator by different ascending nodes*
- *vertical (radial) separation at poles by orbits with different eccentricity vectors (periodic motion of libration has to be compensated by regular manoeuvres)*





# Interferometric Cartwheel

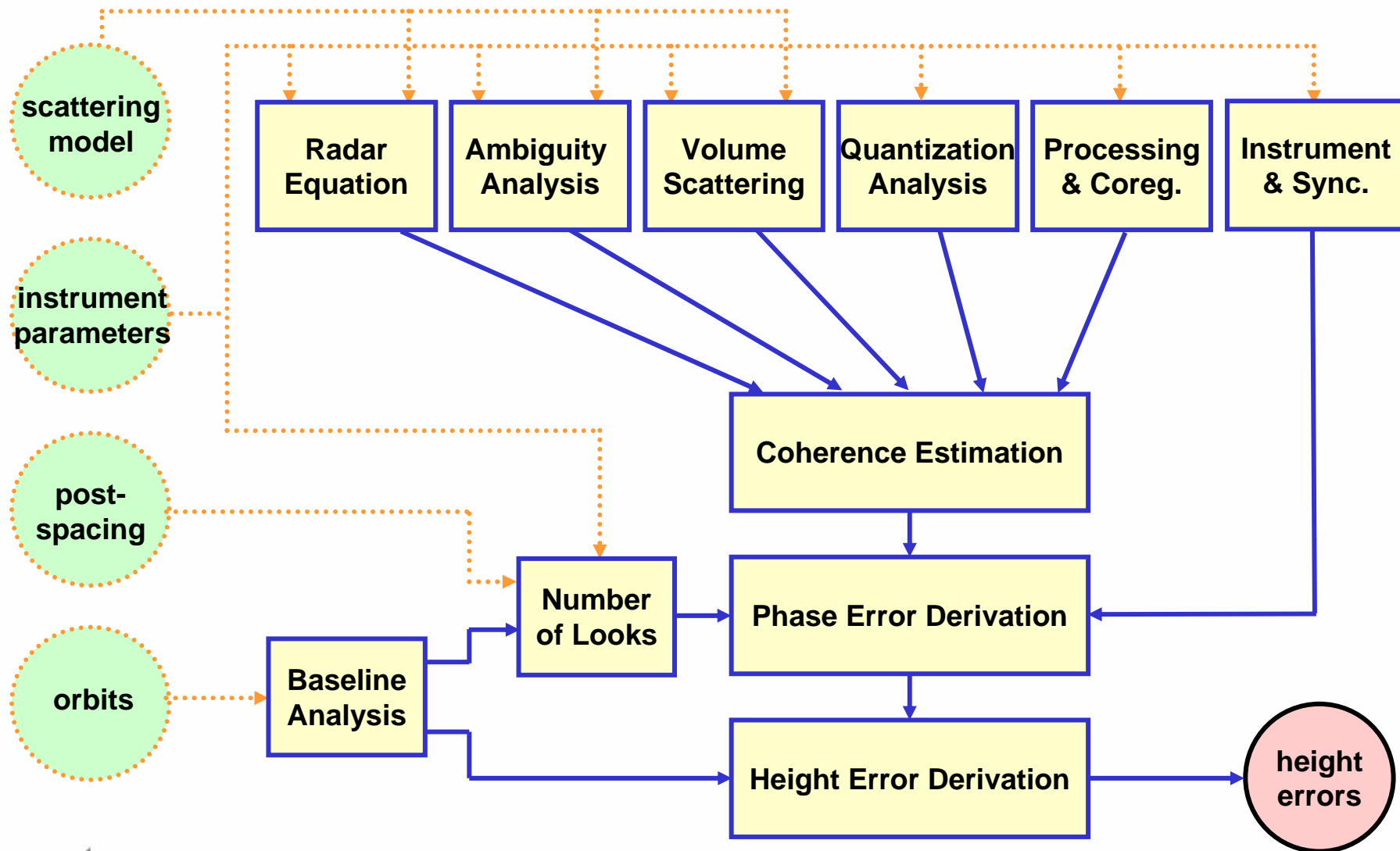


## **Interferometric Cartwheel (Massonnet, 1998)**

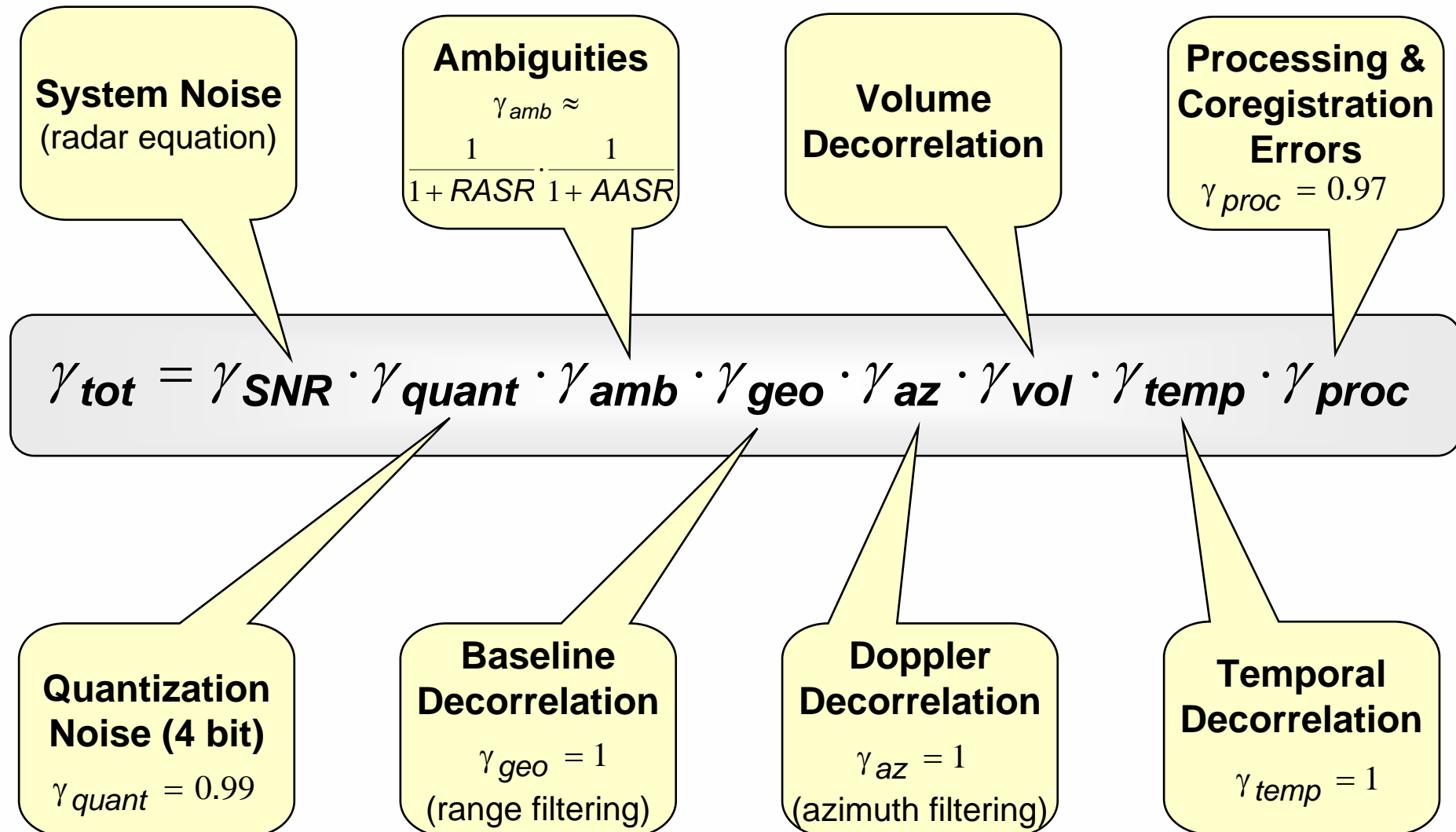
- *all satellites share the same orbital plane*
- *arguments of apogee differ by  $120^\circ$  for a Cartwheel with 3 satellites*
- *provides a stable vertical baseline for all orbit positions*
- *relative movement of the receiver satellites can be approximated by an ellipse*



# Interferometric Performance Analysis



# Coherence Estimation





# Computation of Coherence Loss from Limited SNR

Computation of Noise Equivalent Sigma Zero:

$$NESZ = \frac{4^4 \pi^3 r^3 v \sin(\theta_{inc}) k T B_{rg} FL}{P_{Tx} G_{Tx} G_{Rx} \lambda^3 c_0 \tau_p PRF}$$

Derivation of Single Channel SNR:

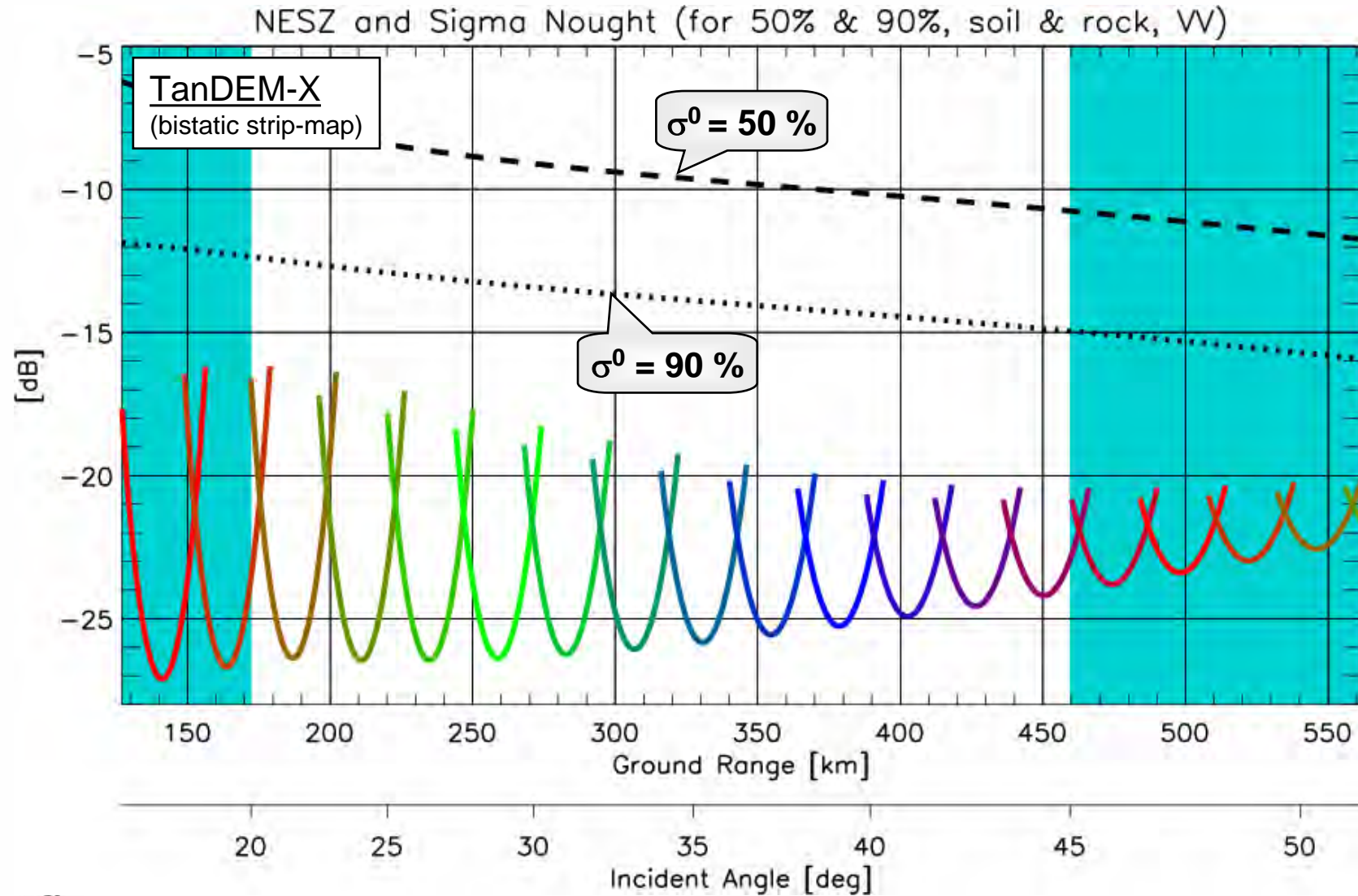
$$SNR[dB] = \sigma_0[dB] - NESZ[dB]$$

Derivation of Coherence:

$$\gamma_{SNR} = \frac{1}{\sqrt{1 + SNR_1^{-1}} \cdot \sqrt{1 + SNR_2^{-1}}}$$

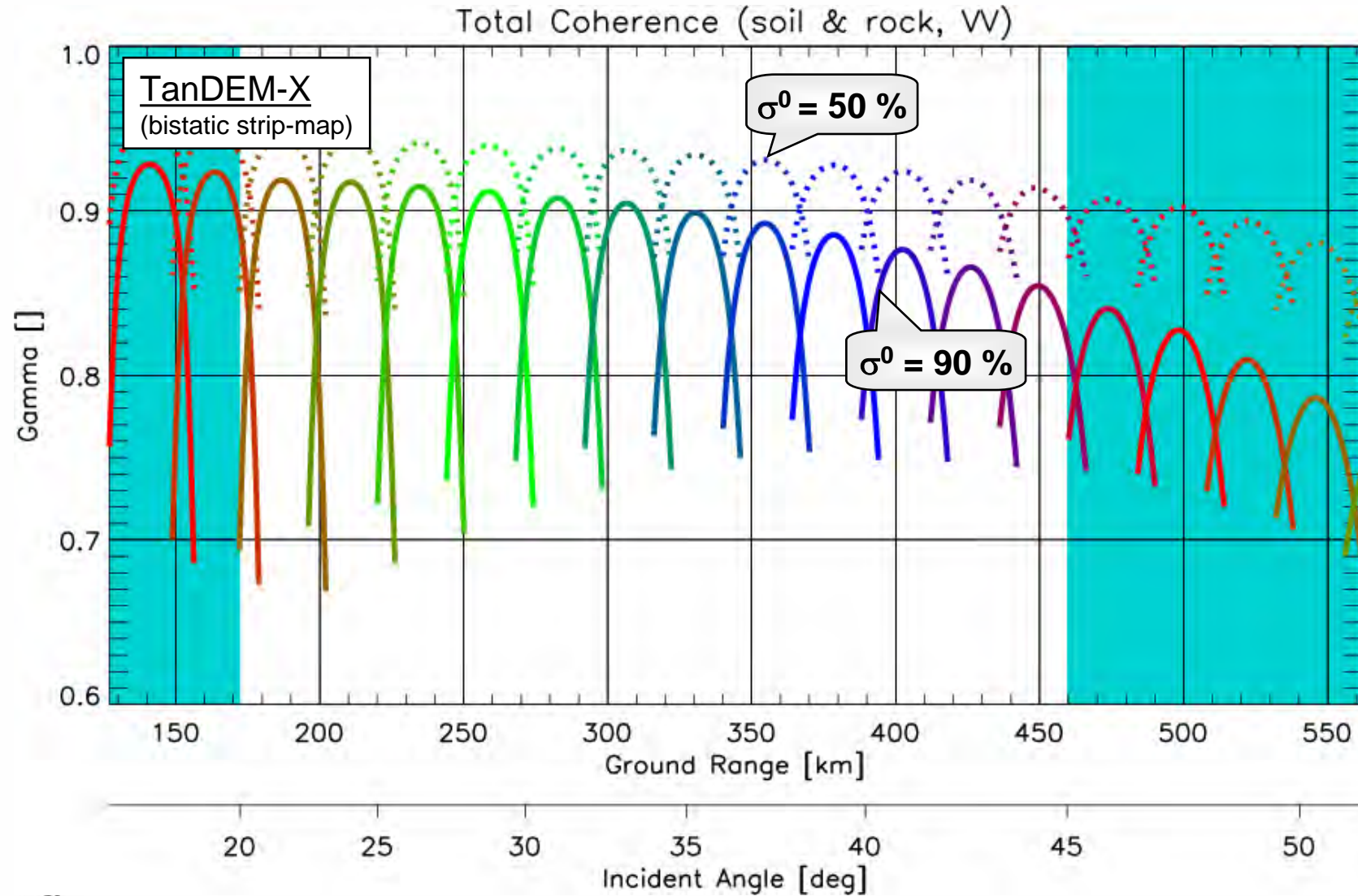
$r$  : slant range  
 $v$  : satellite velocity  
 $\theta_{inc}$  : incident angle  
 $k$  : Boltzman constant  
 $T$  : system temperature (290K)  
 $B_{rg}$  : chirp bandwidth  
 $F$  : system noise figure  
 $L$  : losses  
 $P_{Tx}$  : peak transmit power  
 $G_{Tx}$  : gain of transmit antenna  
 $G_{Rx}$  : gain of receive antenna  
 $\lambda$  : wavelength  
 $c_0$  : velocity of light  
 $\tau_p$  : pulse duration  
 $PRF$  : pulse repetition frequency  
 $\sigma_0$  : backscatter coefficient

# Noise Equivalent Sigma Zero (Example)





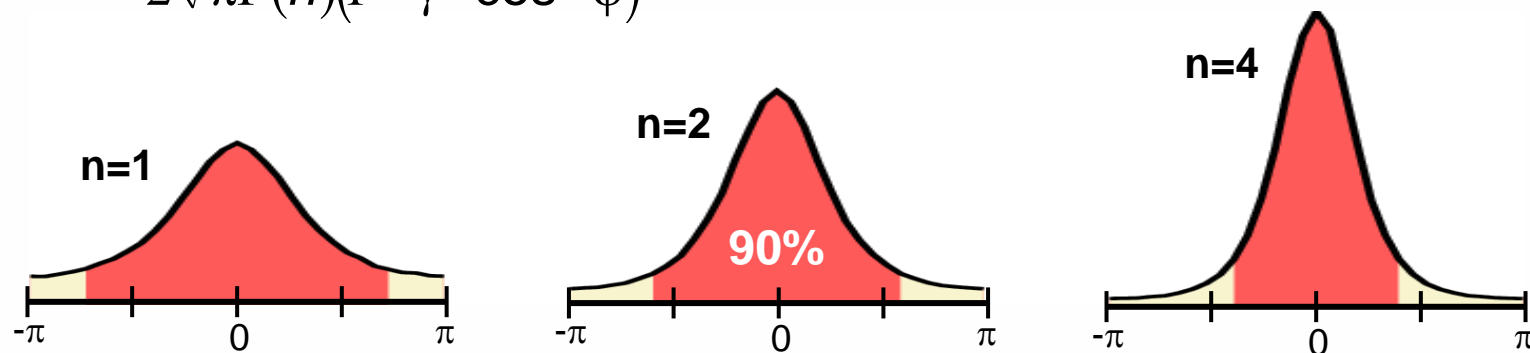
## Total Coherence (Example)



# Derivation of Interferometric Phase Errors

## Phase Error PDF:

$$p_{\varphi}(\varphi) = \frac{\Gamma\left(n + \frac{1}{2}\right)(1 - \gamma^2)^n \gamma \cos \varphi}{2\sqrt{\pi}\Gamma(n)(1 - \gamma^2 \cos^2 \varphi)^{n+1/2}} + \frac{(1 - n^2)^n}{2\pi} F\left(n, 1; \frac{1}{2}; \gamma^2 \cos^2 \varphi\right) \quad (\text{cf. Lee et al., 1994})$$



## Standard Deviation:

$$\sigma_{\varphi} = \sqrt{\int_{-\pi}^{\pi} \varphi^2 p_{\varphi}(\varphi) \cdot d\varphi}$$

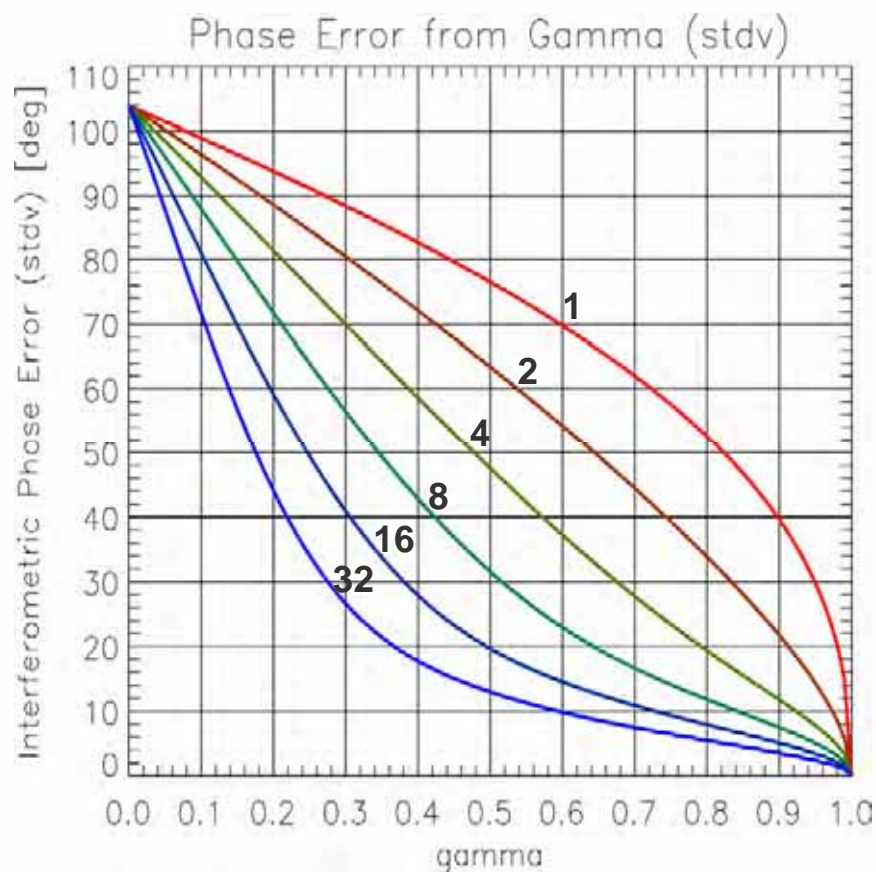
## 90 Percentile:

$$\Delta\varphi^{90\%} : \int_{-\Delta\varphi^{90\%}}^{\Delta\varphi^{90\%}} p_{\varphi}(\varphi) \cdot d\varphi = 0.9$$

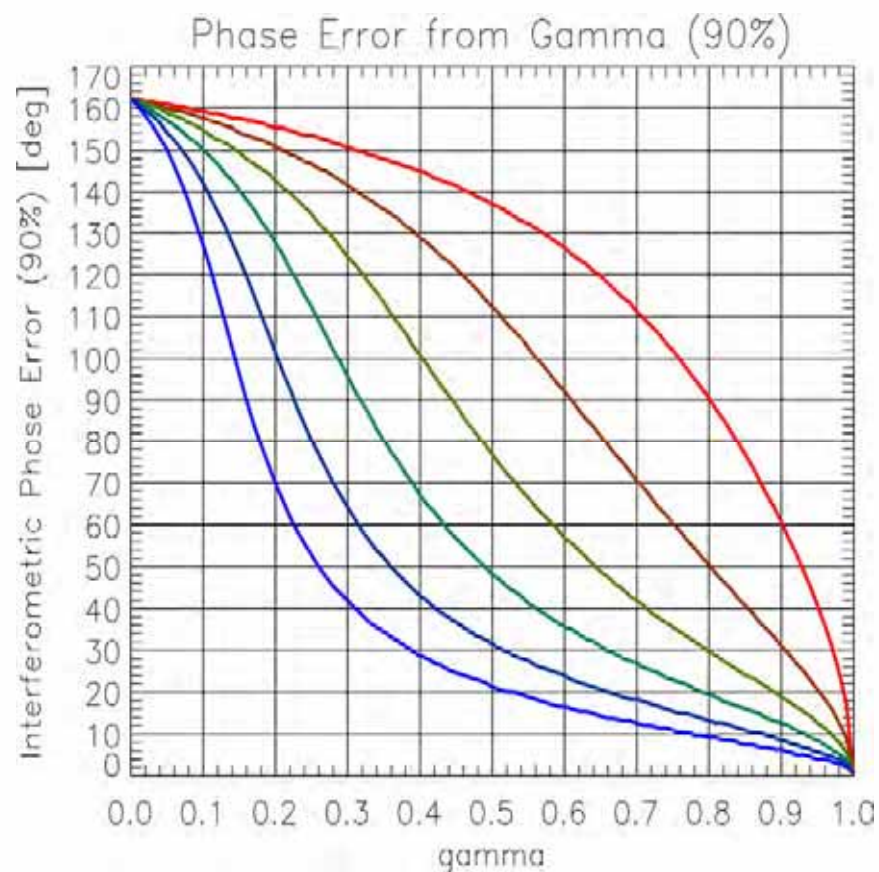


# Interferometric Phase Errors

## Standard Deviation



## 90 Percentile



# Independent Looks

## Ground Range Resolution

(with range filtering)

$$\Delta rg = \frac{c_0}{2B_{rg} \sin(\theta_{inc} - \alpha)} \cdot \frac{B_{\perp, crit}}{B_{\perp, crit} - B_{\perp}}$$

with critical baseline

$$B_{\perp, crit} = \frac{2B_{rg} \lambda r \tan(\theta_i - \alpha)}{c_0}$$

and

- $c_0$  : velocity of light
- $B_{rg}$  : chirp bandwidth
- $\lambda$  : wavelength
- $r$  : slant range
- $\theta_i$  : incident angle
- $\alpha$  : local slope
- $B_{\perp}$  : effective baseline

## Azimuth Resolution

(with azimuth filtering)

$$\Delta az = \frac{v_{grd}}{B_{proc} - \Delta f} \quad \text{with } \Delta f = |f_{Dop,1} - f_{Dop,2}|$$

with Doppler Centroid approximation

$$f_{Dop,i} = \frac{1}{\lambda} \left[ \frac{\mathbf{v}_{Tx} \cdot \mathbf{p}_{Tx}}{\|\mathbf{p}_{Tx}\|} + \frac{\mathbf{v}_{Rx,i} \cdot \mathbf{p}_{Rx,i}}{\|\mathbf{p}_{Rx,i}\|} \right]$$

and

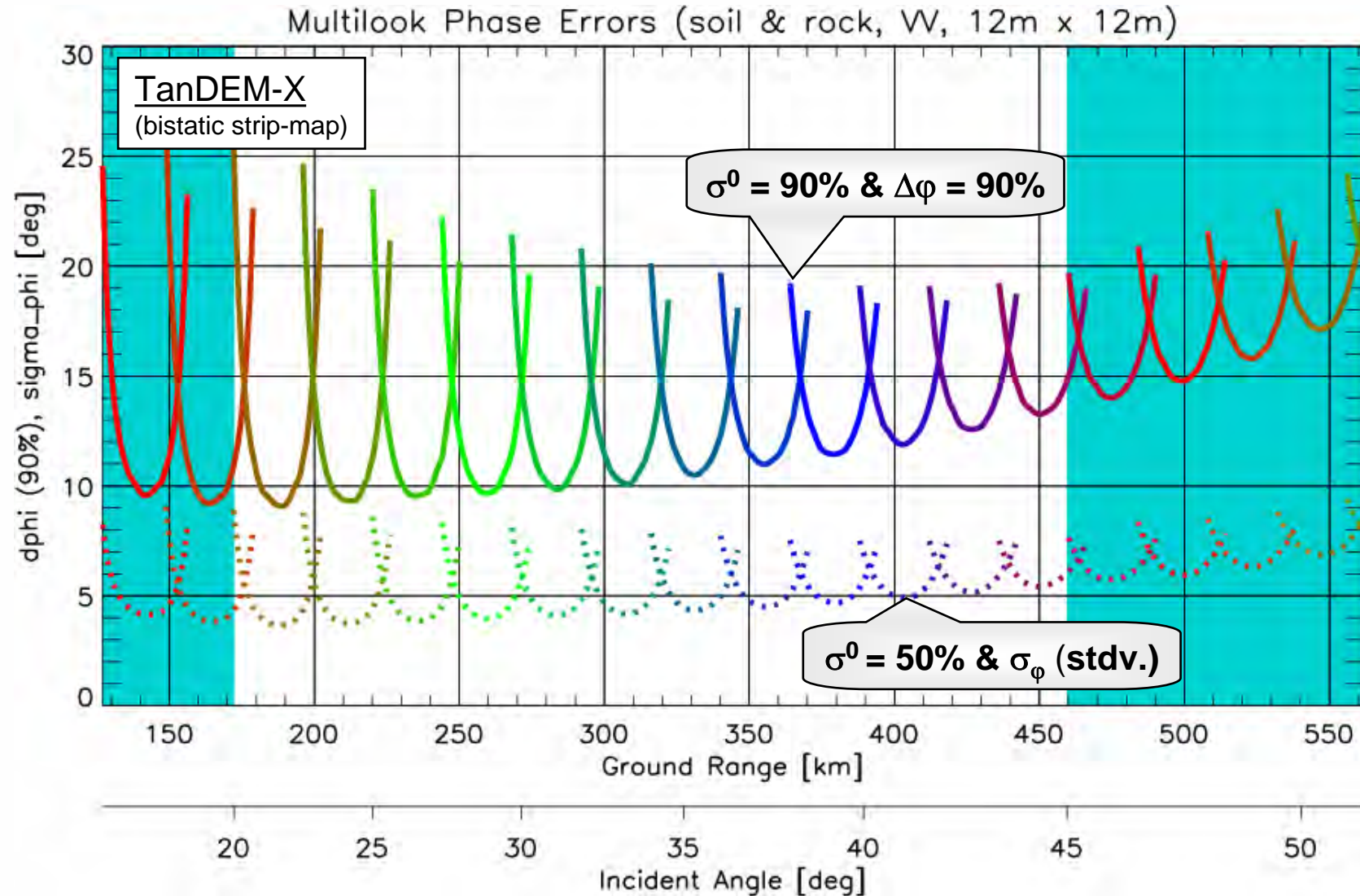
- $v_{grd}$  : satellite velocity on ground
- $B_{proc}$  : processed bandwidth ( $B = 2266 \text{ Hz}$ )
- $\Delta f$  : relative shift of Doppler Centroids
- $\lambda$  : wavelength
- $\mathbf{v}_{\{Tx,Rx\}}$  : velocity vector (transmitter/receiver)
- $\mathbf{p}_{\{Tx,Rx\}}$  : vector from satellite to scene center

$$N_{looks} = \frac{\Delta x}{\Delta rg} \cdot \frac{\Delta y}{\Delta az} \quad \text{with independent post spacing } \Delta x \text{ and } \Delta y$$





# Multilook Interferometric Phase Errors (Example)



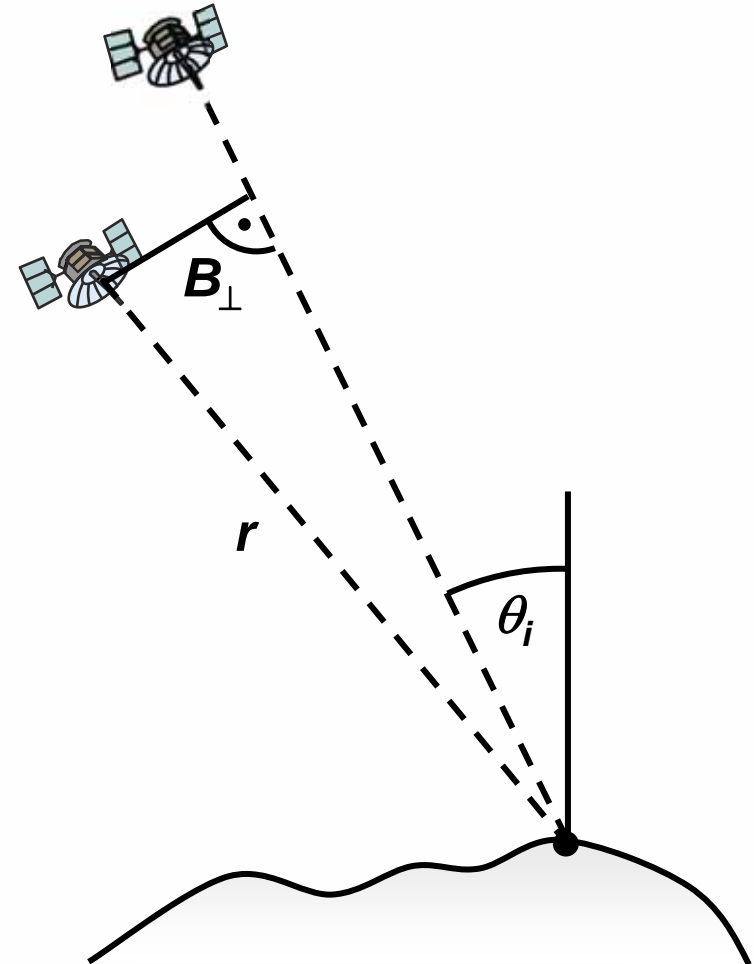
## Derivation of Relative Height Errors

Standard deviation of height errors:

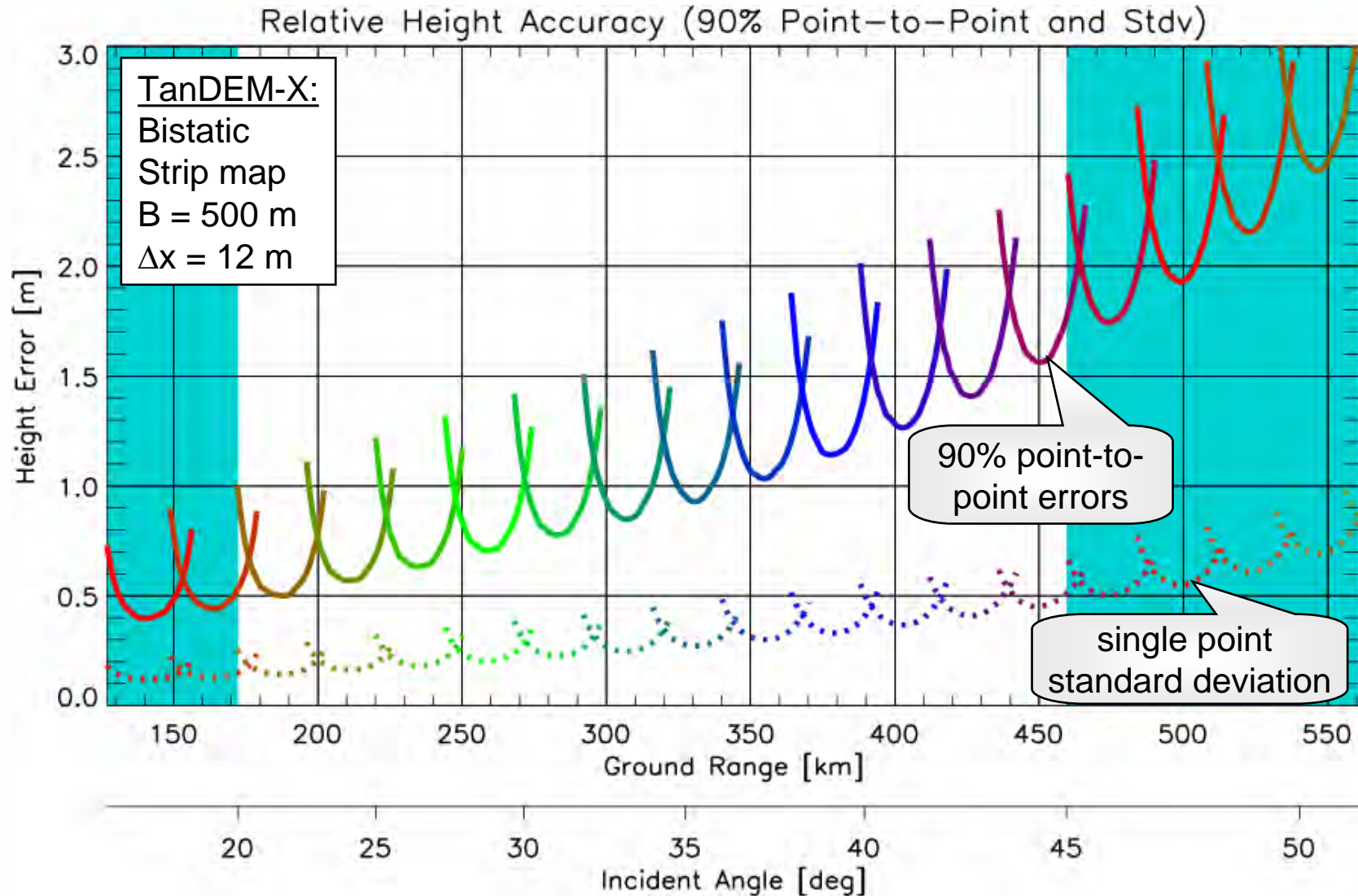
$$\sigma_h = \frac{\lambda r \sin(\theta_i)}{2\pi B_{\perp}} \cdot \sigma_{\varphi} \quad (\text{flat terrain})$$

HRTI-3 requires point-to-point height errors at 90% confidence levels:

$$\Delta h_{\text{HRTI}} \approx \frac{\lambda r \sin(\theta_i)}{2\pi B_{\perp}} \cdot \varphi_{90\%} \cdot \sqrt{2}$$

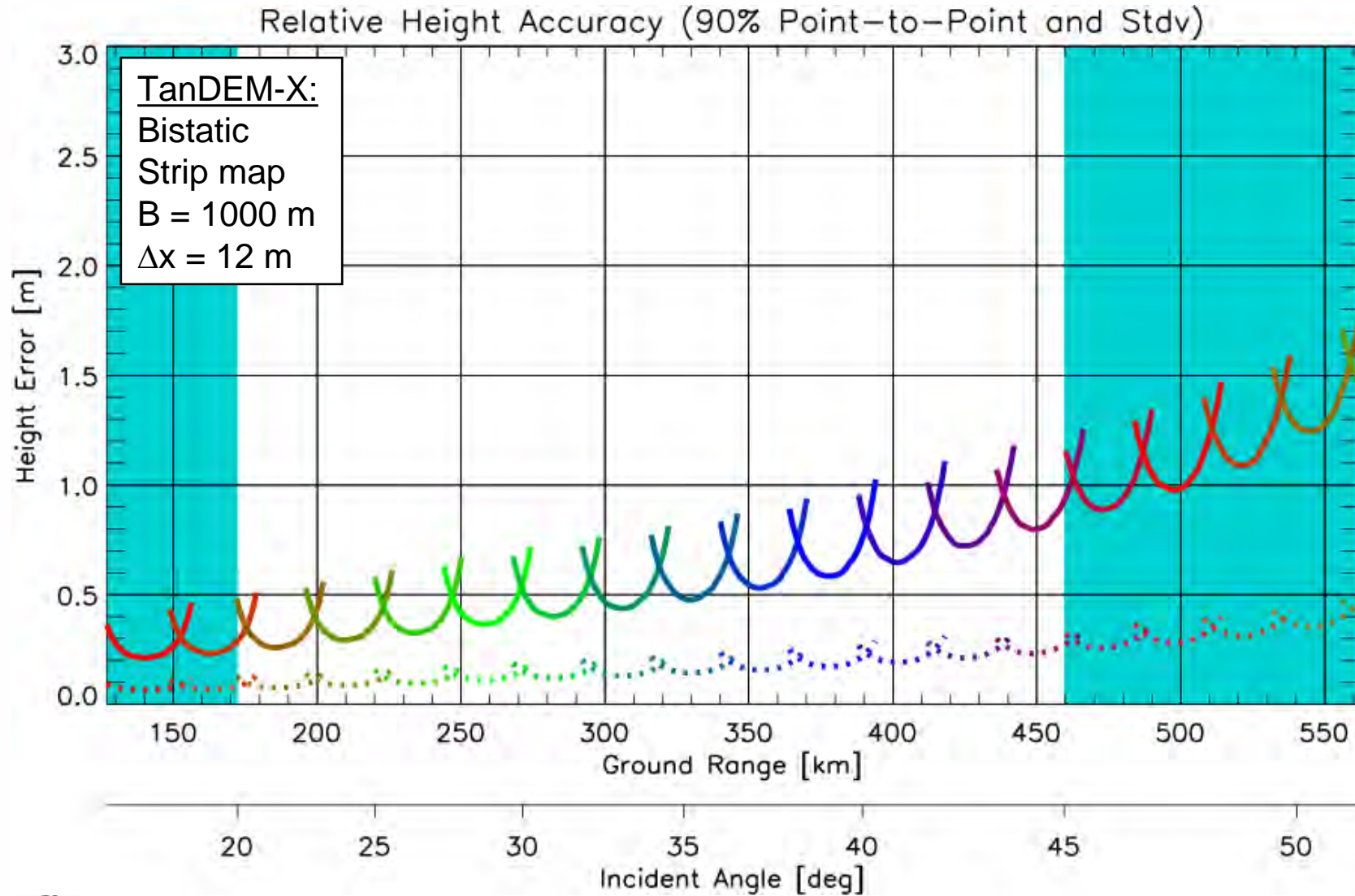


## Relative Height Errors (Example)





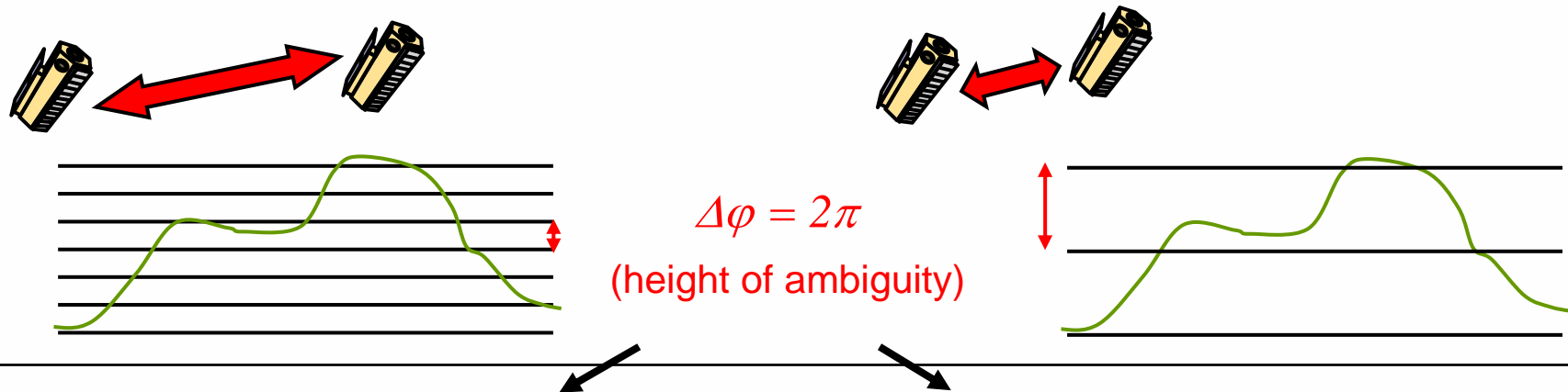
## Relative Height Errors (Example)





## Relative Height Errors (Example)

TanDEM-X enables large baselines which allow for ultra high resolution DEMs with height accuracies in the sub-meter range, but ...



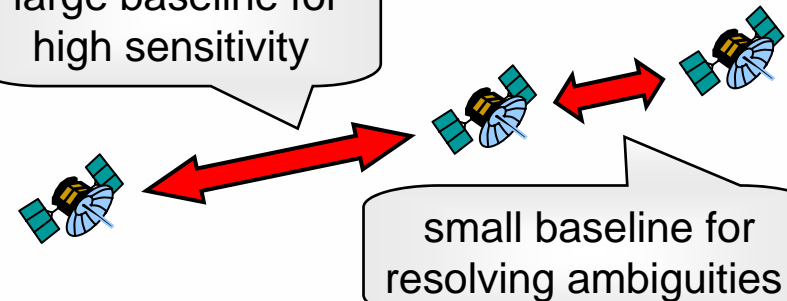
### Compromise on Accuracy for Global DEM



- use reduced baselines
  - additional acquisitions for difficult terrain
- acquisition scenario for global DEM

### Multi-Baseline Data Acquisitions

large baseline for high sensitivity



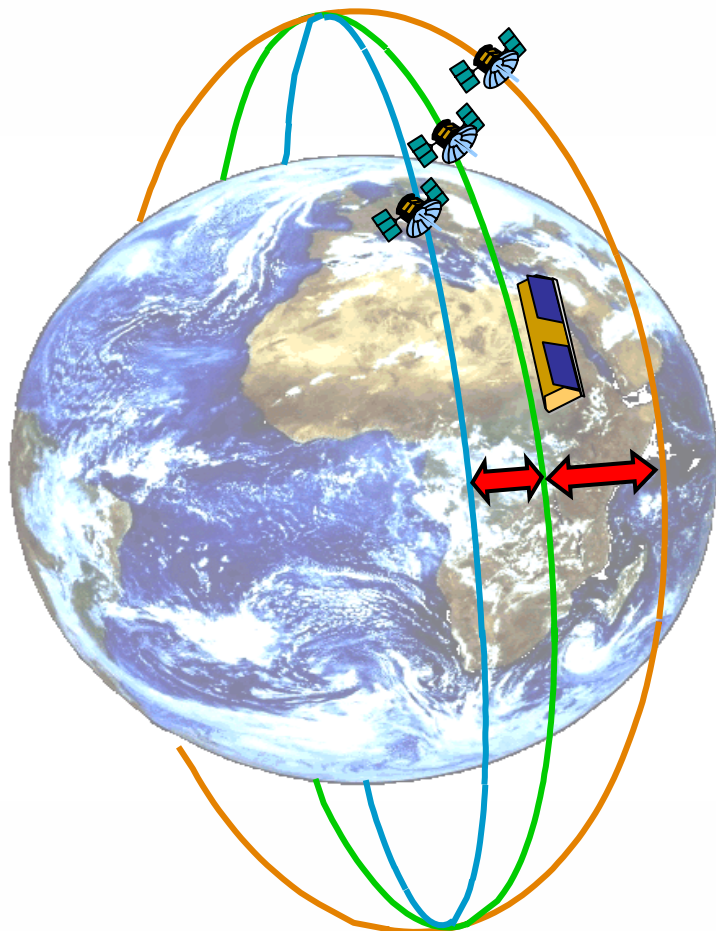
small baseline for resolving ambiguities





# Multi-Baseline Data Acquisitions (Example)

## Trinodal Pendulum

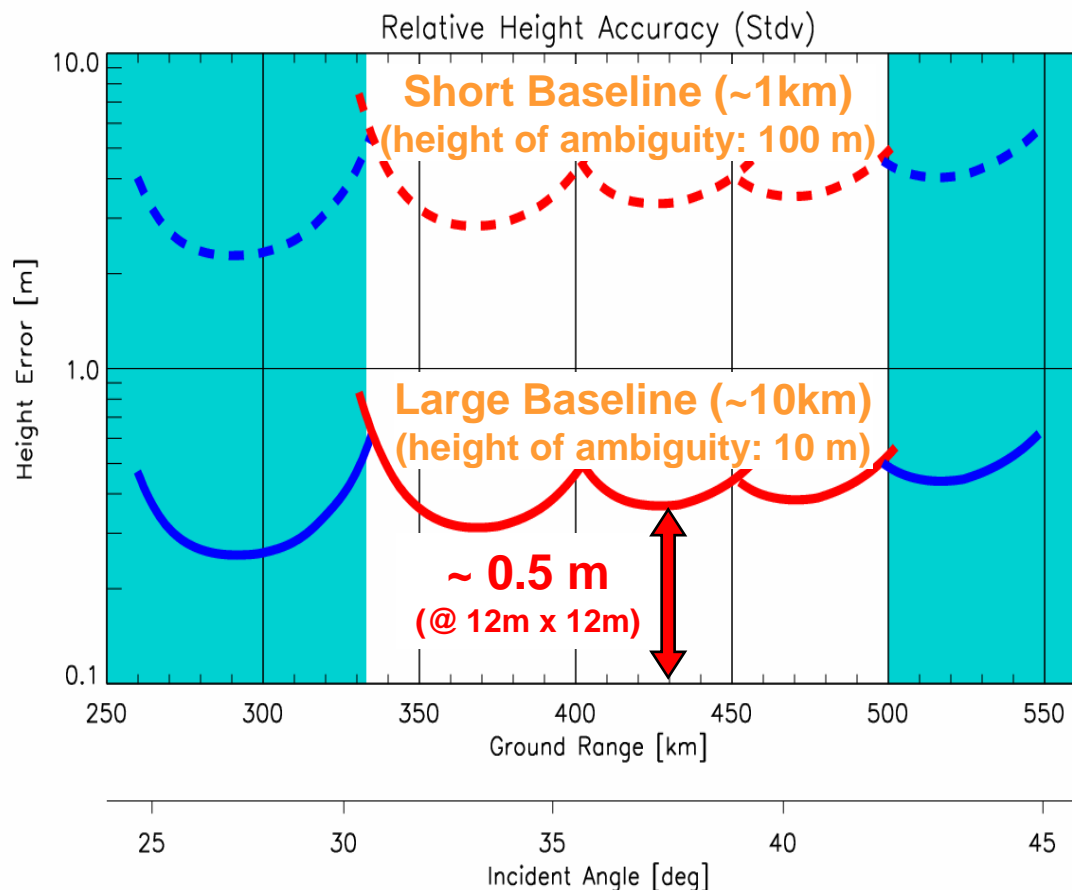


→ *multiple baselines with fixed baseline ratio in one pass*



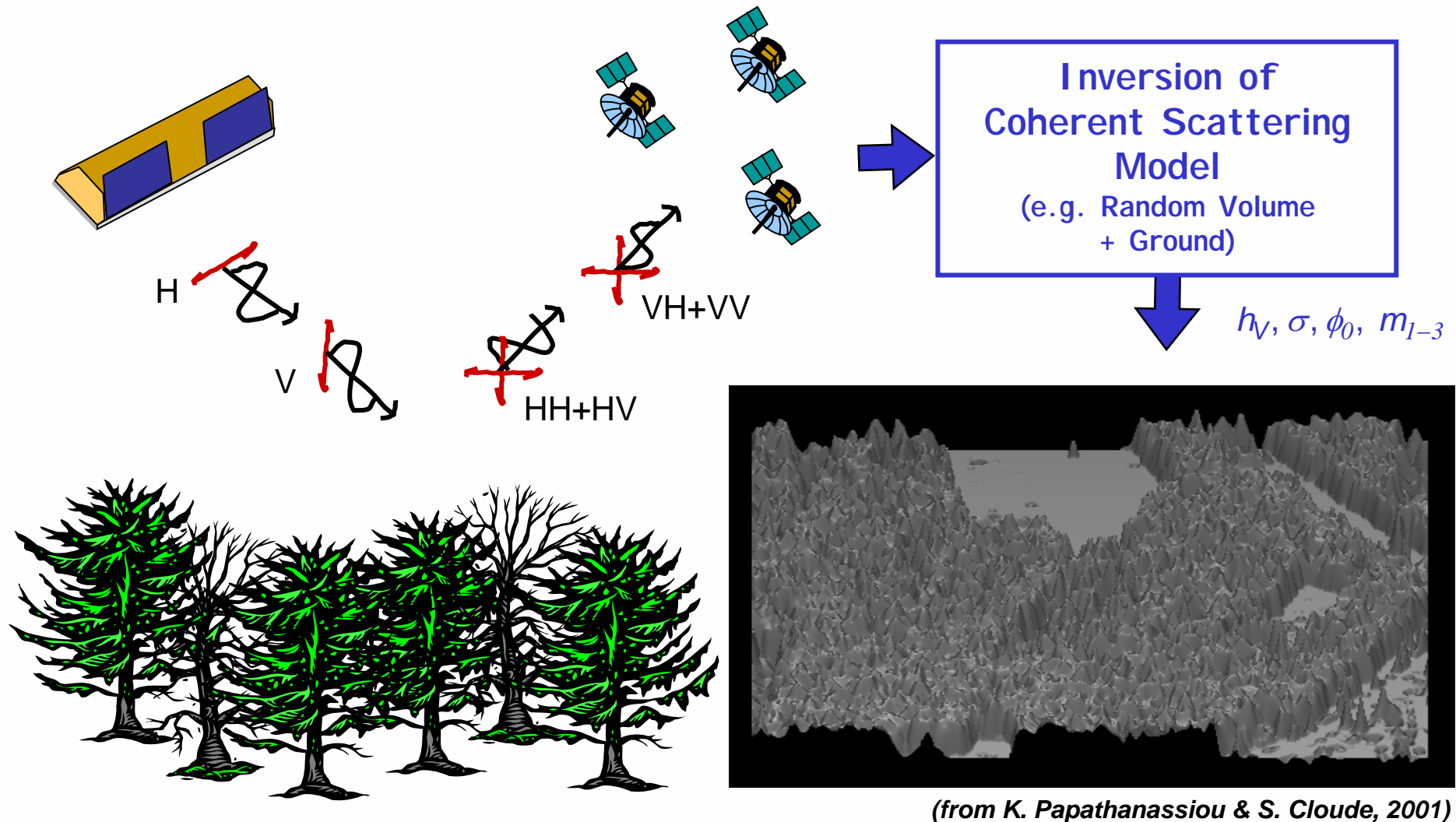
Deutsches Zentrum  
für Luft- und Raumfahrt e.V.  
in der Helmholtz-Gemeinschaft

## Performance Example (TerraSAR-L)



**Large Baseline → Excellent Height Accuracy**  
**Small Baseline → Easy Phase Unwrapping**

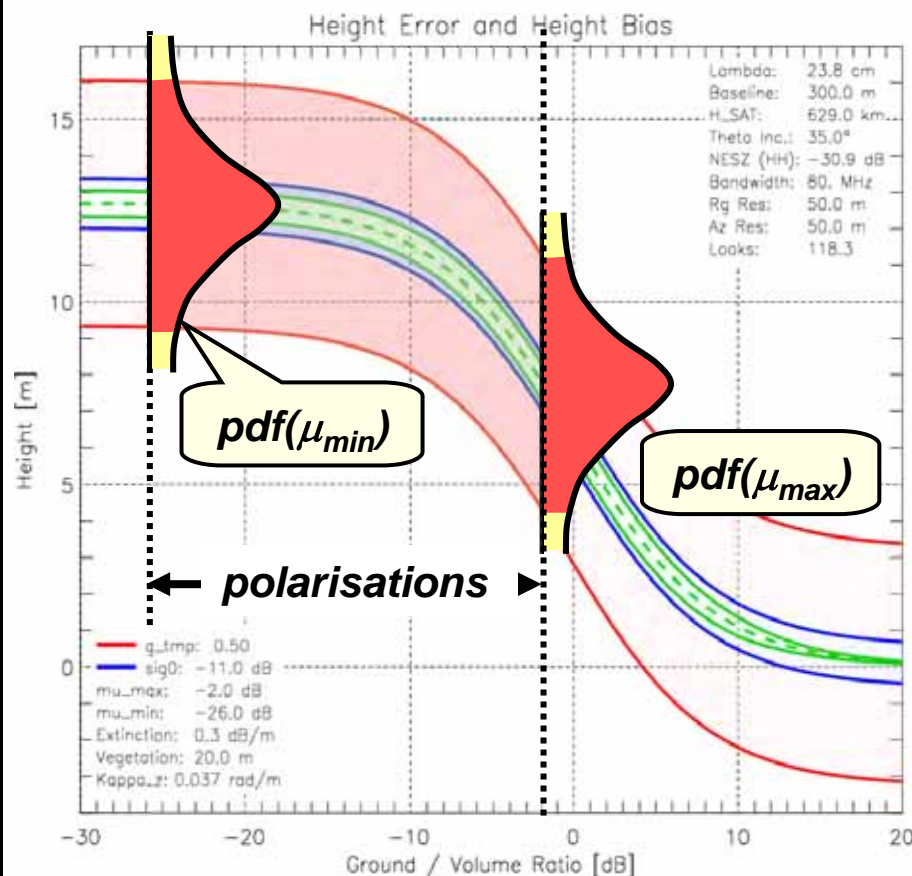
# Polarimetric SAR Interferometry



# PollnSAR L-Band Performance Example

## Monostatic Repeat Pass System

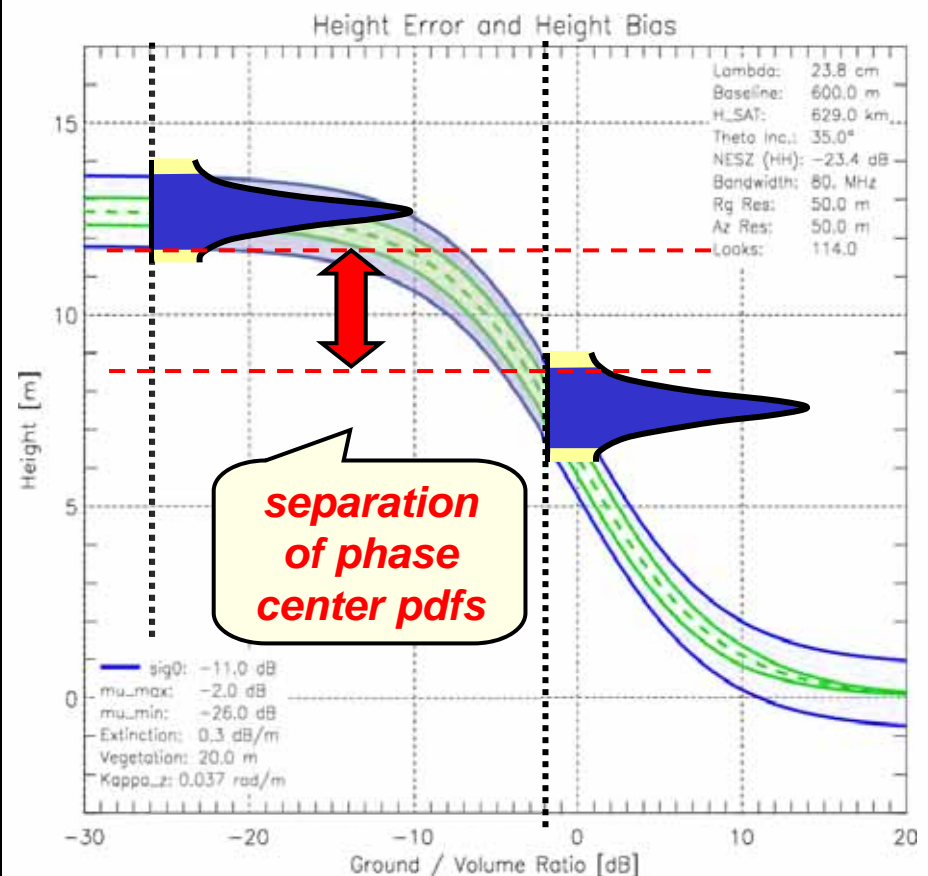
( $\gamma_{temp} = 0.5$ )



→ Poor Inversion Accuracy

## Multistatic Single Pass System

( $\gamma_{temp} = 1.0$ , but smaller Rx antennas)



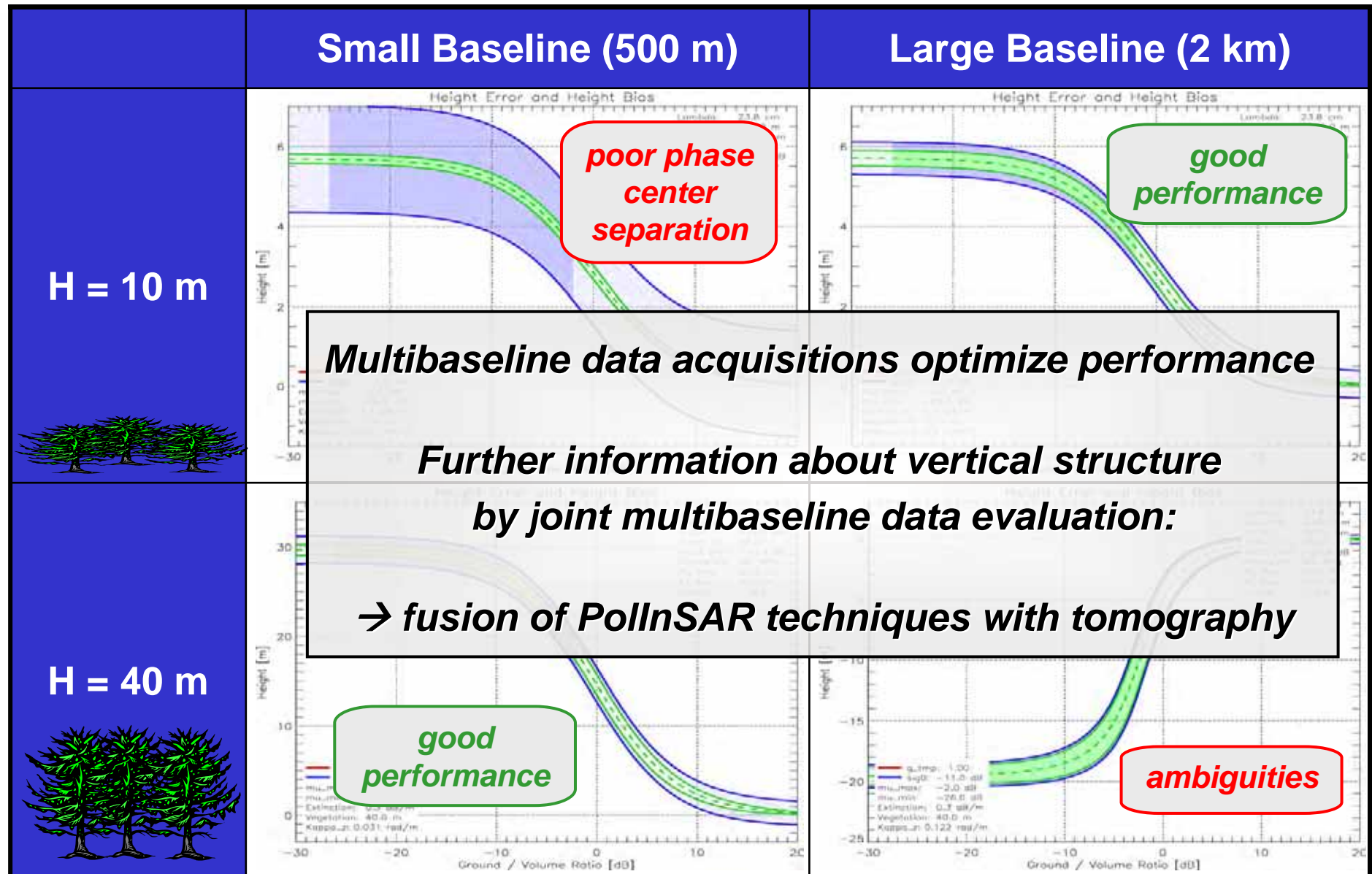
→ Good Inversion Accuracy



Deutsches Zentrum  
für Luft- und Raumfahrt e.V.  
in der Helmholtz-Gemeinschaft



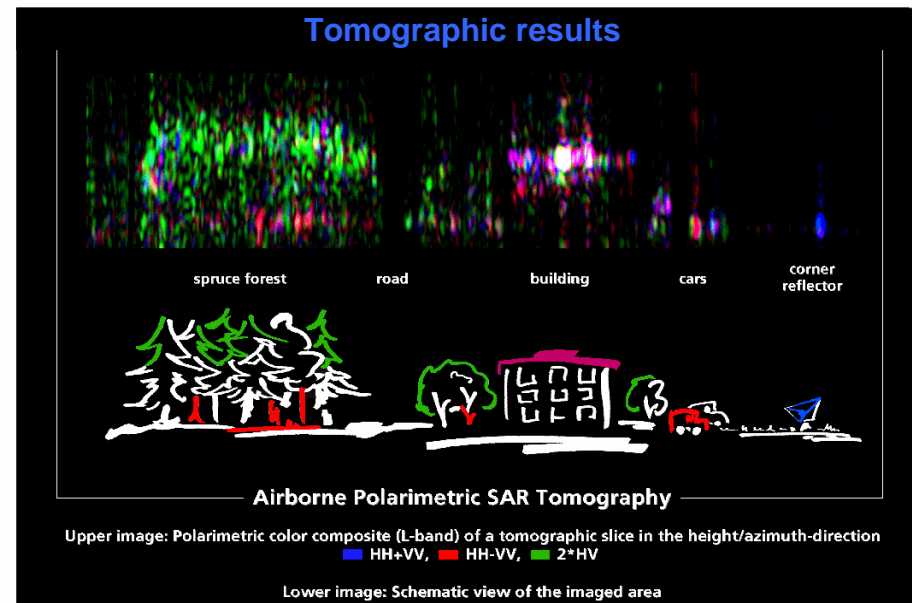
# PollnSAR L-Band Performance Example



# Tomography with Micro-Satellite Array

**Basic Idea:** Cluster of receiver satellites to form an additional aperture in elevation

- allows real three-dimensional imaging, i.e. a geometric resolution in height direction
- avoids the intrinsic height ambiguity in conventional SAR imaging
- accurate modeling and retrieval of vegetation parameters
- not affected by layover or foreshortening effects
- cross-track distance between the satellites defines the height ambiguity value for tomographic processing
- total tomographic baseline defines the height resolution



(from A. Reigber & A. Moreira, 1999)

# Tomography with Semi-Active Micro-Satellite Array

## Fundamental relations:

- height resolution:

$$\Delta h \approx \frac{\lambda \cdot r_0 \cdot \sin \theta_{inc}}{L \cdot \cos \theta_{look}}$$

- required sampling:

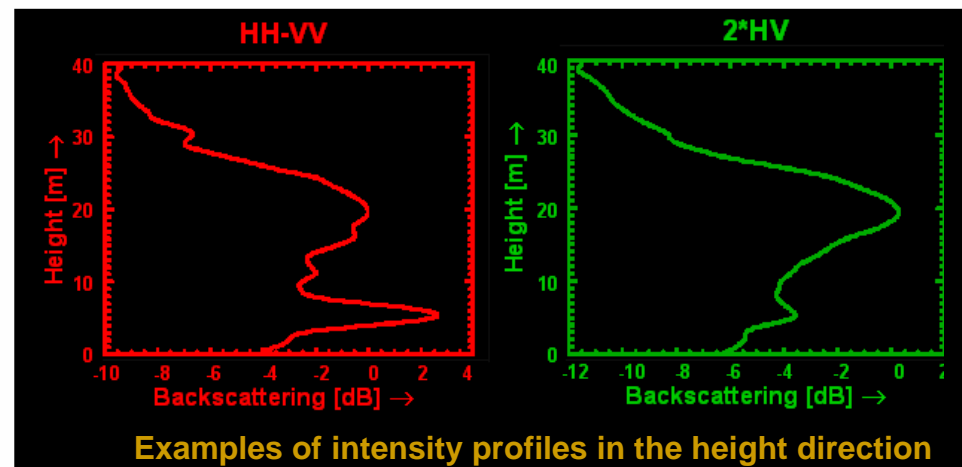
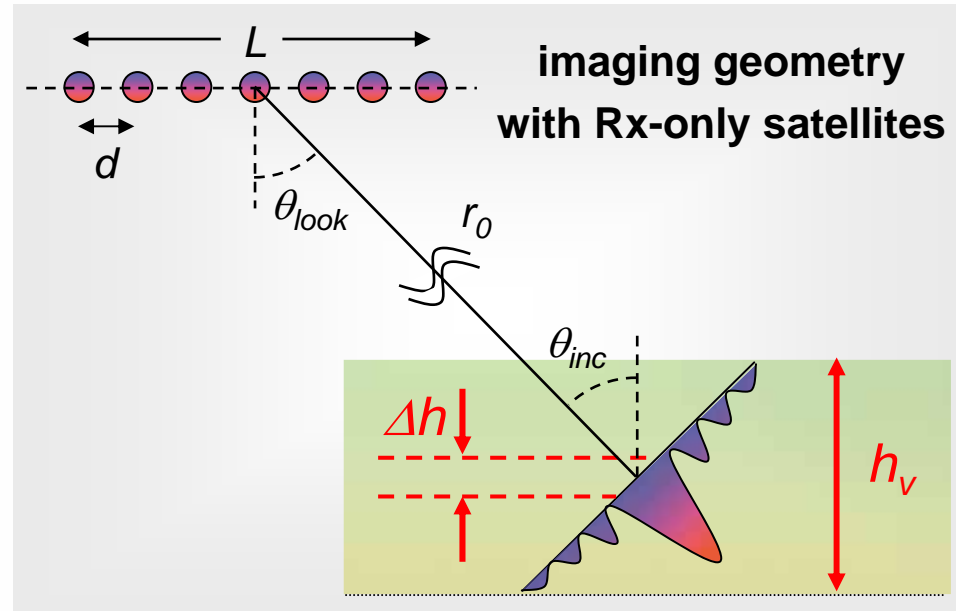
$$d < \frac{\lambda \cdot r_0 \cdot \sin \theta_{inc}}{h_v \cdot \cos \theta_{look}}$$

## Example:

$\lambda$	0.23 m
$r_0$	700 km
$\theta_{inc}$	30°
$\theta_{look}$	30°
$L$	20 km
$d$	2 km

$$\Delta h = 3 \text{ m}$$

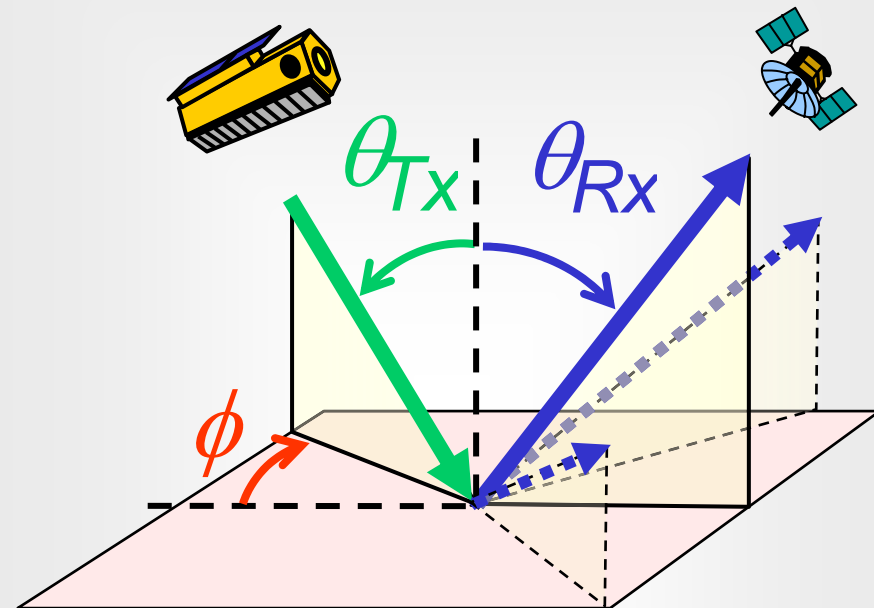
$$h_v < 30 \text{ m}$$



# Multistatic SAR Imaging

- improved detection, segmentation, and classification in SAR images
- separation of different scattering mechanisms (e.g. coherent from non-coherent components)
- radargrammetry and multi-shadow evaluations
- speckle reduction without resolution degradation
- multibaseline coherence analyses
- acquisition of bi- and multistatic Doppler spectra (e.g. multiple ocean wave spectra)
- downward looking receivers (fusion with other sensors)

## Extended Observation Space



*Multiple*

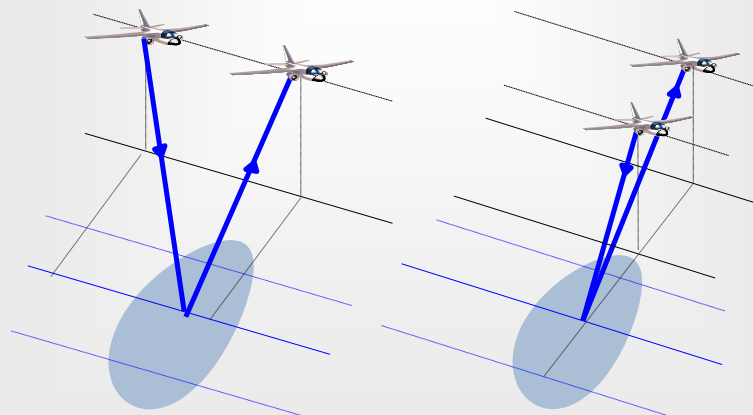
- Scattering angles  $\theta_{Rx}$
- Out-of-plane angles  $\phi$
- Doppler frequencies

+ polarimetry (e.g.  $\sigma_{HV} \neq \sigma_{VH}$ )



# Multistatic Scattering Coefficients: Example

*Bistatic Airborne Radar Experiment:  
(February 2003, Nimes, France)*

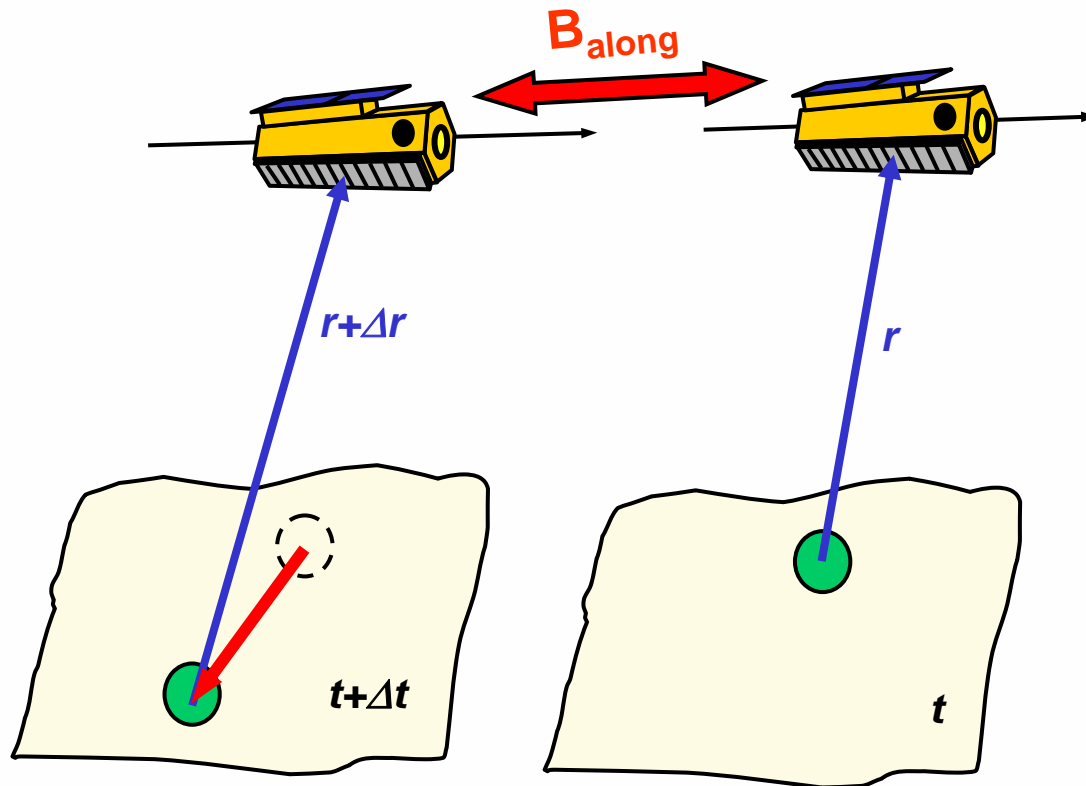


*Color composite of  
three bistatic images:*

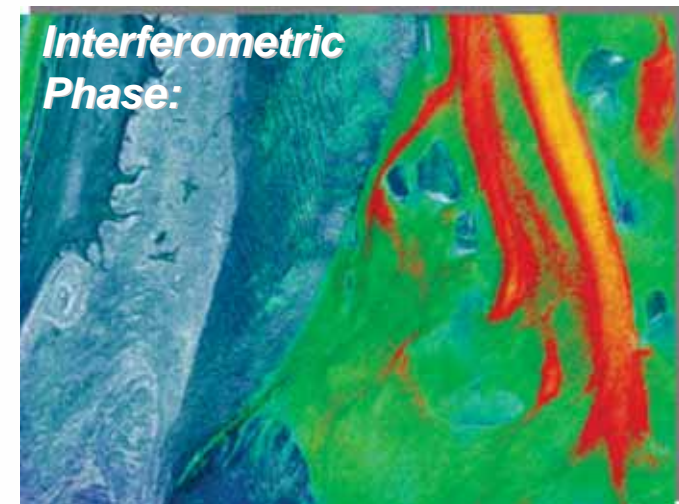
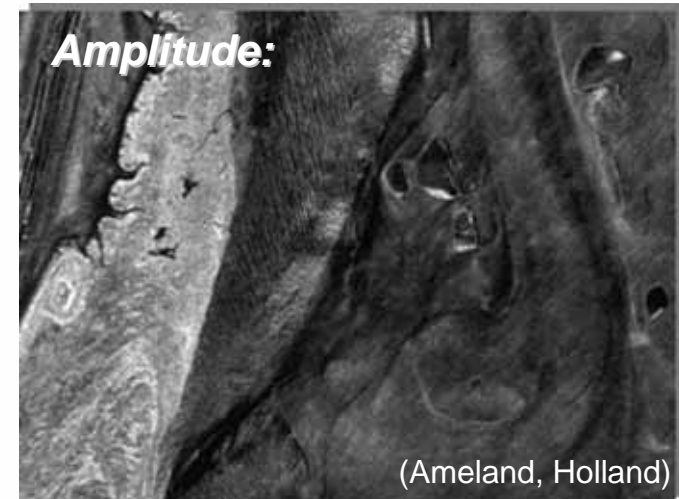


Deutsches Zentrum  
für Luft- und Raumfahrt e.V.  
in der Helmholtz-Gemeinschaft

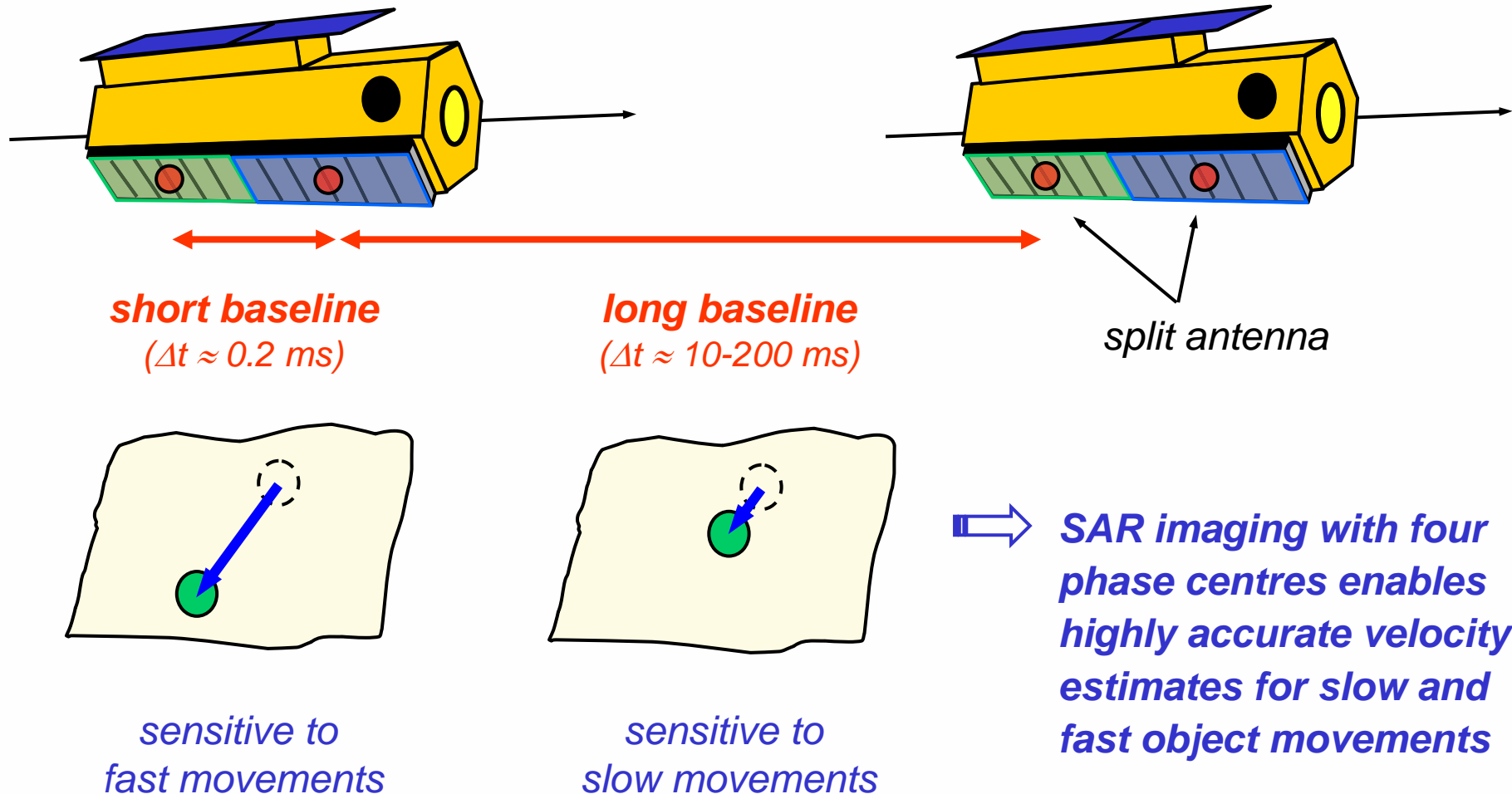
# Along-Track Interferometry



*highly accurate measurements of the radial displacement between two radar observations separated by a short time lag*



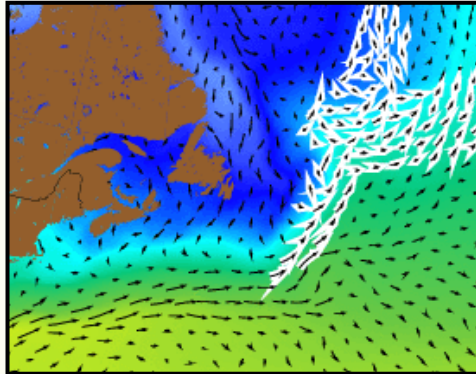
# SAR Imaging with Four Phase Centres





# Applications of Along-Track Interferometry

***Ocean  
Currents***



***Coastal  
Surveillance***



***Ice Drift &  
Ice Flow***



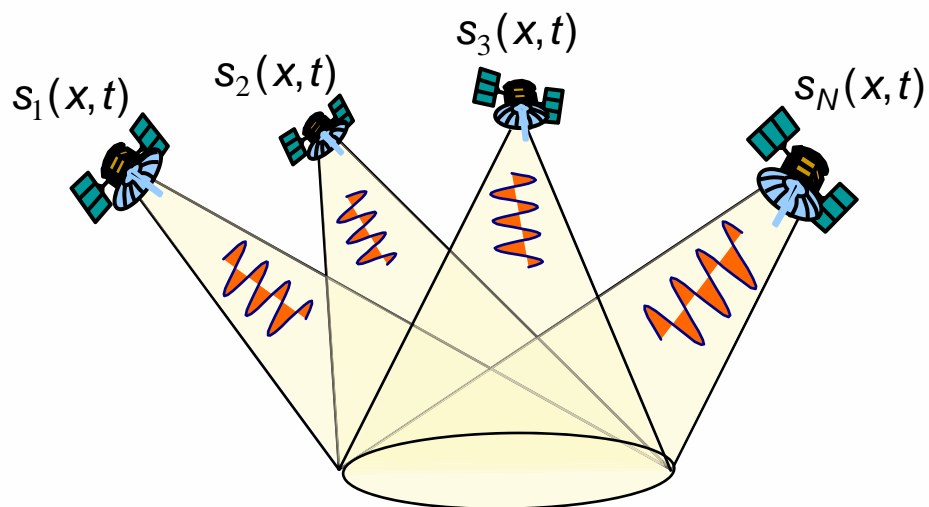
***Traffic  
Monitoring***



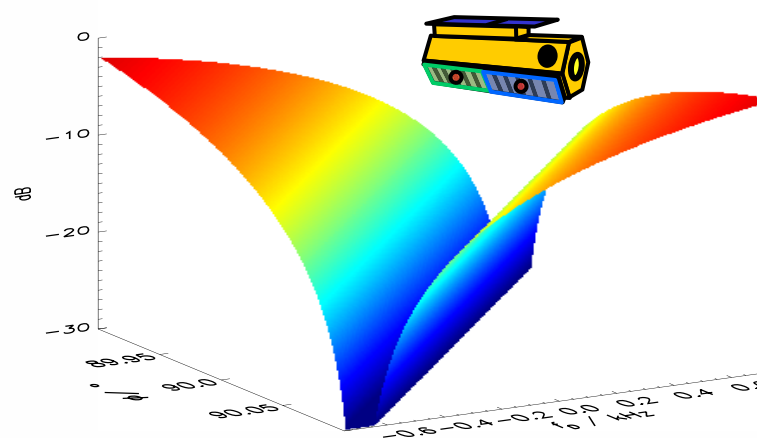
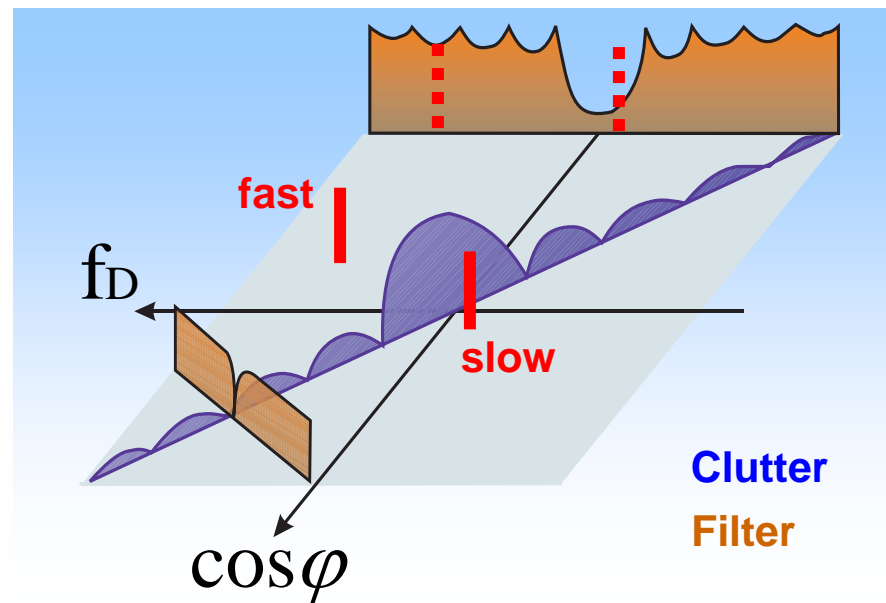
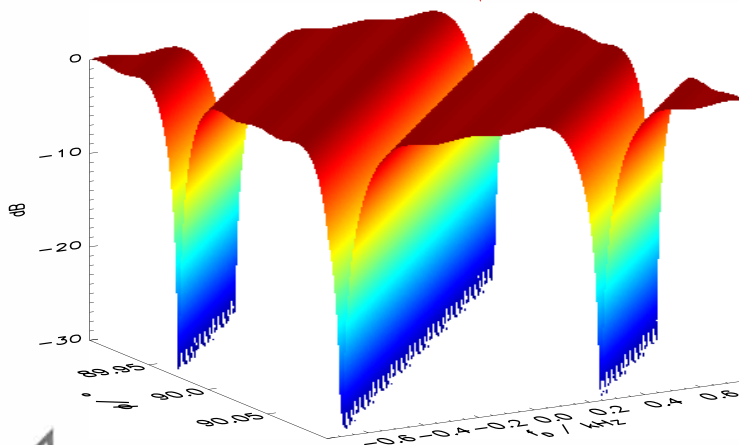




# Multi-Baseline ATI and GMTI



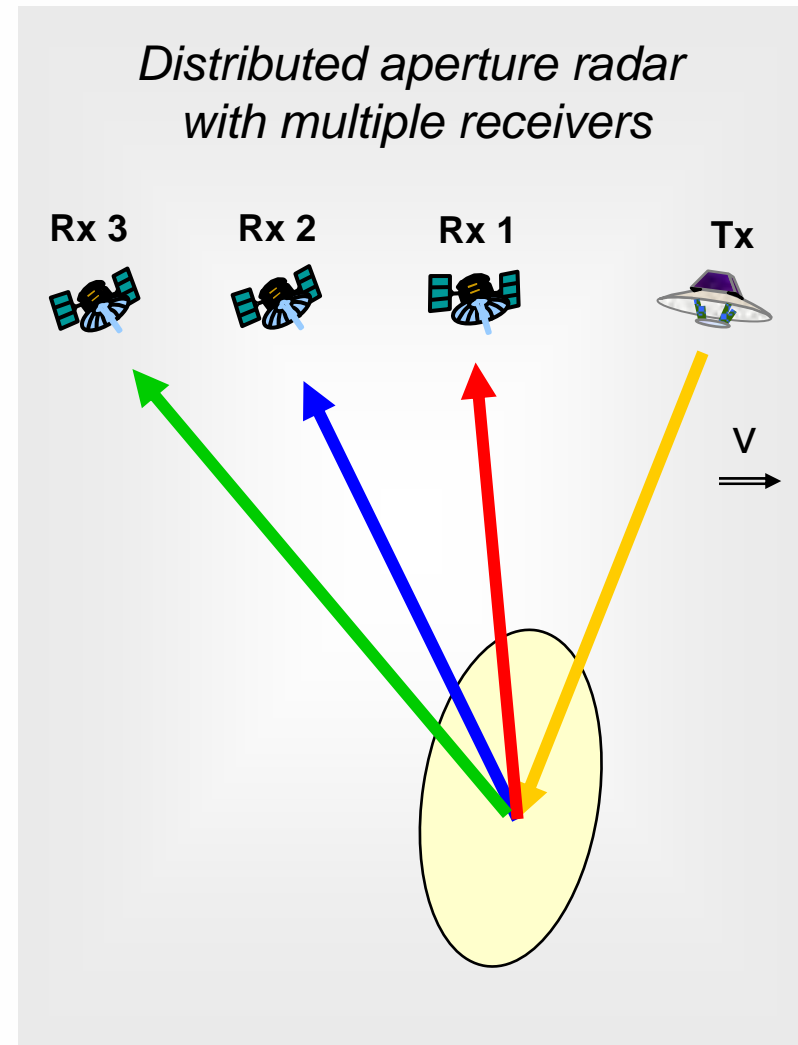
*steep notch enables good clutter suppression*



Deutsches Zentrum  
für Luft- und Raumfahrt e.V.  
in der Helmholtz-Gemeinschaft

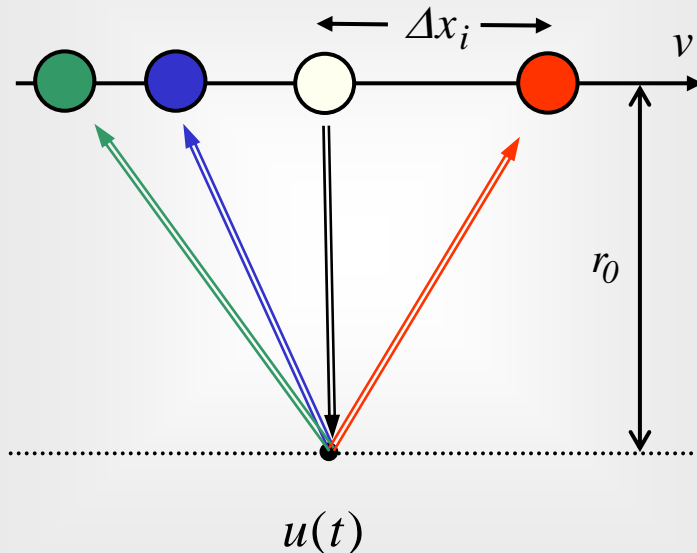
# Ambiguity Reduction and Wide Swath Imaging

- single transmitter illuminates **wide image swath**
- multiple receivers record scattered signal
- N receivers allow **reduction of PRF** by a factor of 1/N **without** raising **azimuth ambiguities**:
  - increase of swath width by factor N at **full azimuth resolution** (as opposed to ScanSAR)
  - variability in optimum receiver displacement:
$$x_i - x_1 \approx \frac{2 \cdot v}{PRF} \left( \frac{i-1}{N} + k_i \right) \quad i \in \{1, 2, \dots, N-1\}, k_i \in \mathbb{Z}$$
  - reconstruction possible for other displacements
  - performance can be optimized by PRF adaptation
  - requires stable oscillators or RF synchronization and accurate estimation of relative displacement
- major application: **high resolution** imaging of a **wide image swath** with **small antennas** (e.g. distributed L-Band SAR with multiple microsatellites)

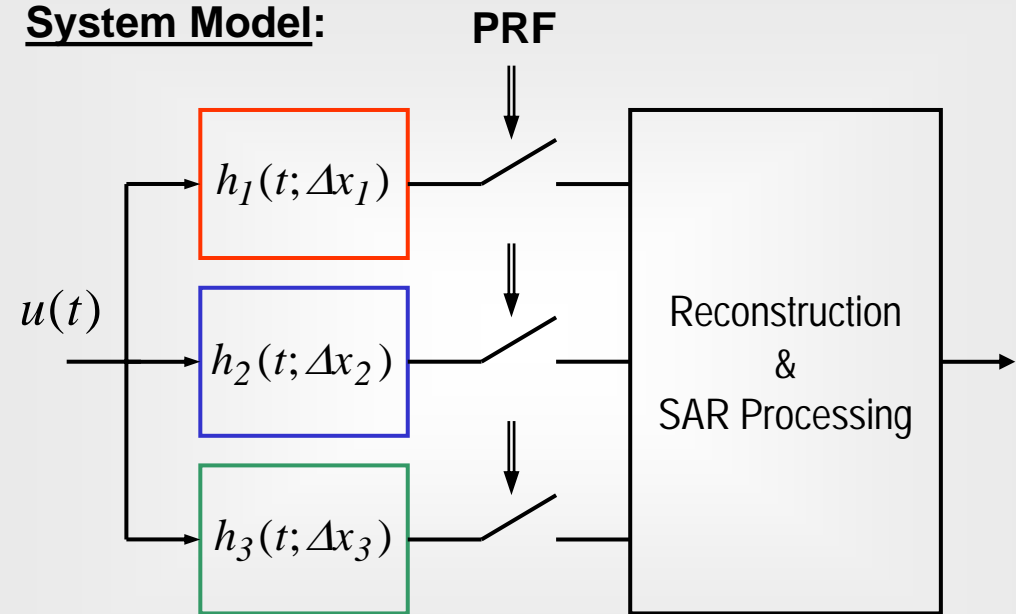


# Multi-Channel Model for Sparse Array

## Multiple Aperture Array:



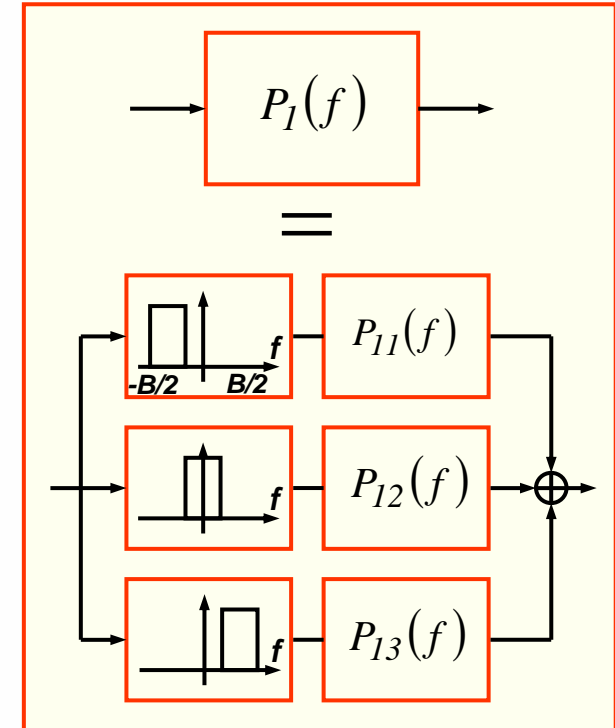
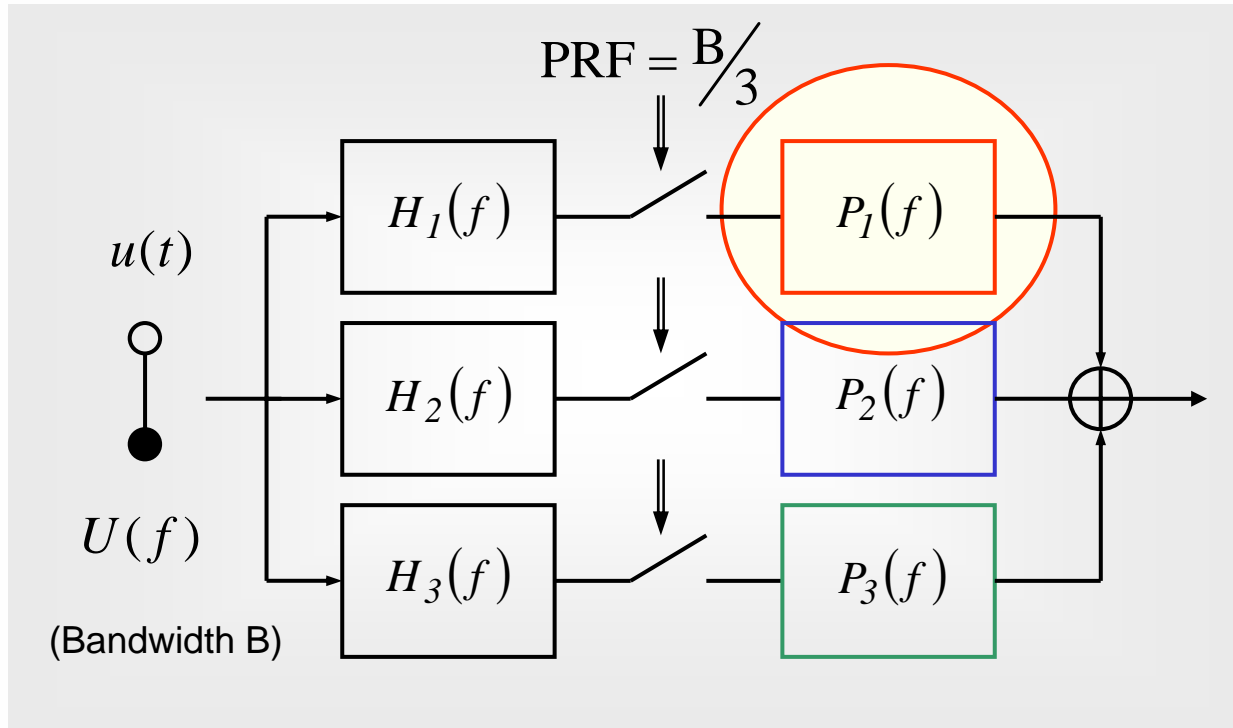
## System Model:



## Bistatic Azimuth Impulse Response:

$$h_i(t; \Delta x_i) = A_{Tx}(vt) \cdot A_{Rx,i}(vt - \Delta x_i) \cdot \exp \left[ -j \frac{2\pi}{\lambda} \left( \sqrt{r_0^2 + (vt)^2} + \sqrt{r_0^2 + (vt - \Delta x_i)^2} \right) \right]$$

# Coherent Reconstruction

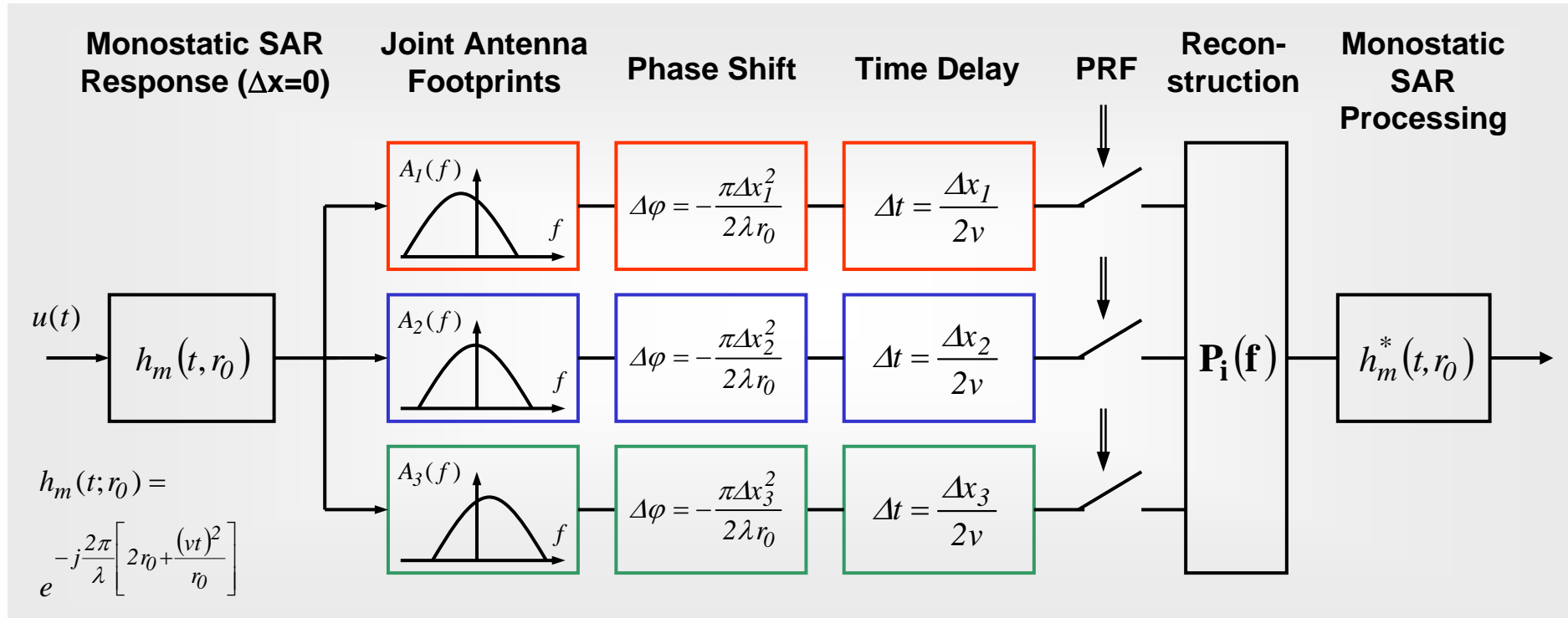


$$\mathbf{A}(\mathbf{f}) = \begin{bmatrix} H_1(f) & H_2(f) & H_3(f) \\ H_1(f + B/3) & H_2(f + B/3) & H_3(f + B/3) \\ H_1(f + 2B/3) & H_2(f + 2B/3) & H_3(f + 2B/3) \end{bmatrix} \rightarrow \mathbf{A}^{-1}(\mathbf{f}) = \begin{bmatrix} P_{11}(f) & P_{12}(f + B/3) & P_{13}(f + 2B/3) \\ P_{21}(f) & P_{22}(f + B/3) & P_{23}(f + 2B/3) \\ P_{31}(f) & P_{32}(f + B/3) & P_{33}(f + 2B/3) \end{bmatrix}$$

(cf. A. Papoulis, 1977 & J.L. Brown, 1981)



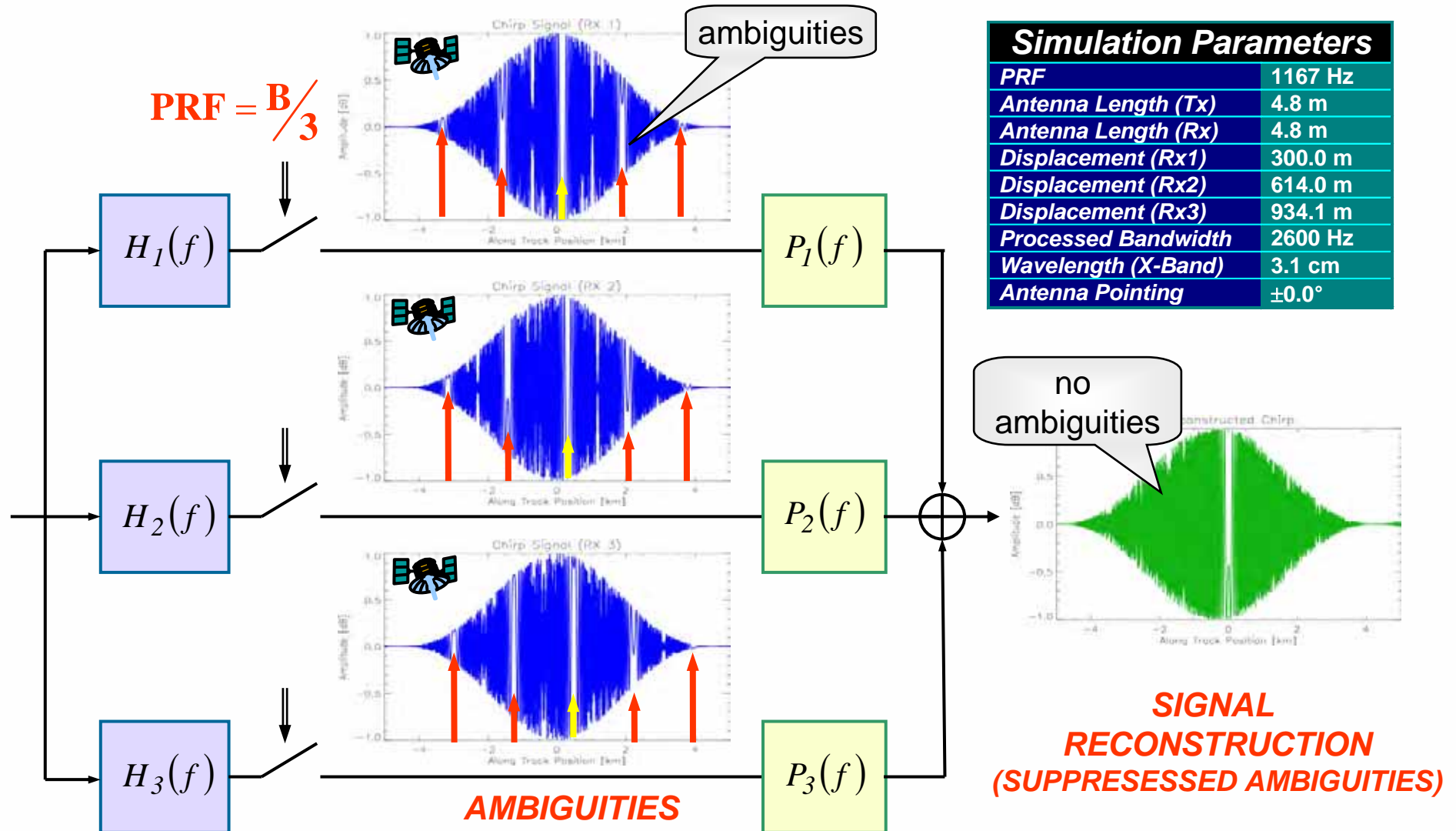
# Model with Quadratic Phase Approximation



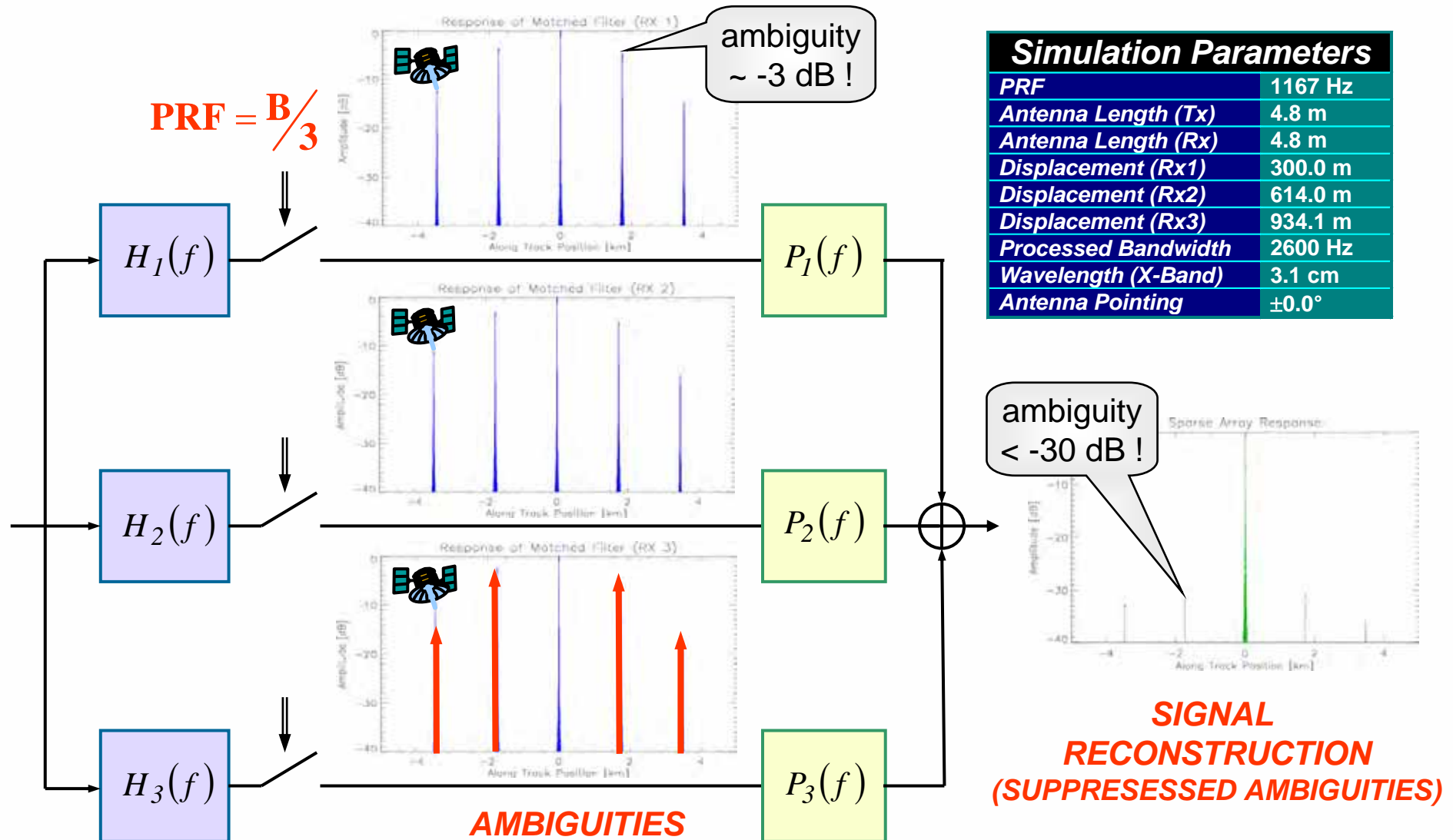
## Bistatic Azimuth Impulse Response:

$$h_i(t; \Delta x_i) \cong A_{Tx}(vt) \cdot A_{Rx,i}(vt - \Delta x_i) \cdot \exp\left[-j\frac{4\pi}{\lambda}r_0\left(1 + \frac{\Delta x_i^2}{8r_0^2}\right)\right] \cdot \exp\left[-j\frac{2\pi}{\lambda}\frac{\left(vt - \frac{\Delta x_i}{2}\right)^2}{r_0}\right]$$

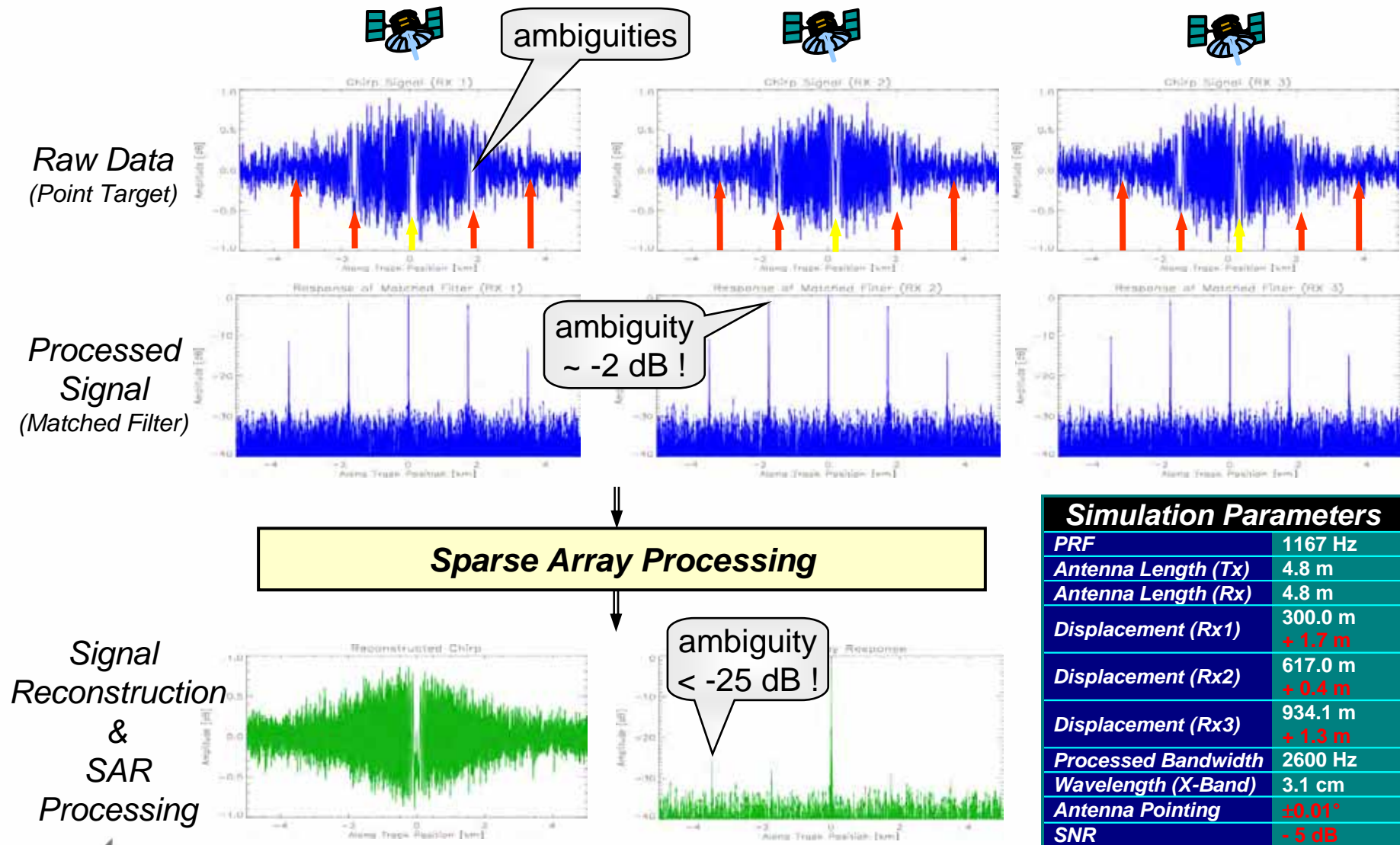
# Sparse Array Reconstruction: Raw Data



# Sparse Array Reconstruction: Focused Data



# Sparse Array Reconstruction: Nonuniform Distance + Noise

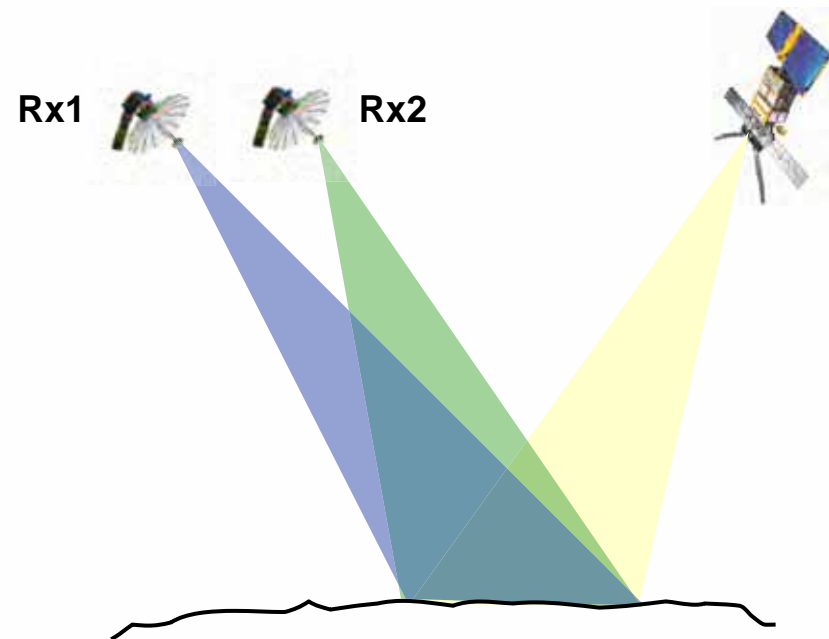
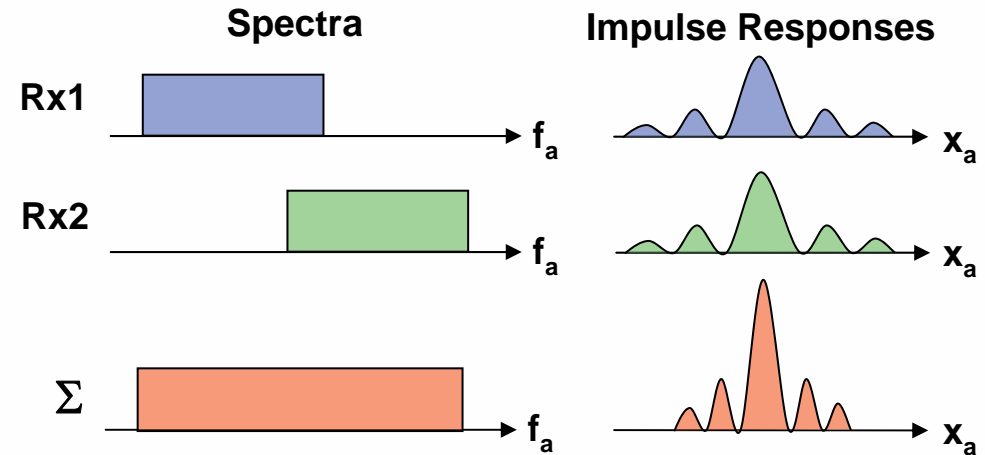




# Superresolution with Multistatic Satellite Arrays

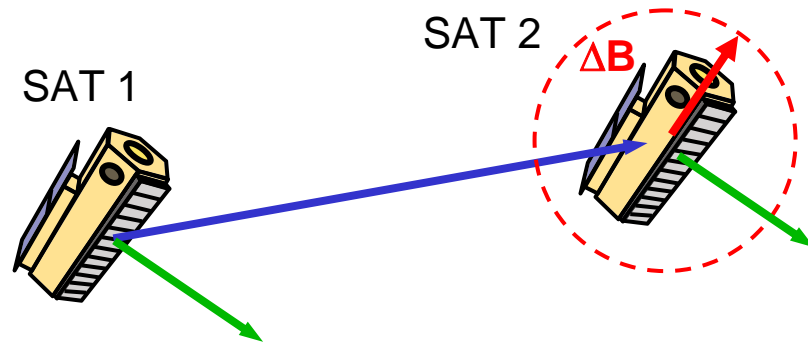
## Increased geometric resolution of SAR images by:

- along-track displacement of receiving satellites :
  - ⇒ different Doppler centroids
  - ⇒ super-resolution in azimuth by coherent combination of shifted Doppler spectra
- across-track displacement of receiving satellites:
  - ⇒ different incident angles
  - ⇒ super-resolution in range by coherent combination of images with different ground range spectra

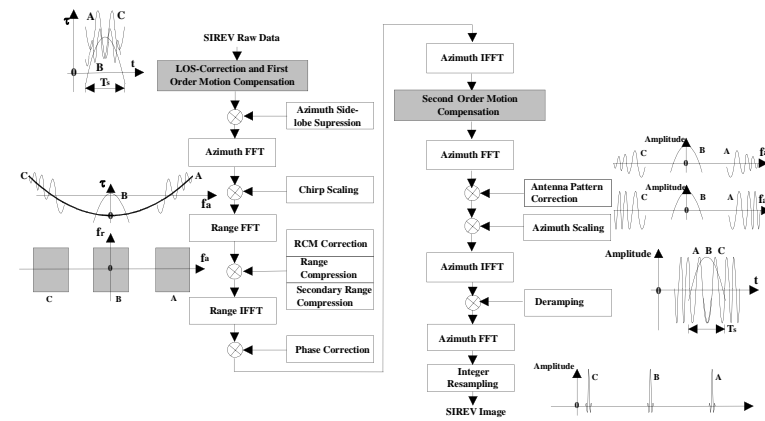


# Challenges in Bistatic and Multistatic SAR Imaging

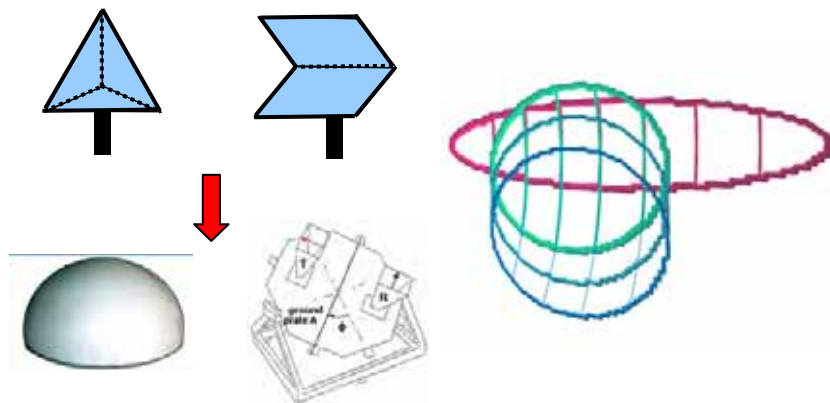
## Orbit Control & Relative Position Sensing



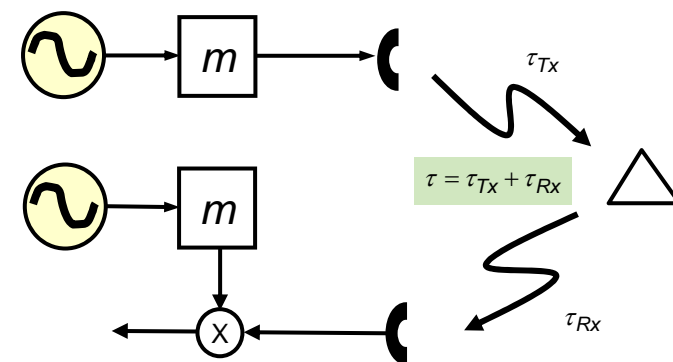
## Bi- and Multistatic SAR Processing



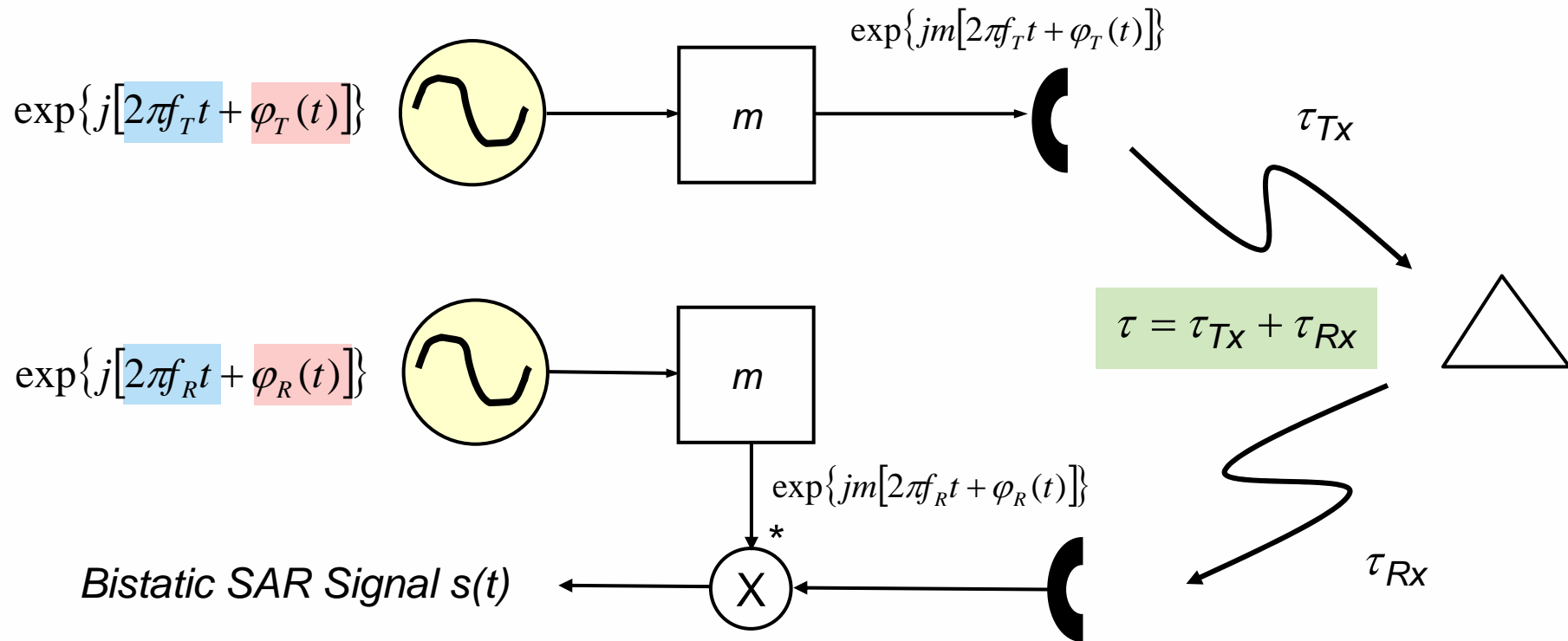
## Bistatic Calibration & Antenna Co-Pointing



## Time and Phase Synchronisation



# System Model for Bistatic Radar



$$s(t) = \exp\{jm[-2\pi f_T \tau + 2\pi(f_T - f_R)t + \varphi_T(t - \tau) - \varphi_R(t)]\}$$

**Azimuth  
Modulation**

**Frequency  
Offset  $\Delta f$**

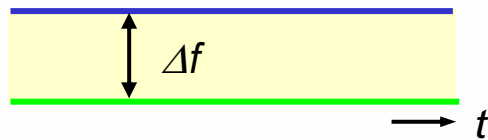
**Phase  
Noise  $\varphi(t)$**



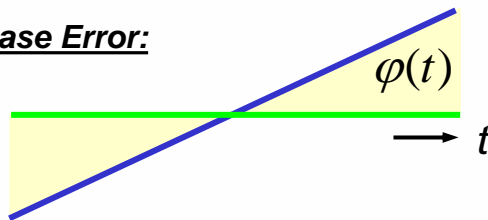
# Degradation of Azimuth Impulse Response

## Constant Frequency Offset

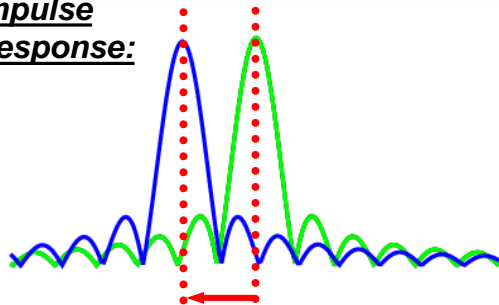
Frequency Offset:



Phase Error:

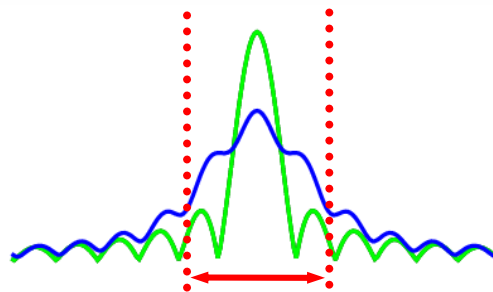
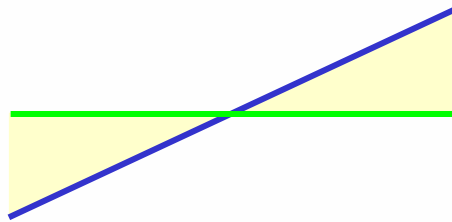


Impulse Response:



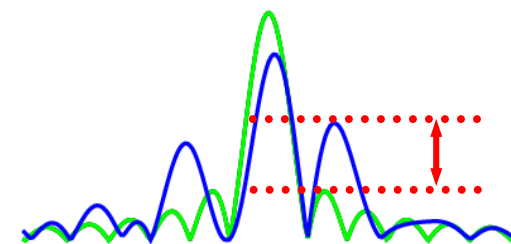
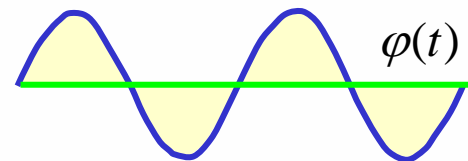
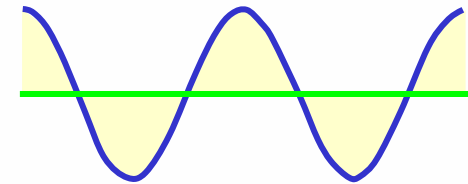
**Linear Displacement**

## Linear Frequency Drift



**Mainlobe Dispersion**

## Higher-Order Phase Error



**Increase of Sidelobes**

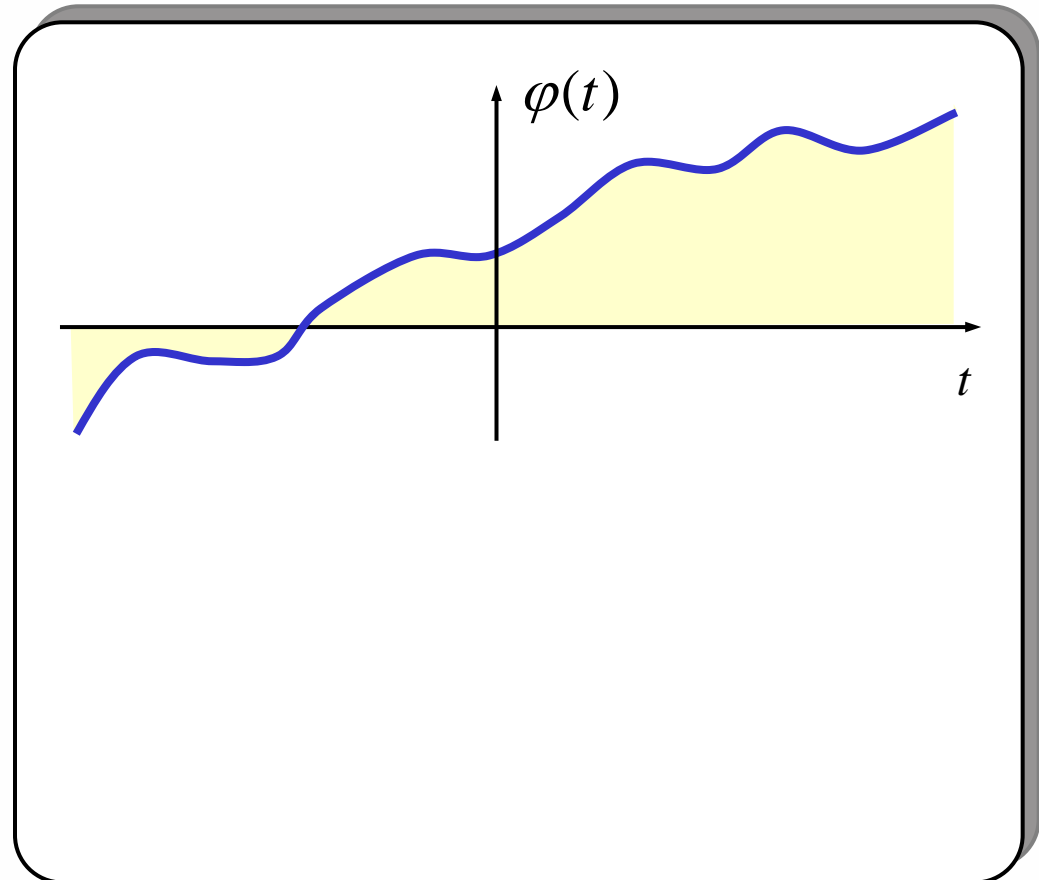






## Model of Oscillator Phase Errors

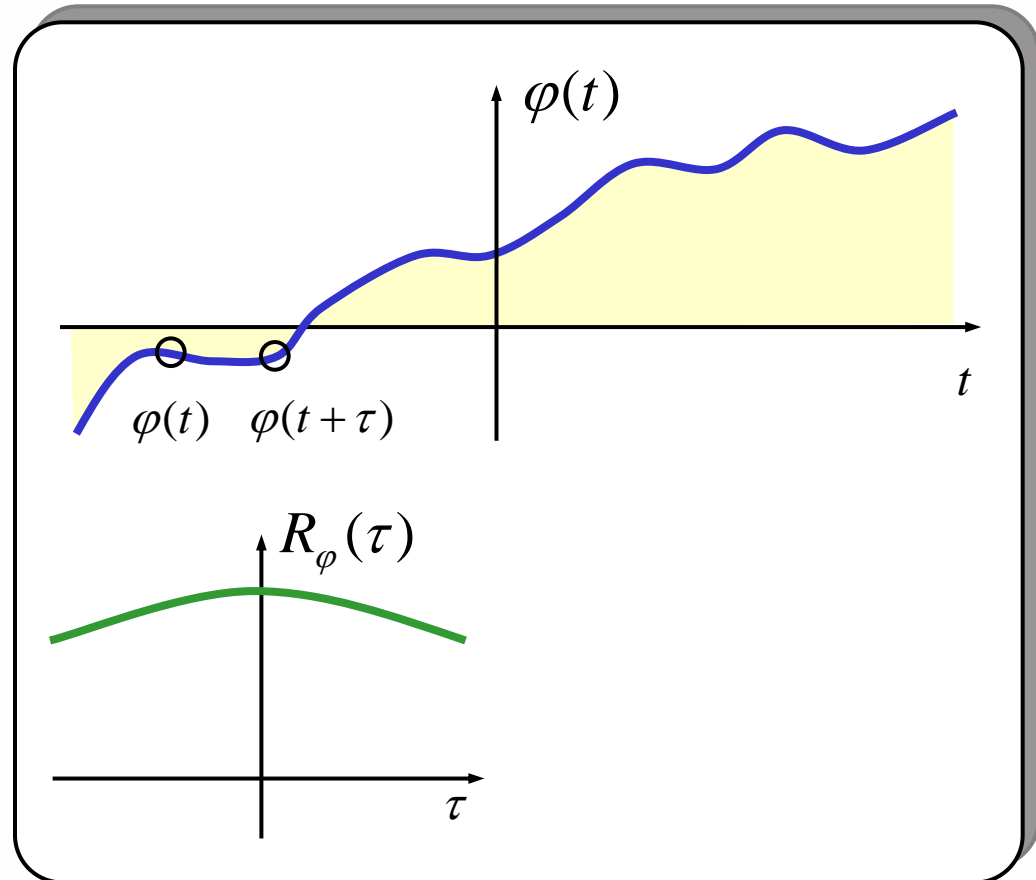
- $\varphi(t)$  is modelled by a stochastic process



## Model of Oscillator Phase Errors

- $\varphi(t)$  is modelled by a stochastic process with acf

$$R_{\varphi}(\tau) = \langle \varphi(t) \cdot \varphi(t+\tau) \rangle$$



## Model of Oscillator Phase Errors

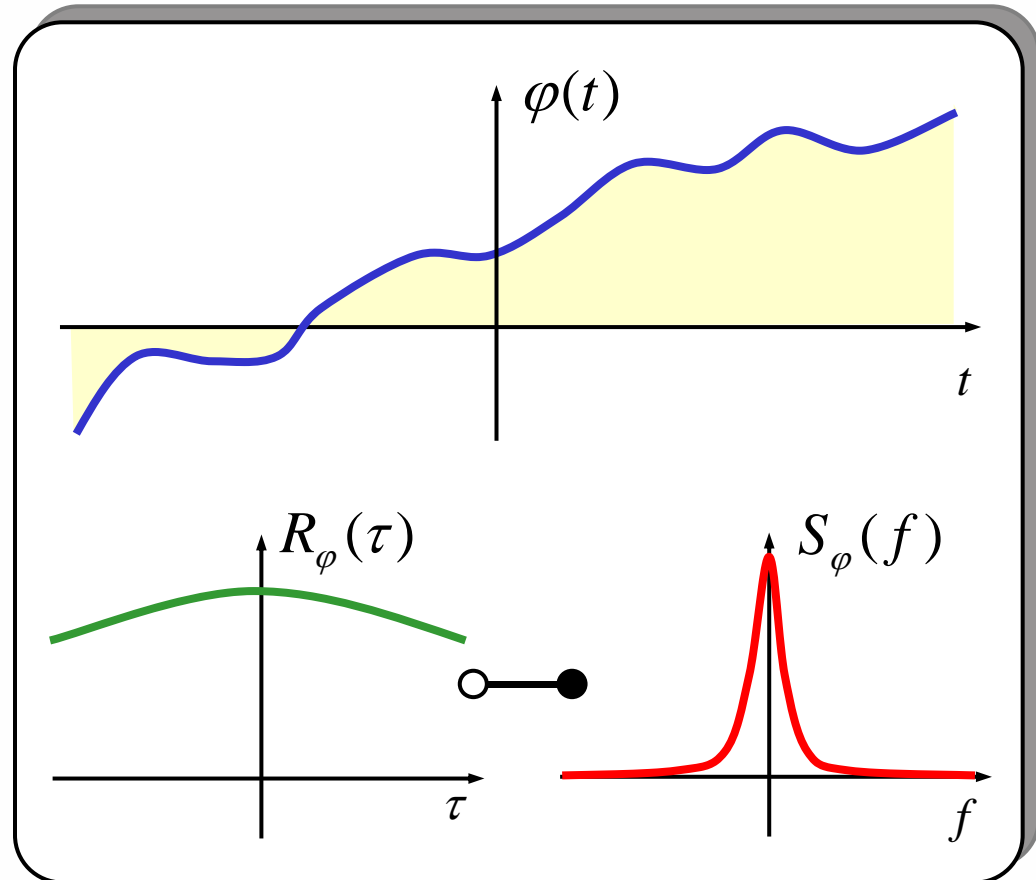
- $\varphi(t)$  is modelled by a stochastic process with acf

$$R_{\varphi}(\tau) = \langle \varphi(t) \cdot \varphi(t+\tau) \rangle$$

- The phase spectrum

$$S_{\varphi}(f) \longleftrightarrow R_{\varphi}(\tau)$$

describes phase fluctuations  
per Hertz bandwidth at Fourier  
frequency  $f$



# Model of Oscillator Phase Errors

- $\varphi(t)$  is modelled by a stochastic process with acf

$$R_{\varphi}(\tau) = \langle \varphi(t) \cdot \varphi(t+\tau) \rangle$$

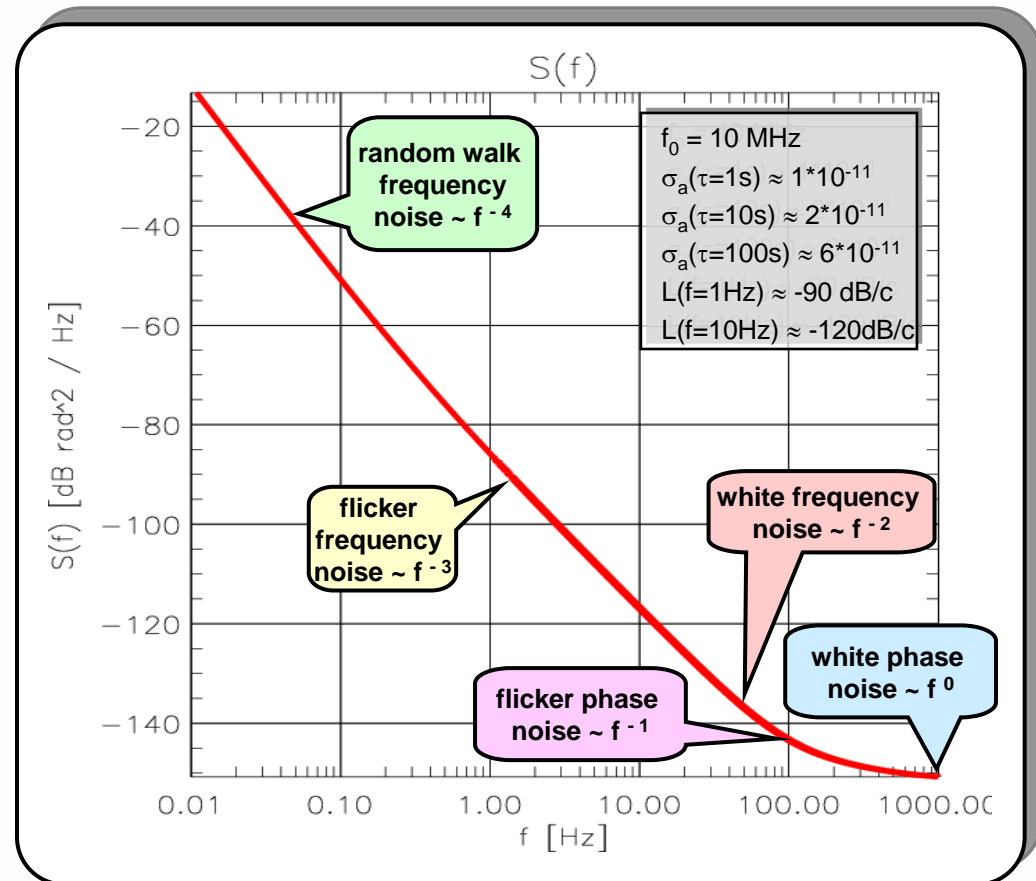
- The phase spectrum

$$S_{\varphi}(f) \longleftrightarrow R_{\varphi}(\tau)$$

describes phase fluctuations per Hertz bandwidth at Fourier frequency  $f$

- $S_{\varphi}(f)$  is often approximated by:

$$S_{\varphi}(f) = \underbrace{a \cdot f^{-4}}_{\text{random walk frequency noise}} + \underbrace{b \cdot f^{-3}}_{\text{frequency flicker noise}} + \underbrace{c \cdot f^{-2}}_{\text{white frequency noise}} + \underbrace{d \cdot f^{-1}}_{\text{flicker phase noise}} + \underbrace{e \cdot f^0}_{\text{white phase noise}}$$

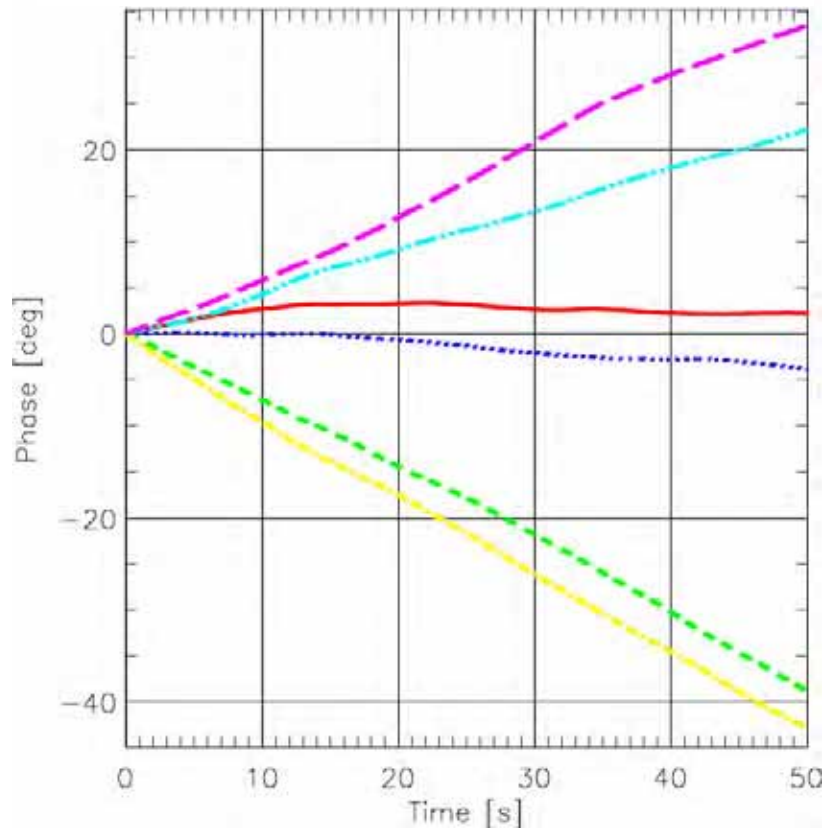




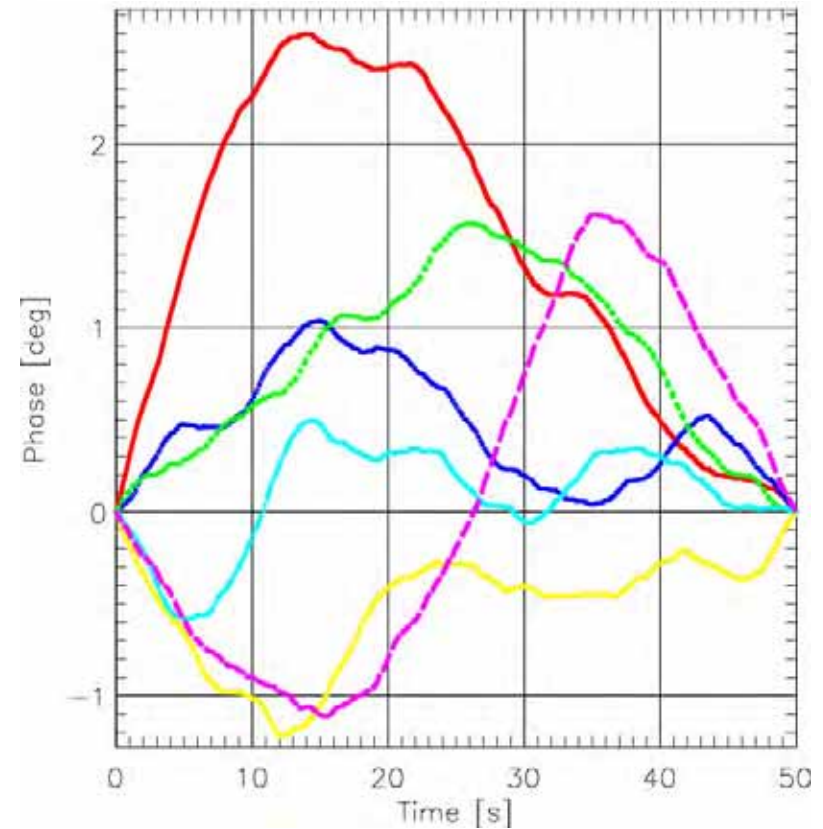


## Typical Realisations from $R_\varphi(\tau) \longleftrightarrow S_\varphi(f)$

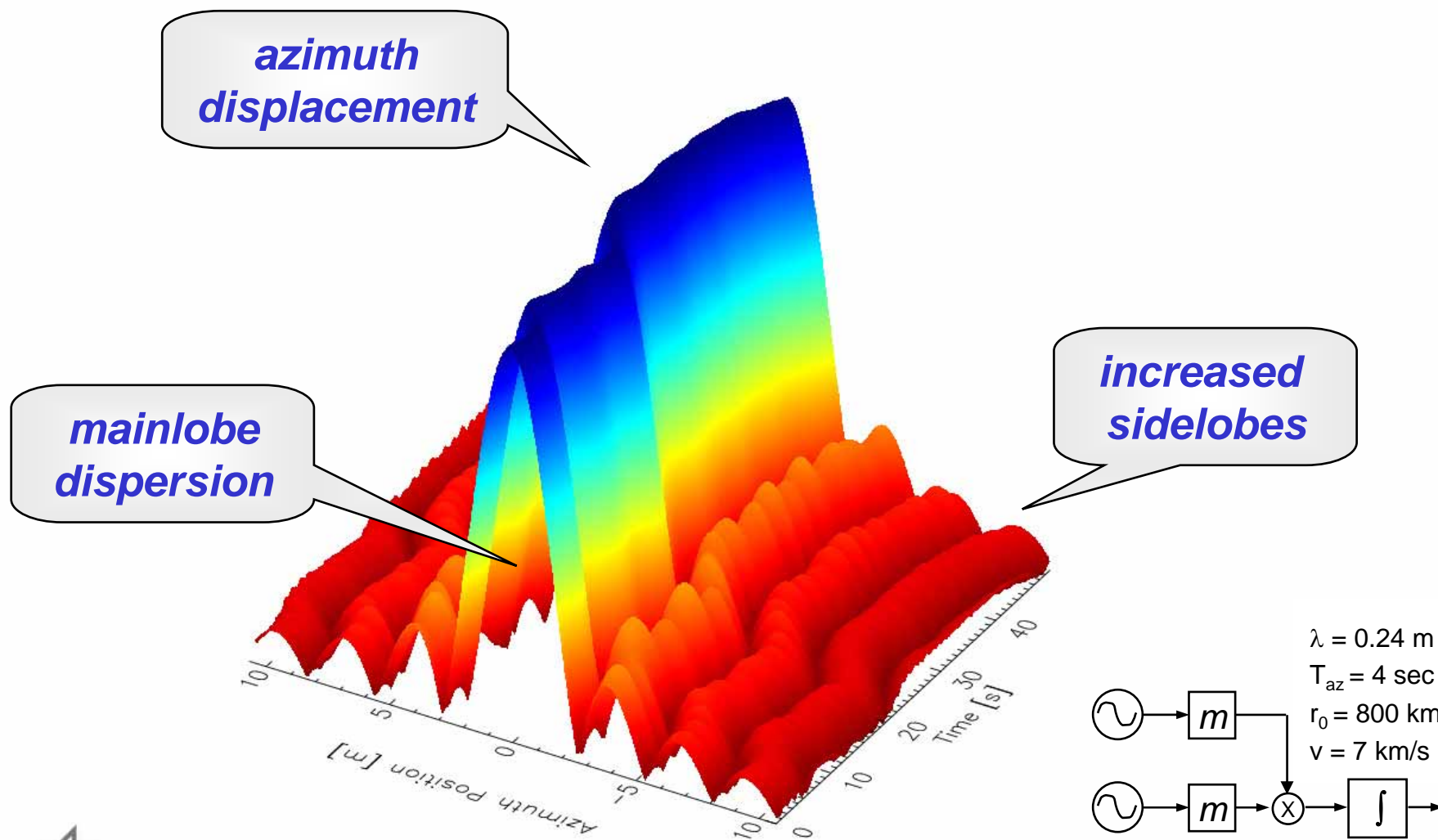
Realisations with  $\varphi(t=0) = 0^\circ$



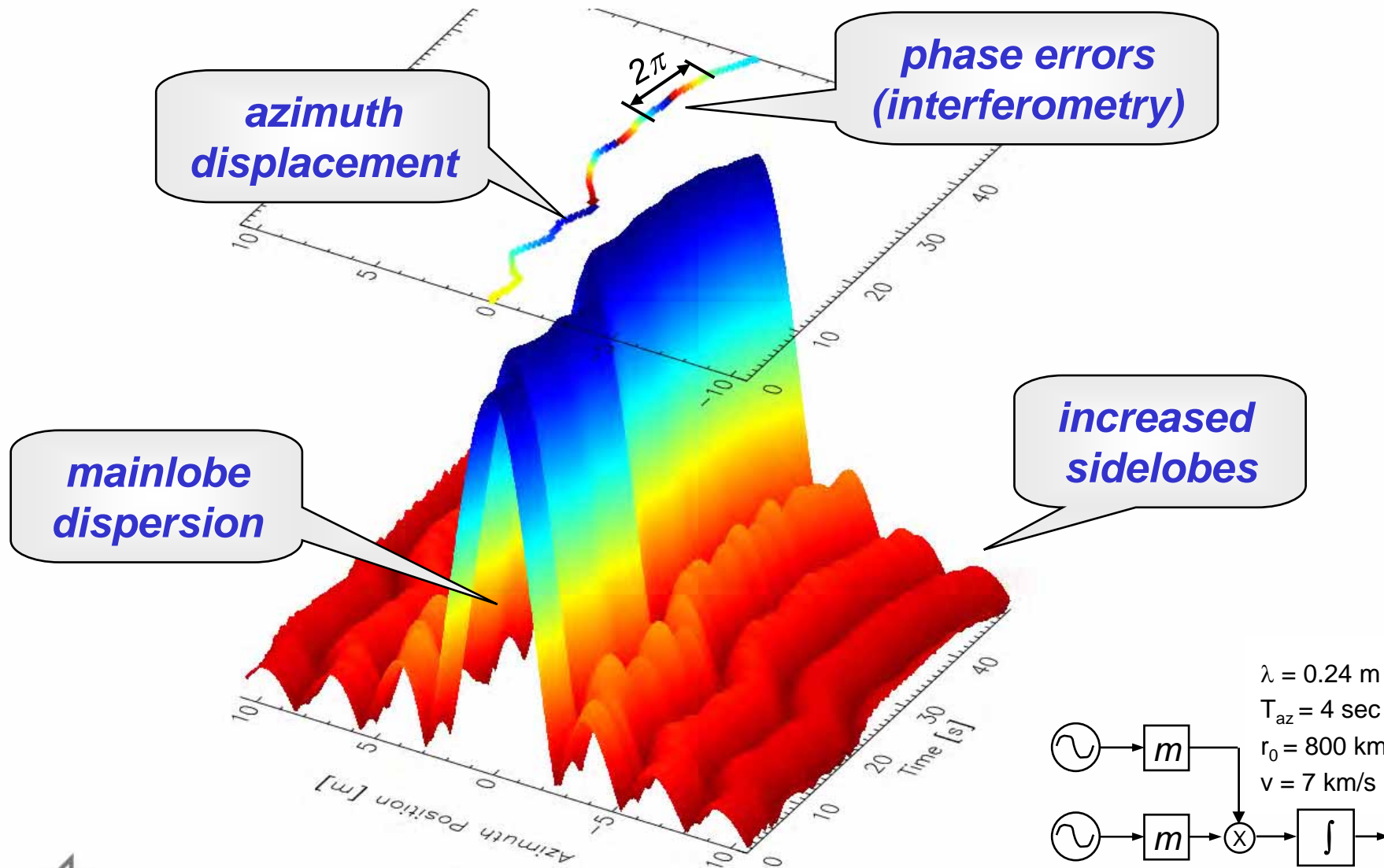
Constant phase ramp removed



# Bistatic SAR Response: Simulation Example



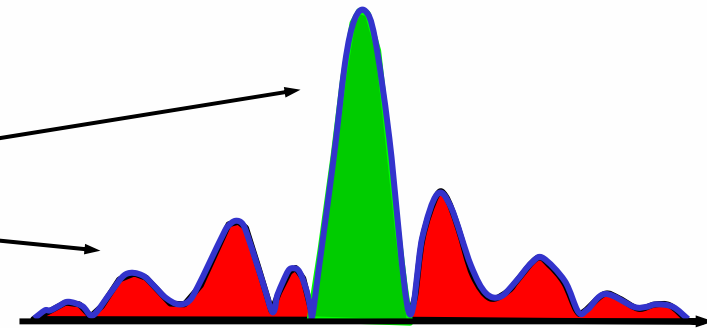
# Bistatic SAR Response: Simulation Example



# Integrated Side-Lobe Ratio (ISLR)

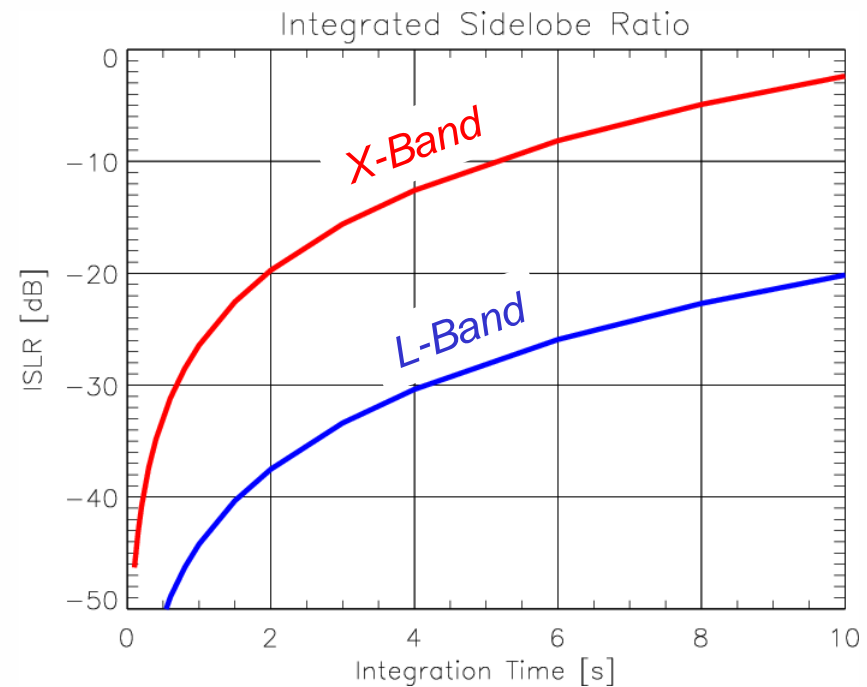
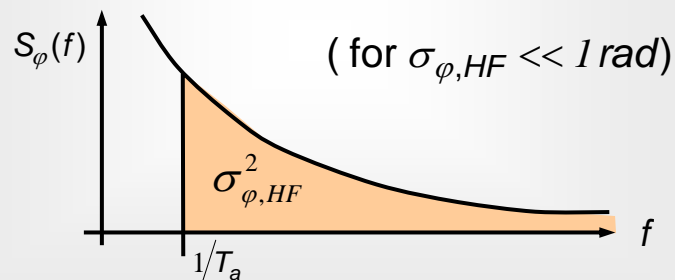
*Definition:*

$$ISLR = \frac{\text{Signal Energy in Mainlobe}}{\text{Integrated Signal Energy in Sidelobes}}$$



*Derivation from Phase Spectrum  $S_\varphi(f)$ :*

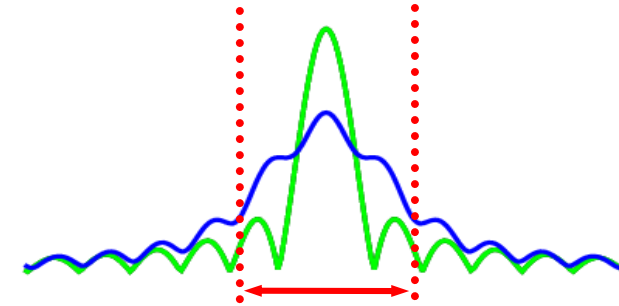
$$ISLR \approx 2 \cdot \left( \frac{f_0}{f_{osc}} \right)^2 \cdot \int_{1/T_a}^{\infty} S_\varphi(f) \cdot df$$





# Mainlobe Dispersion

- mainlobe dispersion is mainly due to quadratic phase errors
- typical requirement:  $\varphi_Q < 45^\circ$  (~ 3 % dispersion)
- quadratic phase errors may be derived from second derivative of phase

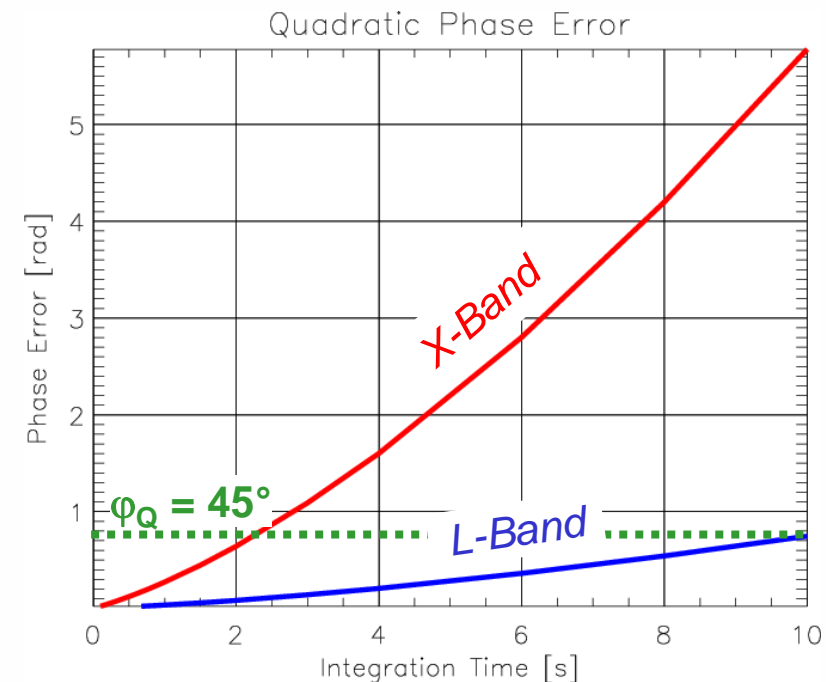


## Quadratic Phase Errors:

$$\sigma_Q^2 = 2 \cdot \left( \frac{f_0}{f_{osc}} \right)^2 \cdot \frac{(\pi T_a)^4}{4} \cdot \frac{1}{T_a} \int_0^{1/T_a} f^4 \cdot S_\varphi(f) \cdot df$$

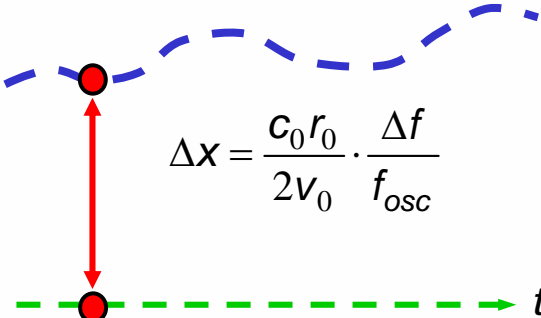
## Mainlobe Dispersion:

$$\frac{\Delta az_{bistat}}{\Delta az_{ideal}} \approx \sqrt{1 + \left( \frac{\sigma_Q}{\pi} \right)^2} \quad \text{for } \varphi_Q < \pi$$



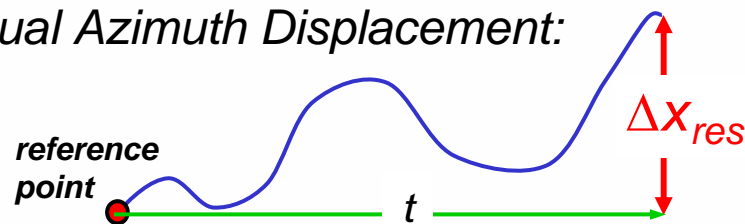
# Azimuth Displacement

- small frequency offset  $\rightarrow$  large azimuth displacement
- possible solutions:
  - use of ground control points
  - simultaneous mono- and bistatic data acquisition
  - relative phase referencing between Tx and Rx

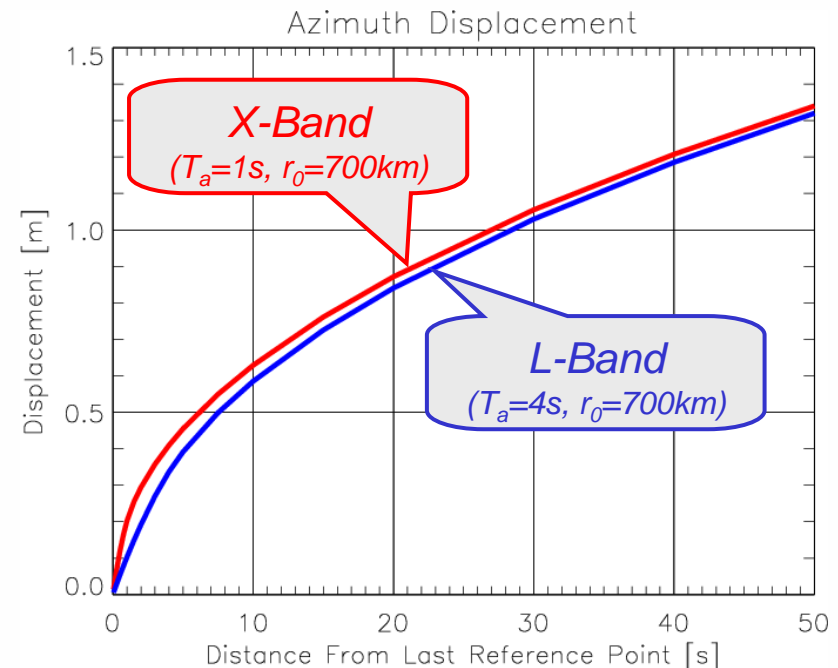


$$\Delta x = \frac{c_0 r_0}{2v_0} \cdot \frac{\Delta f}{f_{osc}}$$

*Residual Azimuth Displacement:*



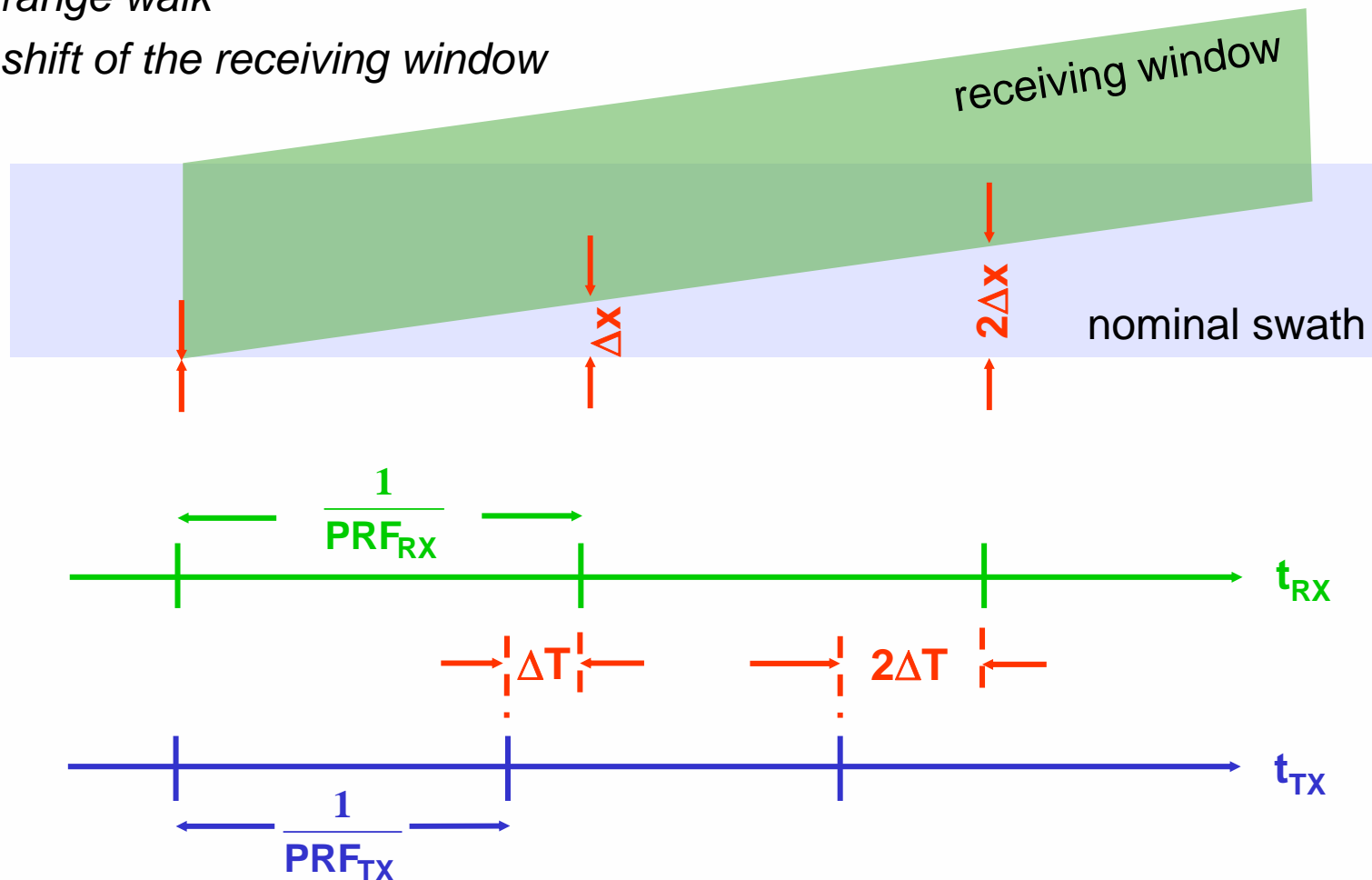
$$\sigma_{\Delta x_{res}}^2 = \left( \frac{c_0 r_0}{v_0} \right)^2 \cdot \int_0^\infty \frac{f^2}{f_{osc}^2} \cdot S_\phi(f) \cdot \left[ \frac{\sin(\pi T_a f)}{\pi T_a f} \right]^2 \cdot \left[ 1 - \left( \frac{\sin(2\pi f t)}{2 \sin(\pi f t)} \right)^2 \right] \cdot df$$



# Pulse Synchronisation

LO frequency offset will lead to slightly different PRFs:

- range walk
- shift of the receiving window

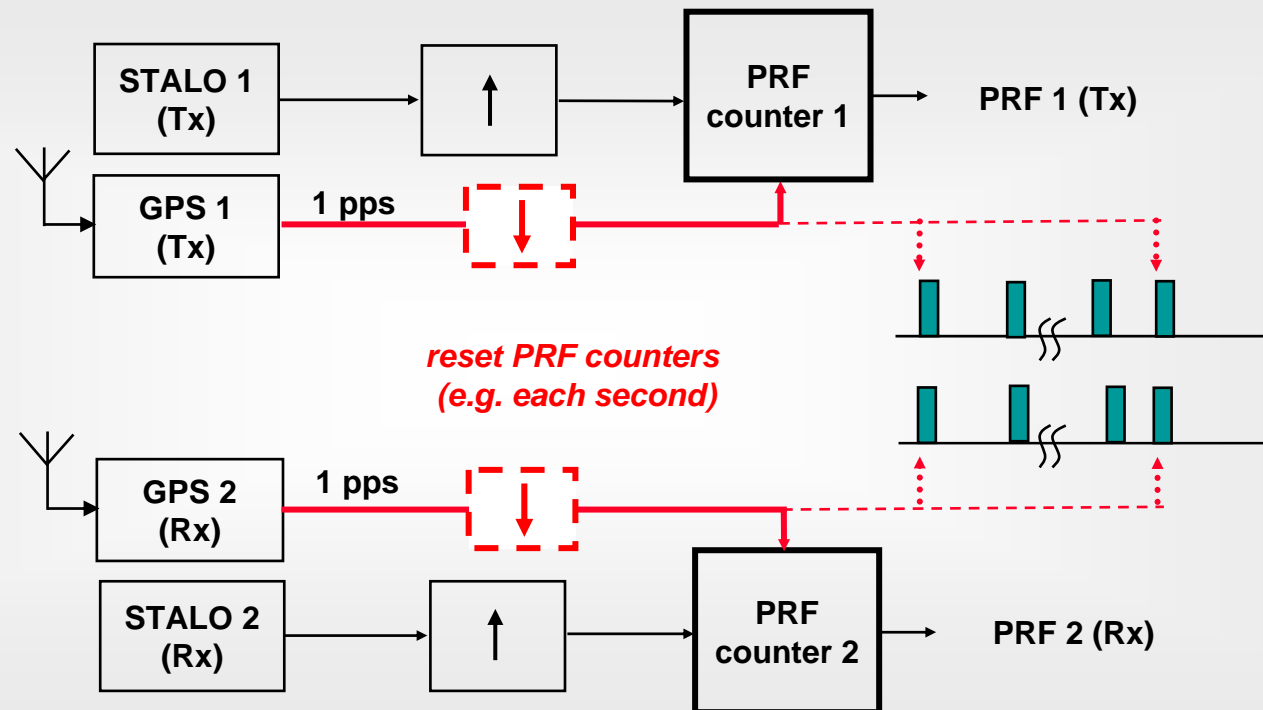


# Pulse Synchronisation

## Solutions to synchronization:

- *Pulse synchronization with link to common reference (e.g. GPS)*
- *Automatic detection of swath signal (e.g. by power analysis of received signal)*
- *Direct data link between transmitter and receiver*
- *Continuous recording*

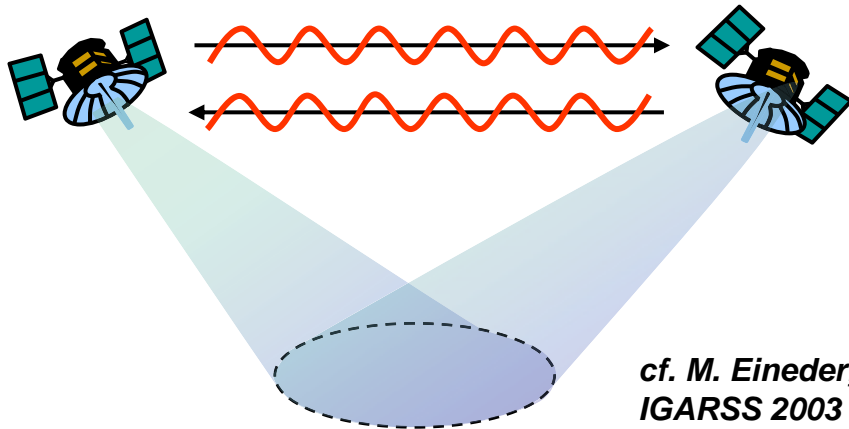
*Example for  
synchronization  
with GPS:*



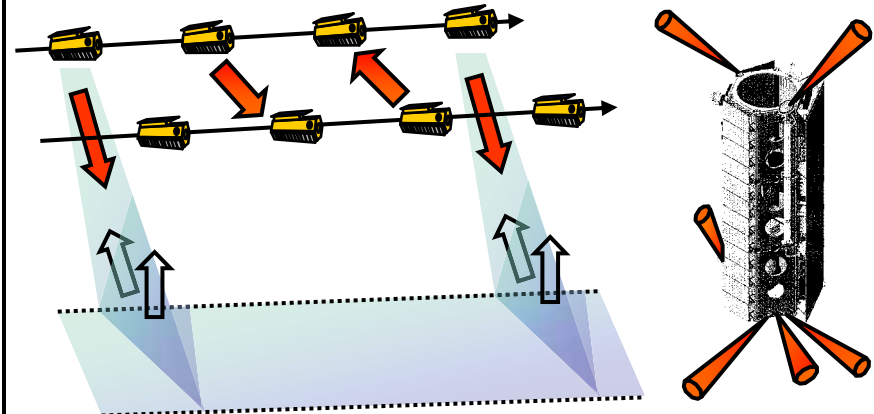


# Solutions for Phase Referencing

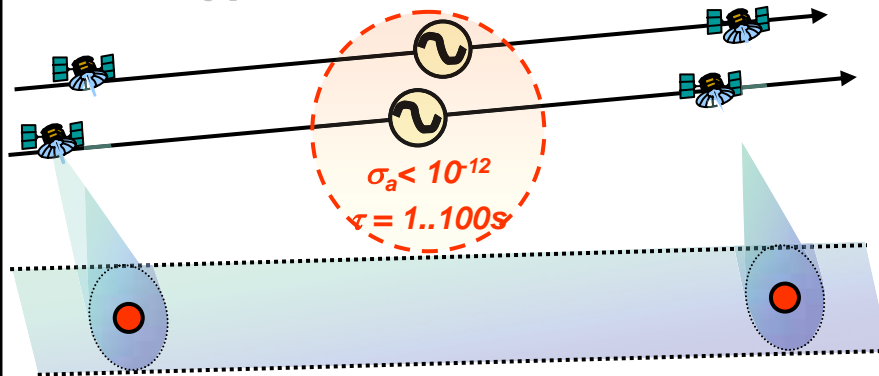
## Continuous USO Synchronisation



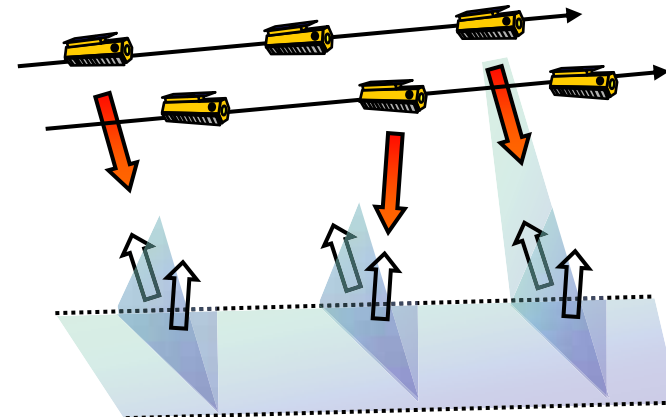
## Mutual Exchange of Radar Pulses



## Ground Control Points + 'Hyper-Stable' Oscillators



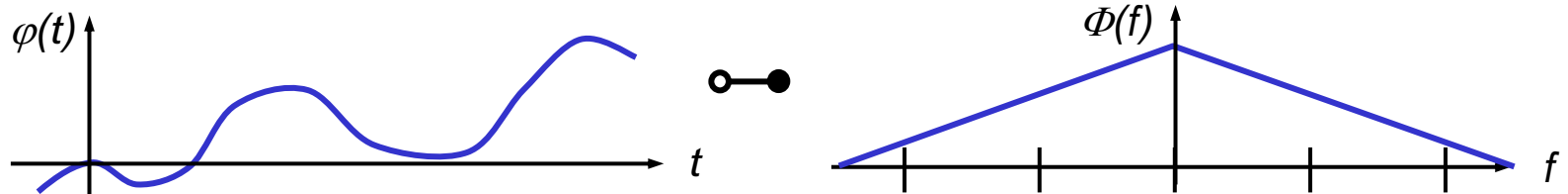
## Alternating Transmit Mode





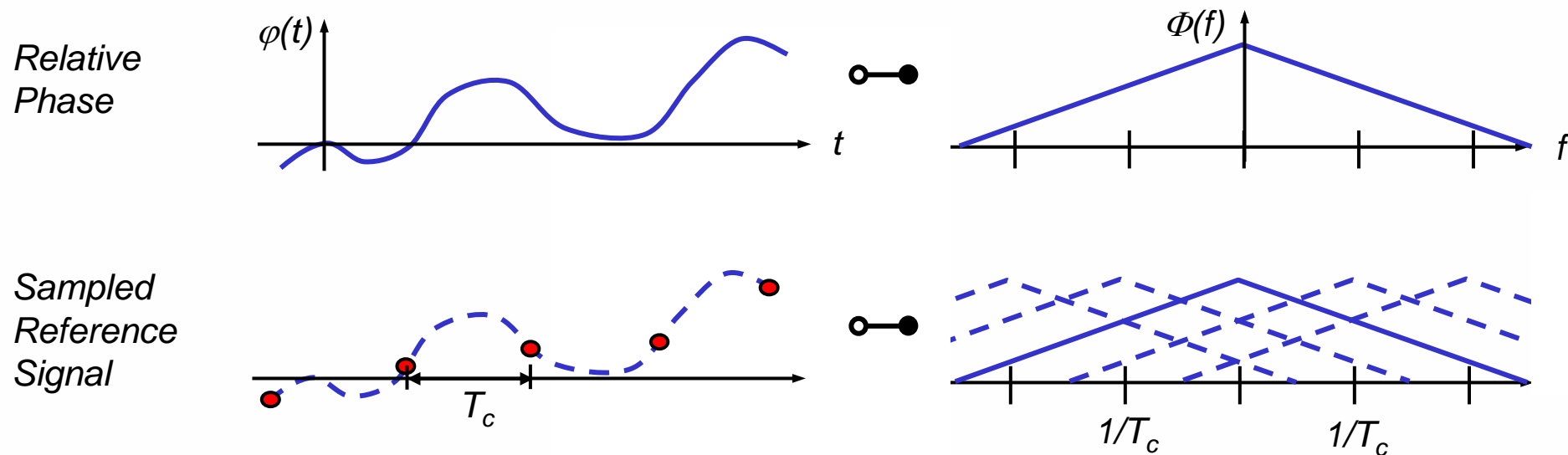
## Derivation of Required Sync Frequency

*Relative  
Phase*



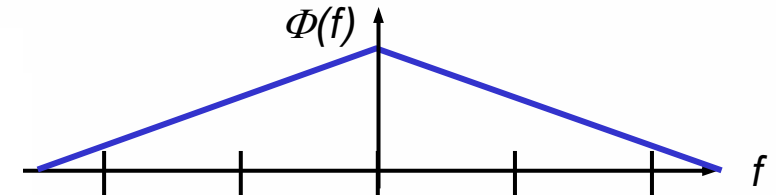
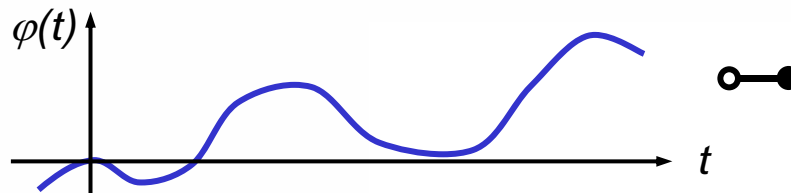


## Derivation of Required Sync Frequency

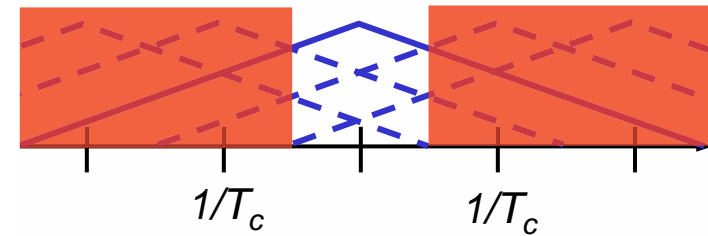
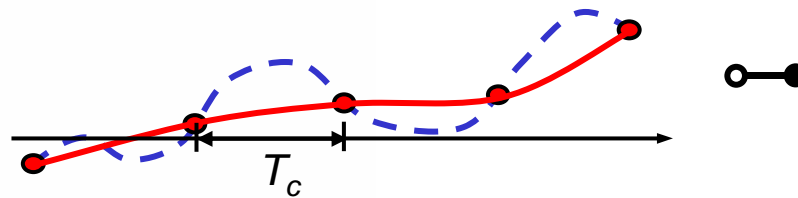


# Derivation of Required Sync Frequency

Relative  
Phase

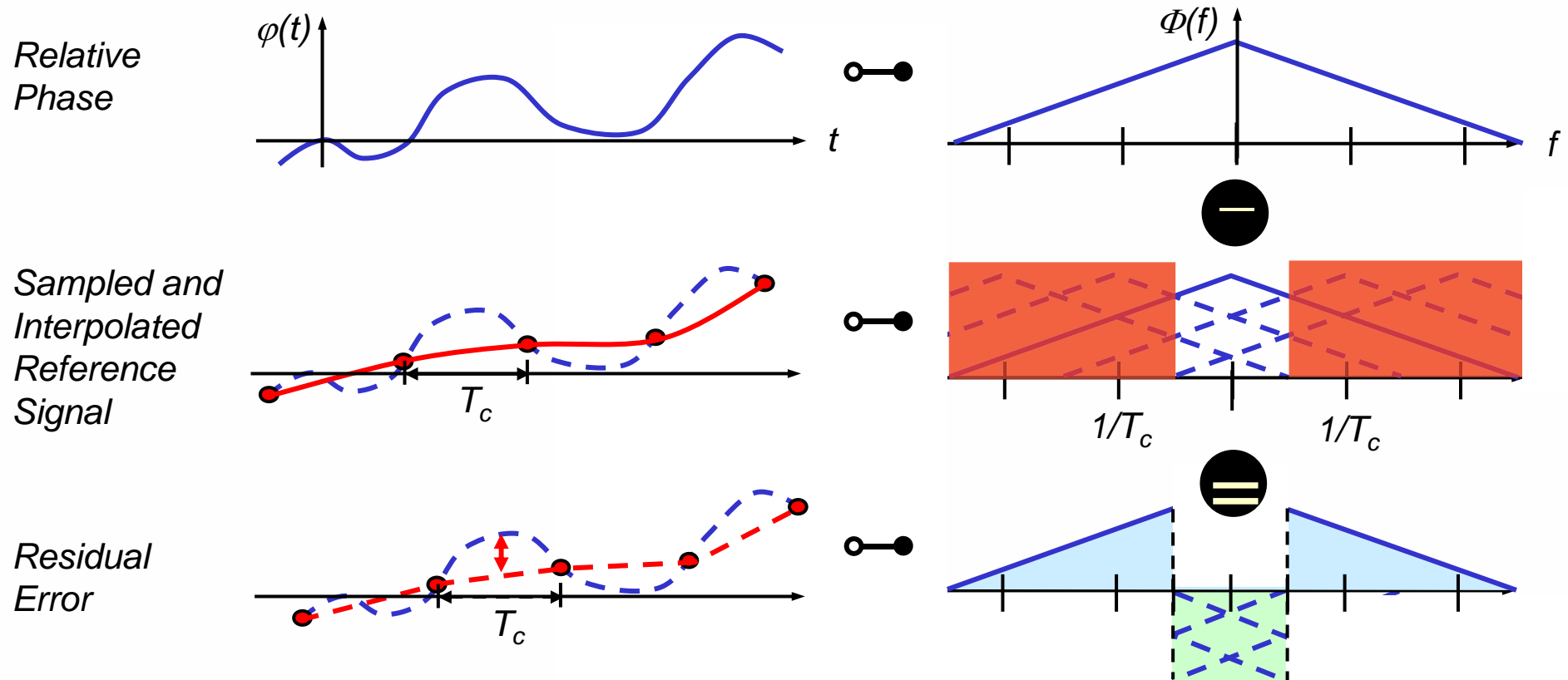


Sampled and  
Interpolated  
Reference  
Signal

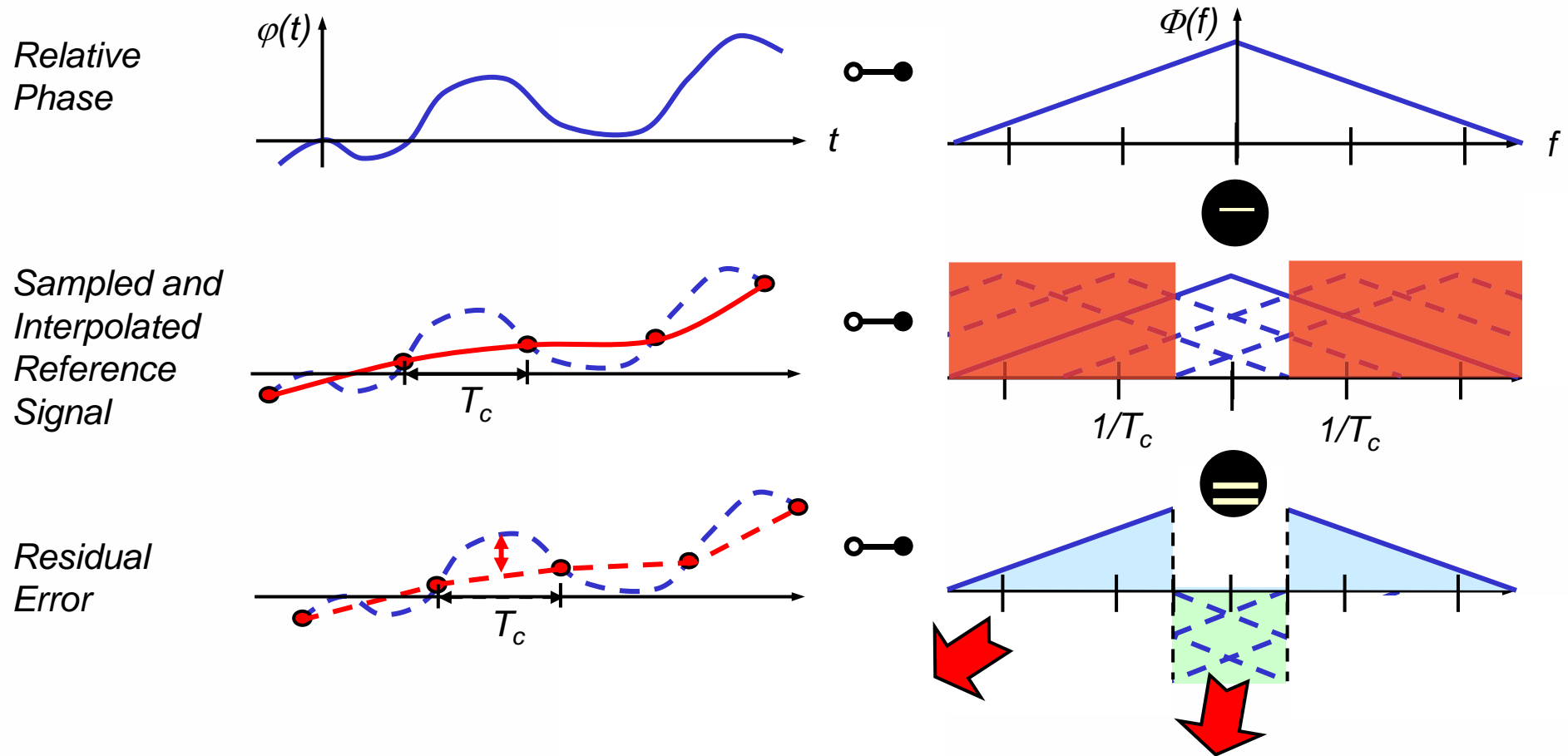




# Derivation of Required Sync Frequency



# Derivation of Required Sync Frequency



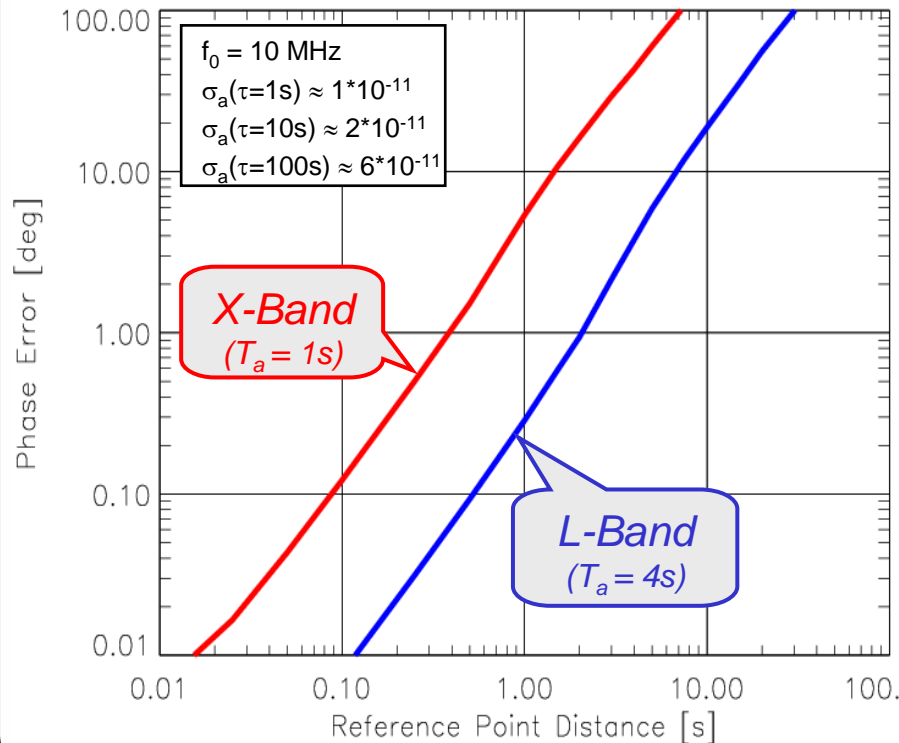
$$\sigma_{\phi}^2 = 2 \cdot \left( \frac{f_0}{f_{\text{osc}}} \right)^2 \cdot \left[ \int_{1/(2T_c)}^{\infty} S_{\phi}(f) \cdot \left( \frac{\sin(\pi T_a f)}{\pi T_a f} \right)^2 \cdot df + \sum_{i=1}^{\infty} \int_{-1/(2T_c)}^{1/(2T_c)} S_{\phi}\left(f + \frac{i}{T_c}\right) \cdot \left( \frac{\sin(\pi T_a f)}{\pi T_a f} \right)^2 \cdot df \right]$$



# Predicted Phase Error

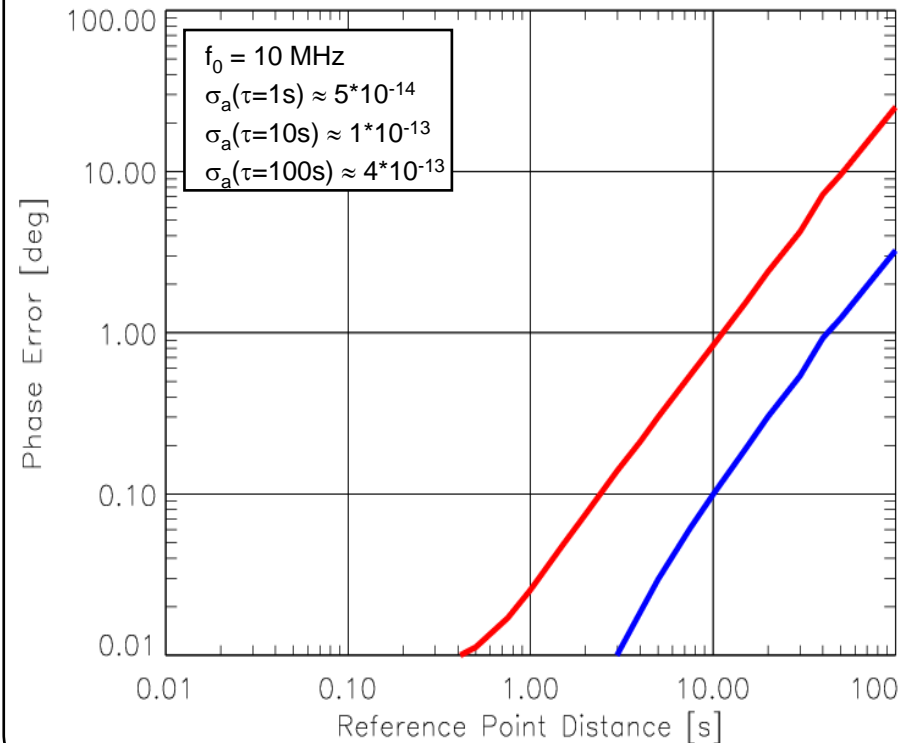
## „Normal USO“

Interferometric Phase Error



## „Hyper-Stable“ Oscillator

Interferometric Phase Error



$$\sigma_{\varphi}^2 = 2 \cdot \left( \frac{f_0}{f_{\text{osc}}} \right)^2 \cdot \left[ \int_{1/(2T_c)}^{\infty} S_{\varphi}(f) \cdot \left( \frac{\sin(\pi T_a f)}{\pi T_a f} \right)^2 \cdot df + \sum_{i=1}^{\infty} \int_{-1/(2T_c)}^{1/(2T_c)} S_{\varphi}\left(f + \frac{i}{T_c}\right) \cdot \left( \frac{\sin(\pi T_a f)}{\pi T_a f} \right)^2 \cdot df \right]$$



Deutsches Zentrum  
für Luft- und Raumfahrt e.V.  
in der Helmholtz-Gemeinschaft

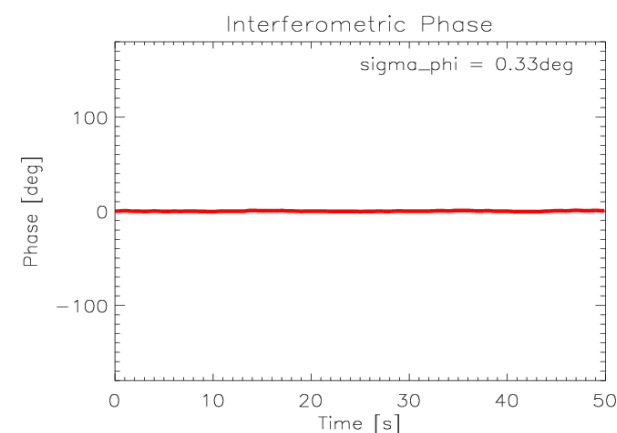
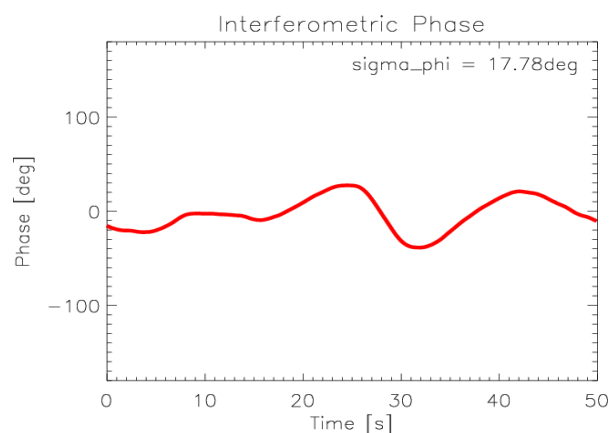
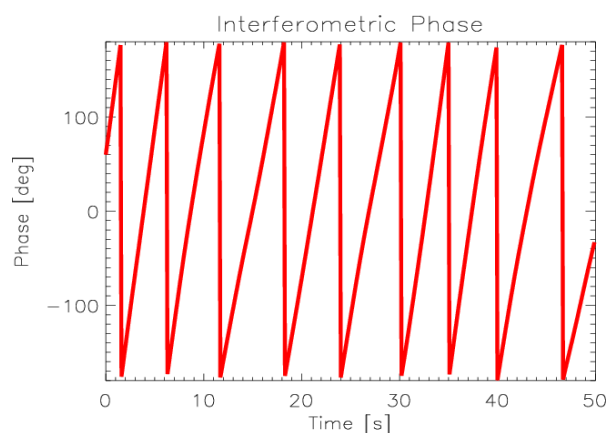
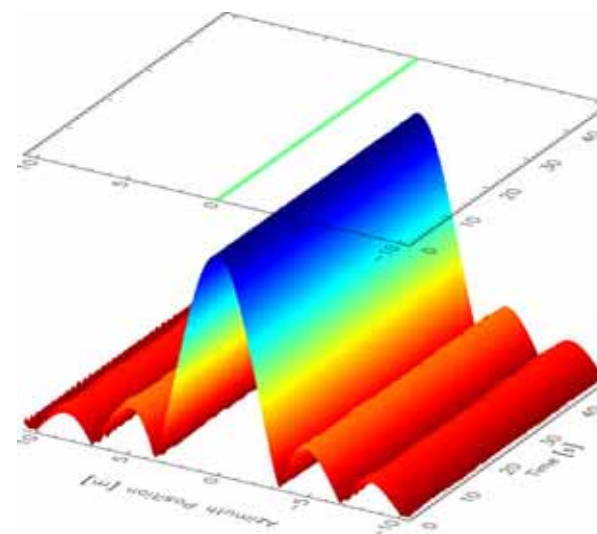
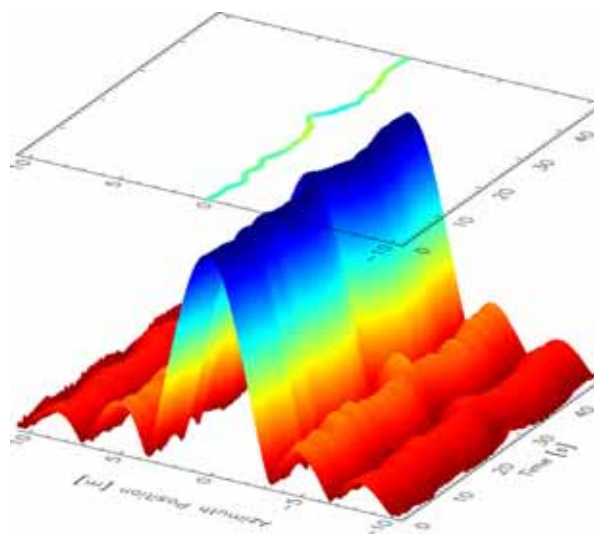
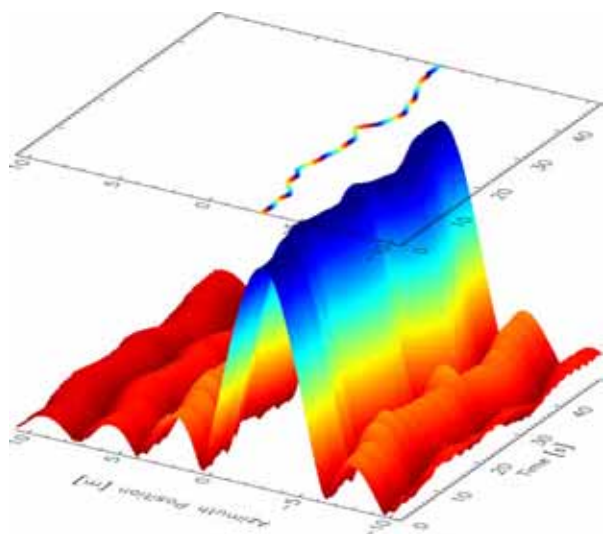


# Predicted Phase Error

No Phase Referencing

$T_C = 10$  sec

$T_C = 1$  sec

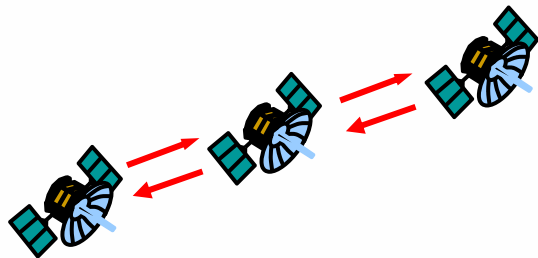




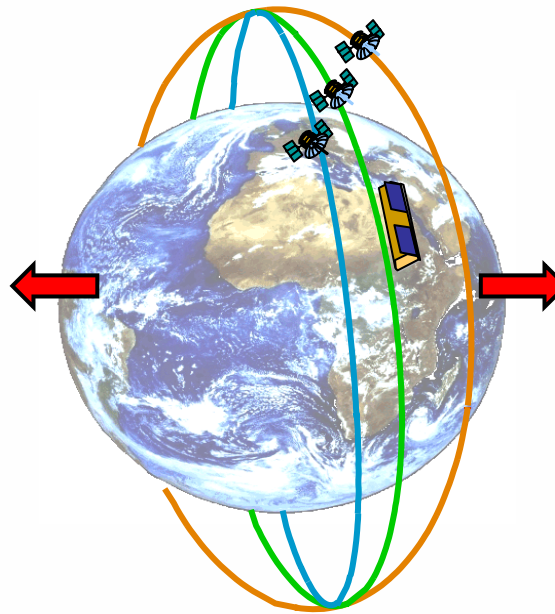
# Close Formation Flight: Collision Avoidance

## Autonomous Control

- *onboard computation of relative orbit maneuvers to ensure safe distance*
- *requires real-time relative position estimation and communication link between the satellites*

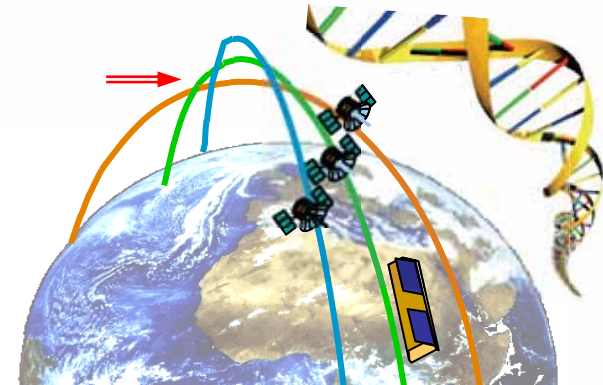


- *has to take into account fuel constraints (DART !)*



## Safe Orbit Design

- *avoid orbit crossings, e.g. by slight eccentricity vector offset (HELIX)*

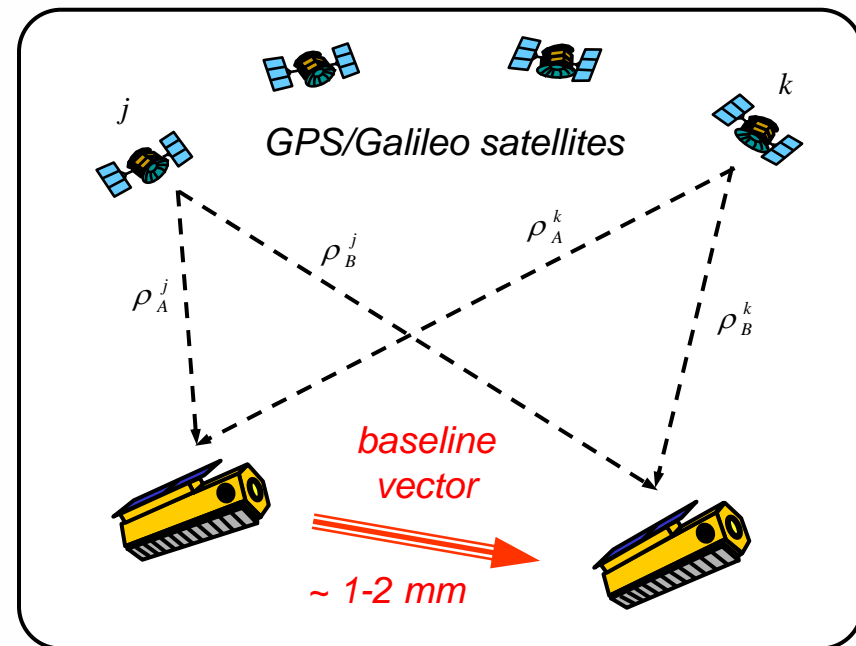
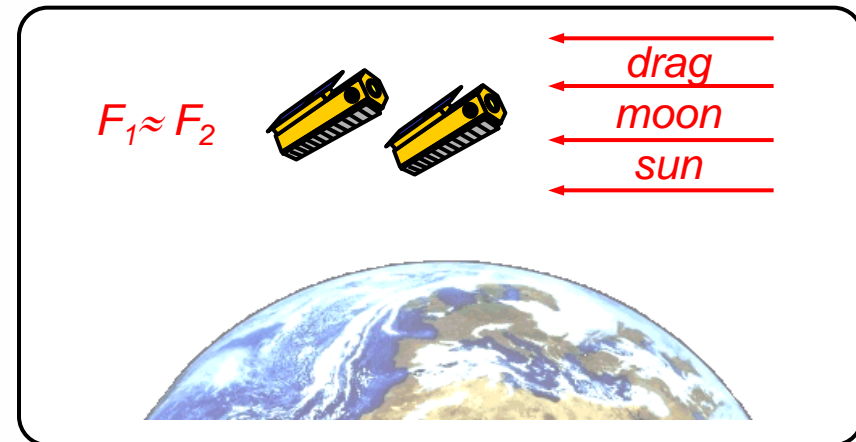


- *arbitrary satellite shifts along their orbits*
- *very short along-track baselines for a given latitude range*



## Baseline Control and Baseline Estimation

- **All satellites are exposed to almost identical orbit perturbations**
  - precise baseline control with low fuel consumption (e.g. to ensure fixed baseline ratios)
  - very small differential acceleration ( $< 100 \text{ nm/s}^2 \rightarrow$  baseline error of less than 1 mm for 1000 km swath)
- **Precise baseline estimation by**
  - double-difference GPS/Galileo carrier-phase measurements
  - accurate orbit propagation model
  - several studies predict a 3-D accuracy in the order of 1-2 mm





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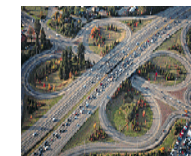
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